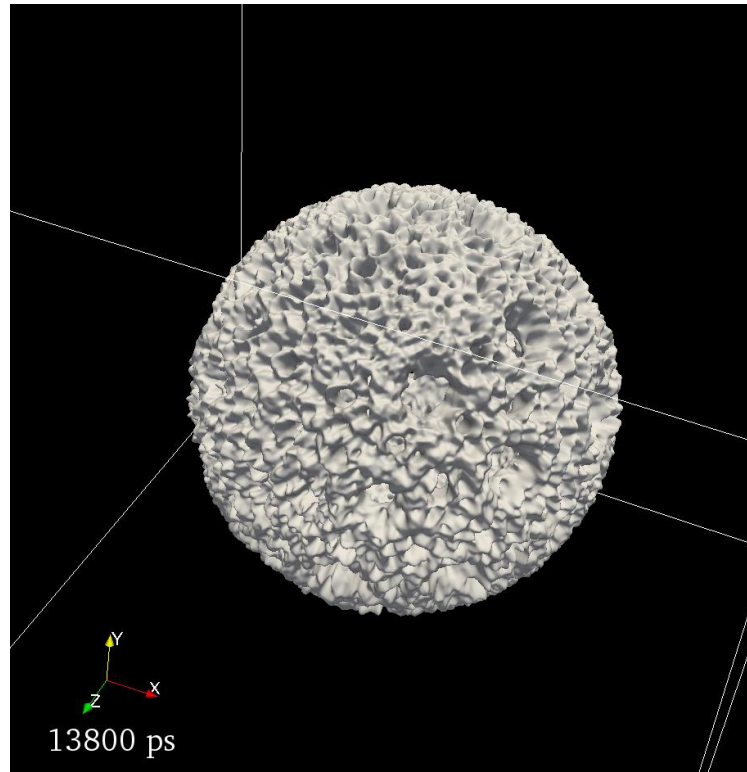


## High Energy Density Physics Modelling at Imperial College.

J.P. Chittenden, B.D. Appelbe, N.P.L. Niasse, J. Pecover, K. McGlinchey,  
C. Walsh, D. Botero-Garcia, J. Tong, A. Seaton, F. Manke

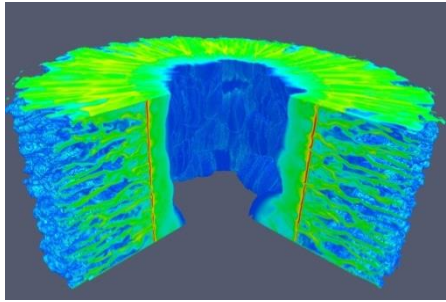
Centre for Inertial Fusion Studies, Imperial College, U.K.  
j.chittenden@imperial.ac.uk



# HEDP modelling suite



**Gorgon**

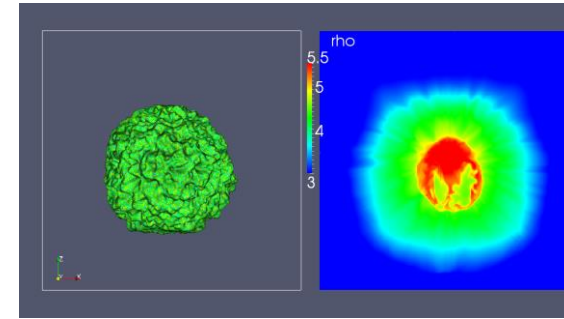


3D  $(x,y,z)$  or  $(r,z,\theta)$  Eulerian, resistive MHD code.  
Fully explicit version using vector potential representation.  
Parallel via domain decomposition.  
Detailed equation of state and transport data.  
non LTE DCA atomic & radiation loss model.

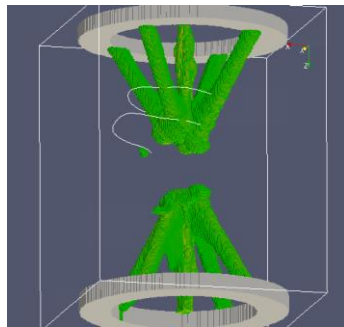


**Chimera**

3D  $(x,y,z)$ ,  $(r,z,\theta)$  or  $(r,\theta,\phi)$  Eulerian, rad-hydro code.  
Fully explicit version with parallel domain decomposition.  
Detailed equation of state and transport data.  
Multi-group diffusive radiation transport using non-LTE opacity data



**Melinda**

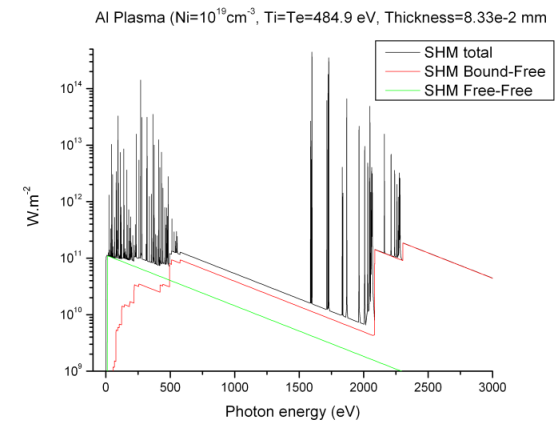


**All codes share common solvers and can be used in combination**

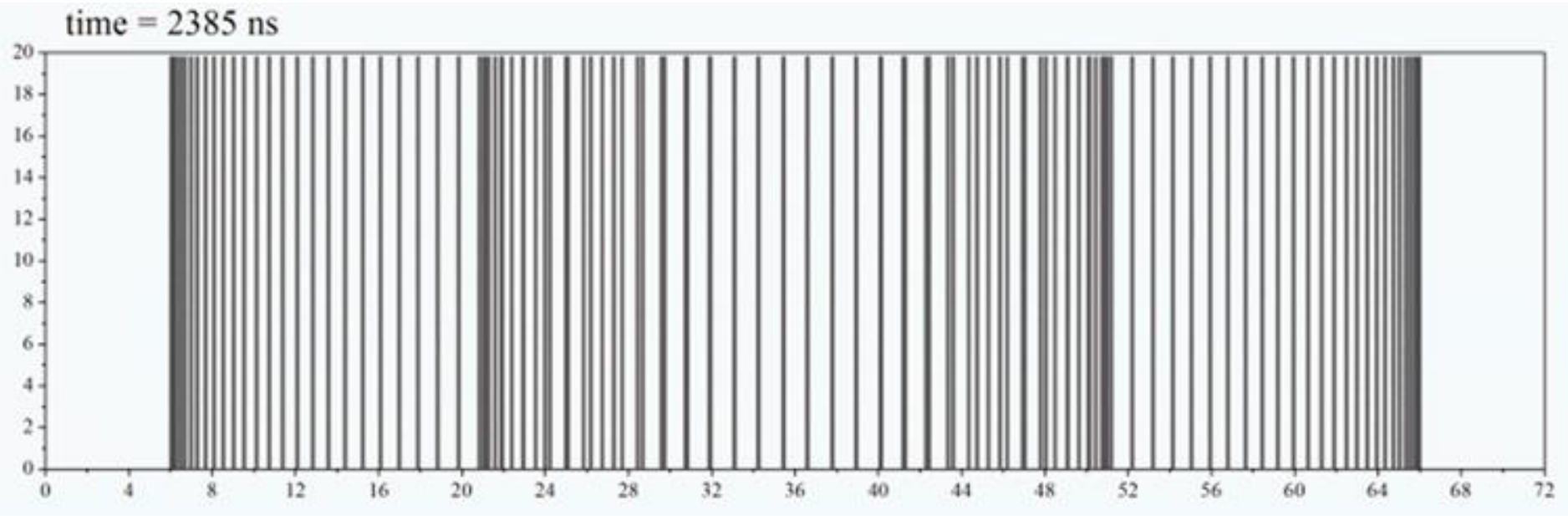


**SpK**

Multi-material CRE DCA emissivities and opacities based on a modified Saha model including n-l splitting with up to 10m excitation level  
Inline resolution of ionization populations, and radiative cooling rates, using tabulated data.  
Offline generation of detailed synthetic spectra and other diagnostics using detailed line profile radiation transport including Stark, lifetime, ion temperature and motion Doppler broadening.



Wire array Z-pinchs are intense (200TW) sources of soft x-rays  
used for inertial confinement fusion research



Z shot 1284 (60mm/30mm) nested Copper (80 wires / 40 wires) 10.97um

4608 processors,  $2880 \times 2880 \times 800 = 6,635,520,000$  computational cells,

24.7 Gigabytes per output, 2.7 Terabytes for this animation:

# K-shell radiation sources require detailed spectral modelling

**Spk** - a simple atomic model for large scale parallel HEDP simulations

Kernel:

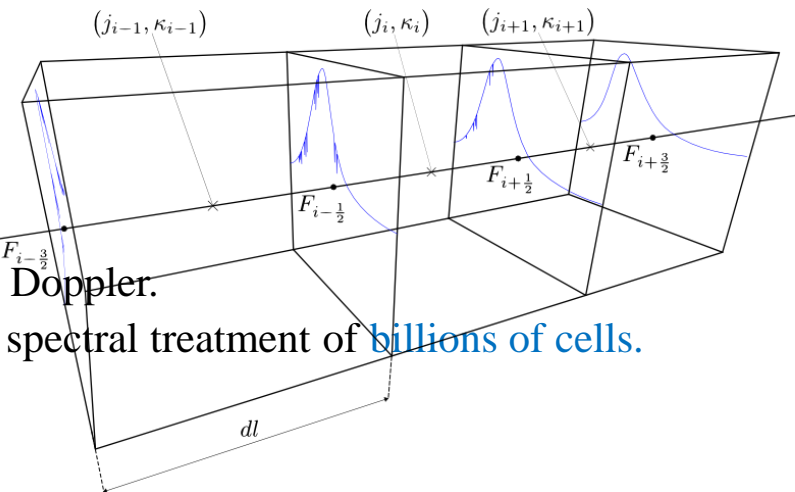
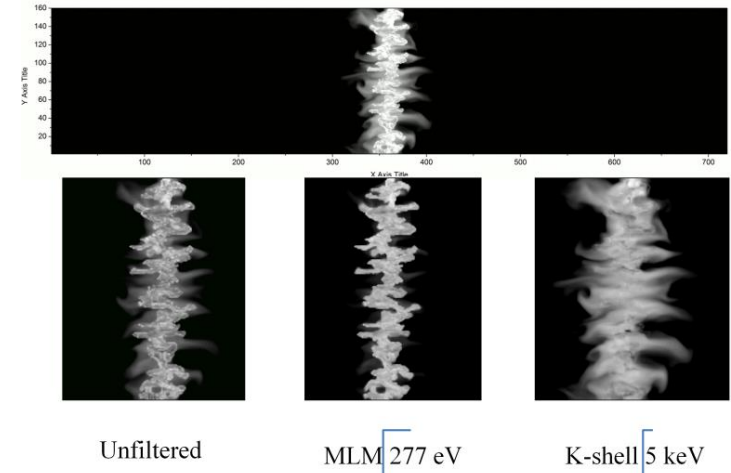
- **SHM + DCA** with nl splitting.
- **Experimental energy levels** from NIST atomic database.
- **Multi-material** for mixed opacities and emissivities.
- Self-consistent EOS (**very experimental for the moment**).

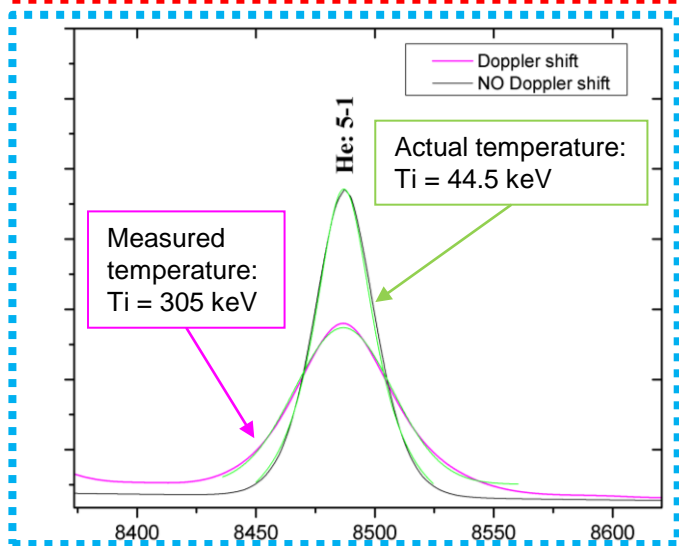
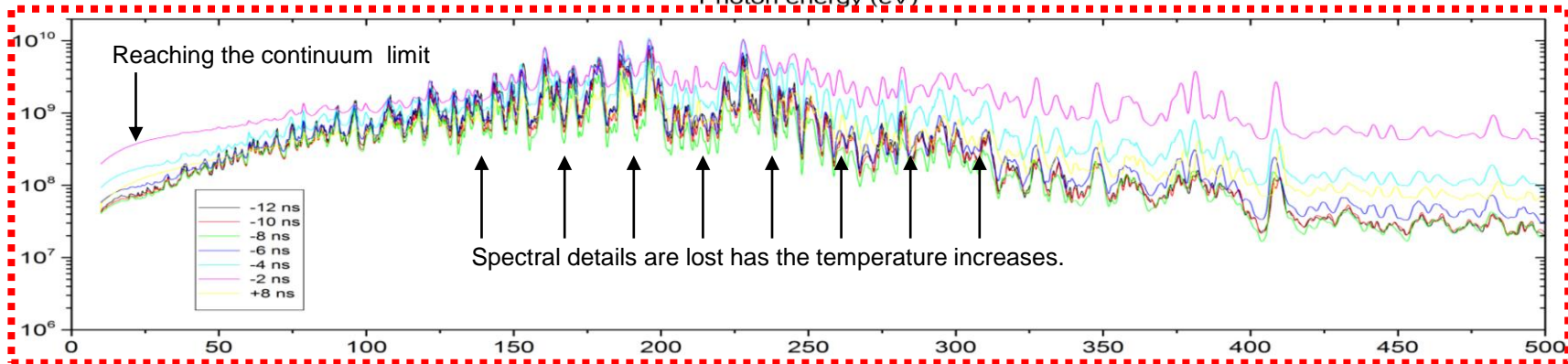
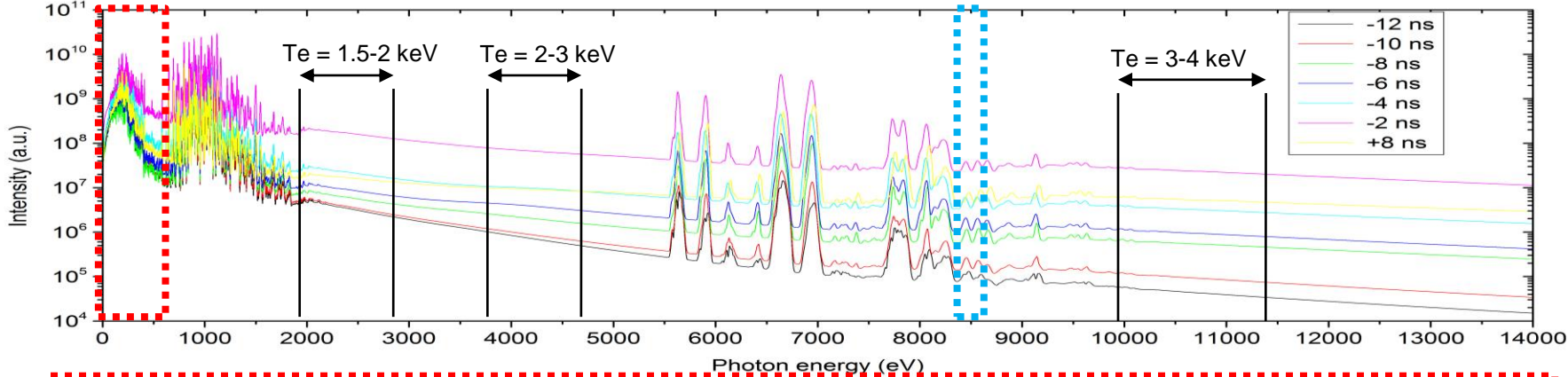
Inline:

- Highly **optimized**.
- Run in **parallel** with Gorgon 3D / Chimera.
- Radiation transport:  $P_{1/3}$  approximation.
- Fast pre-tabulation of **Voigt profiles**.
- **Filtered** synthetic bolometers and PCD signal.
- Self-consistent **non-LTE Te** and **zbar** .

Offline:

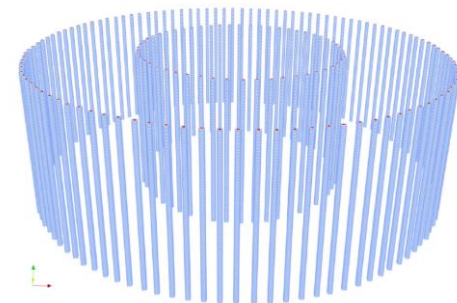
- **Detailed** spectra (FF, BF, BB [with  $\Delta n = 0$ ]).
- **High resolution** line profiles, continuum lowering **Stark & Doppler**.
- A modified **self-balancing binary search tree** to handle the spectral treatment of **billions of cells**.
- **Multi-material** opacity tables generation.
- **Filtered** synthetic MCP images.
- **Spatially integrated** spectra.



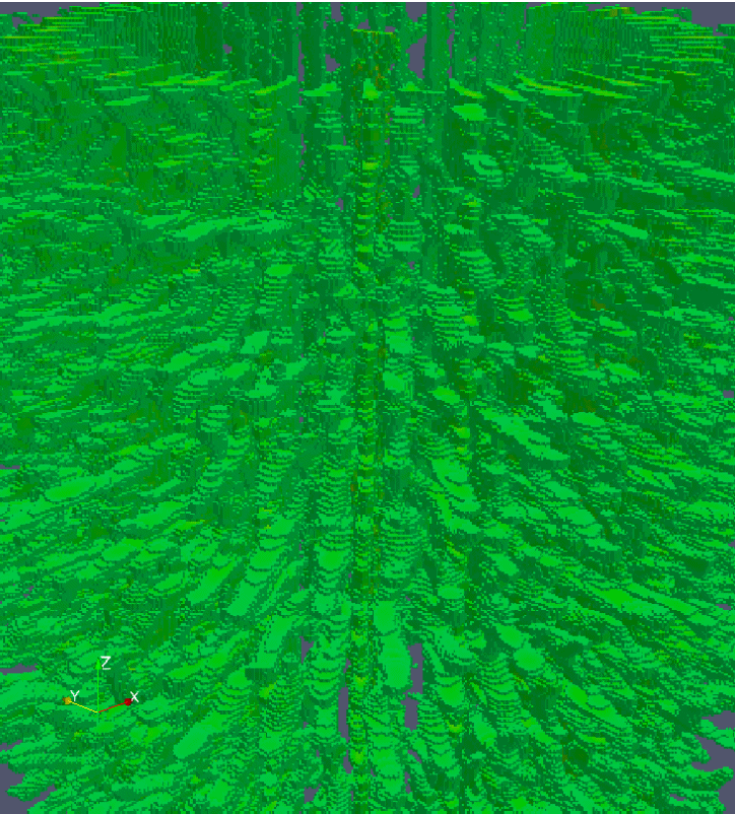


Spk spatially integrated time-dependent synthetic SS spectra, obtained by post processing the results of a ZR cylindrical wire array Z-pinch simulation performed by the Gorgon code.

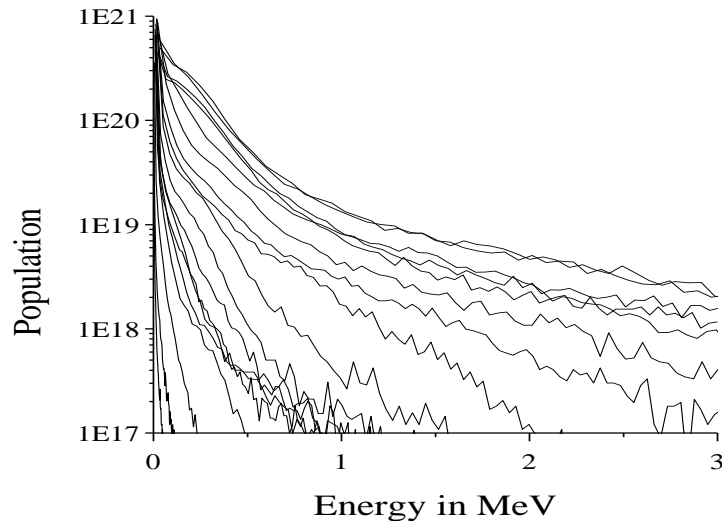
| <b>YIELD</b>     | <b>Total</b> | <b>L-Shell (&gt;1keV)</b> | <b>K-Shell (&gt;5KeV)</b> |
|------------------|--------------|---------------------------|---------------------------|
| Experimental [3] | 940 kJ       | 470 kJ                    | 85 kJ                     |
| SHM (old Spk)    | 994 kJ       | 370 kJ                    | 89-73kJ                   |
| $\Delta X/X$     | 5.4%         | 21%                       | 4-14%                     |



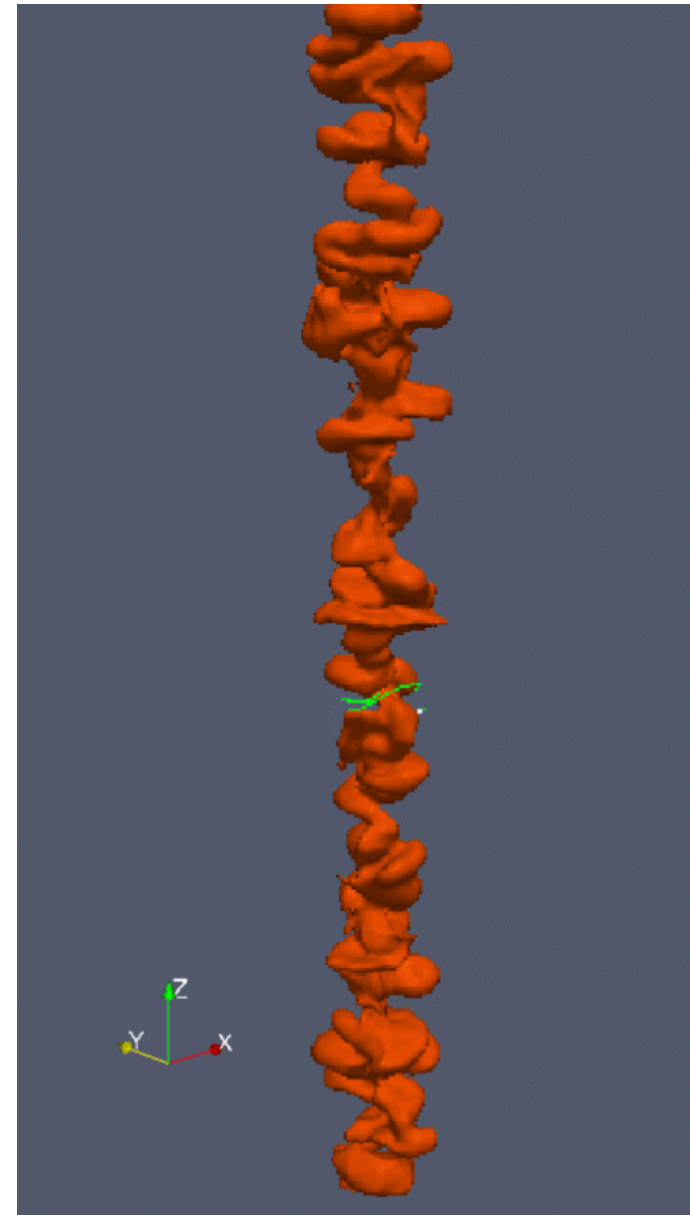
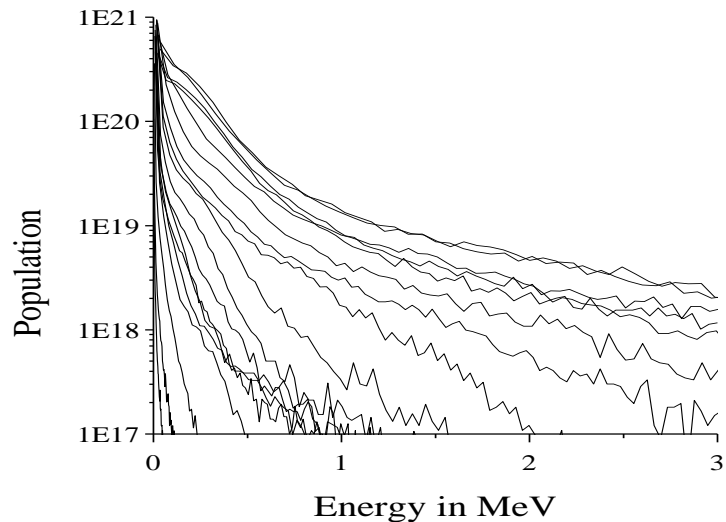
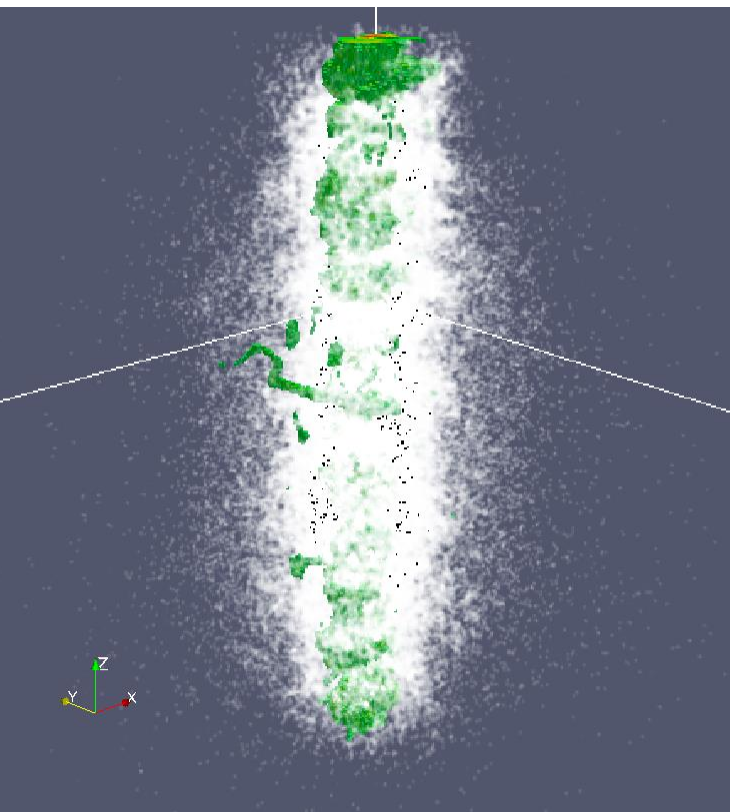
Electrons increase in number & energy and their trajectories become increasingly eccentric in the convoluted field of caused by instabilities



Results were generated using 3000 cores on the ARCHER UK National High Performance Computing Facility



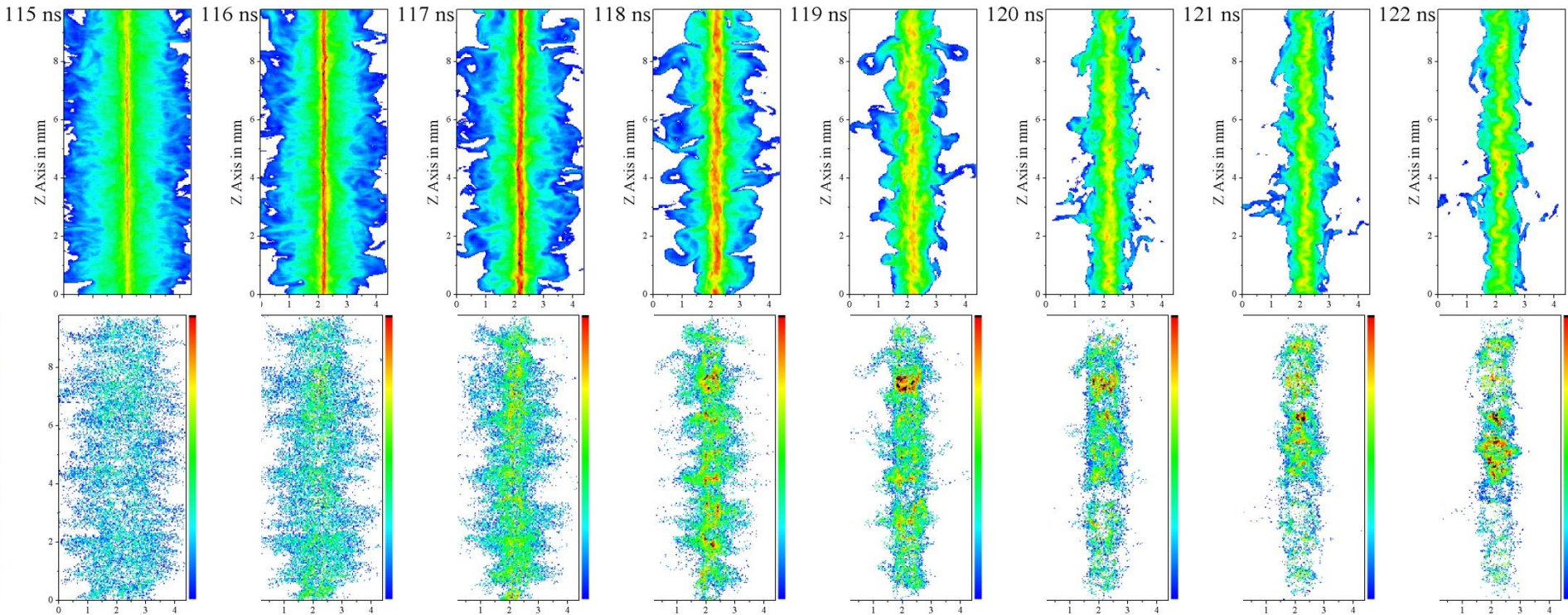
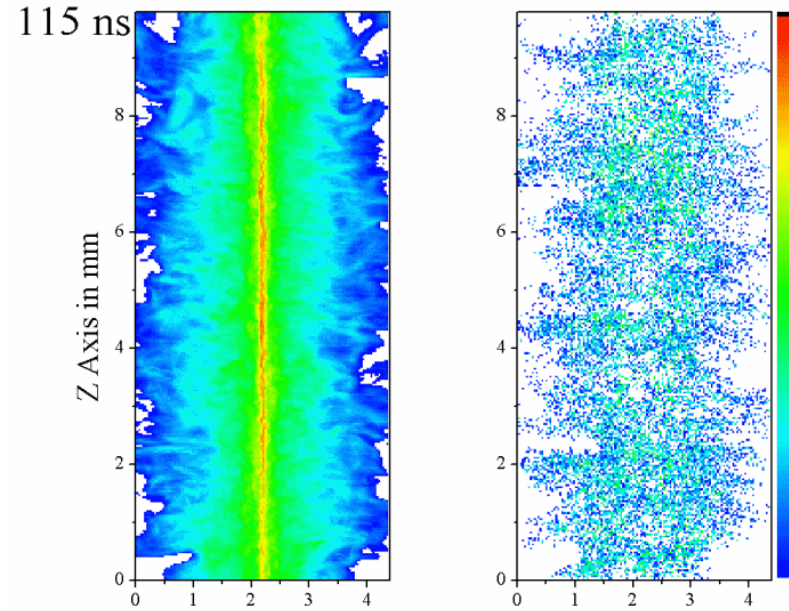
Electrons increase in number & energy and their trajectories become increasingly eccentric in the convoluted field of caused by instabilities



# K-alpha emission peaks after thermal soft X-ray and typically coincides with regions of strong MHD instability

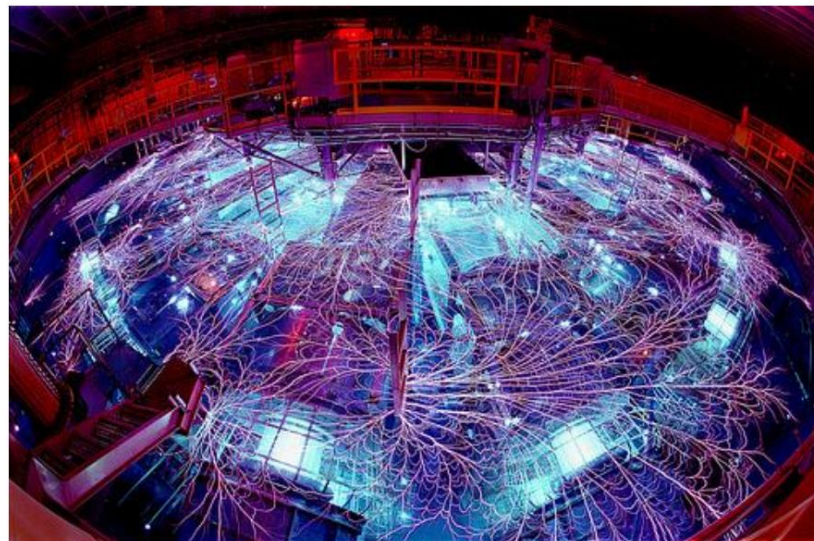
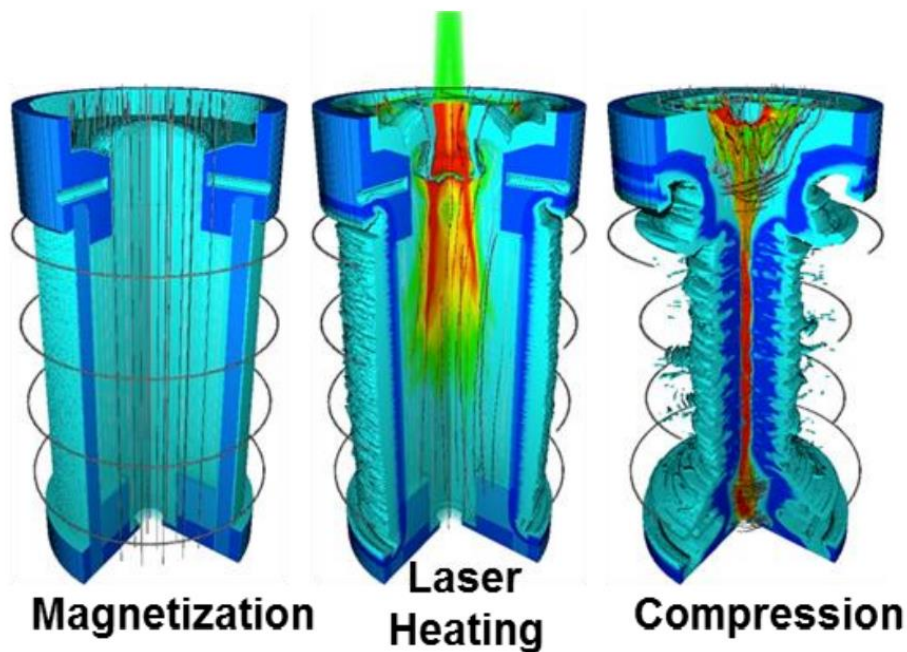
Here the product of fast electron density and background ion density provides a rough estimate of the probability of K-alpha emission – assuming a constant cross-section above a few times the ionization potential ...

$$P_{\text{k-alpha}} \sim n_{\text{fast}} v_{\text{fast}} n_i \sigma E_{\text{k-alpha}}$$

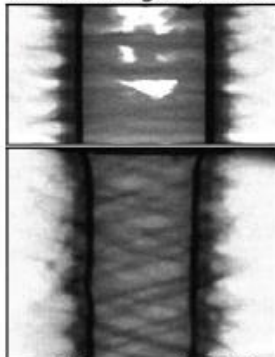




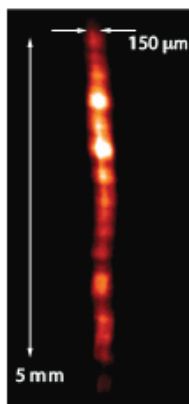
# Magnetised Liner Inertial Fusion (MagLIF) research at Sandia National Laboratory



Without Magnetic Field



With Magnetic Field



3123 eV  
imager  
data  
from  
z2584

Z2591

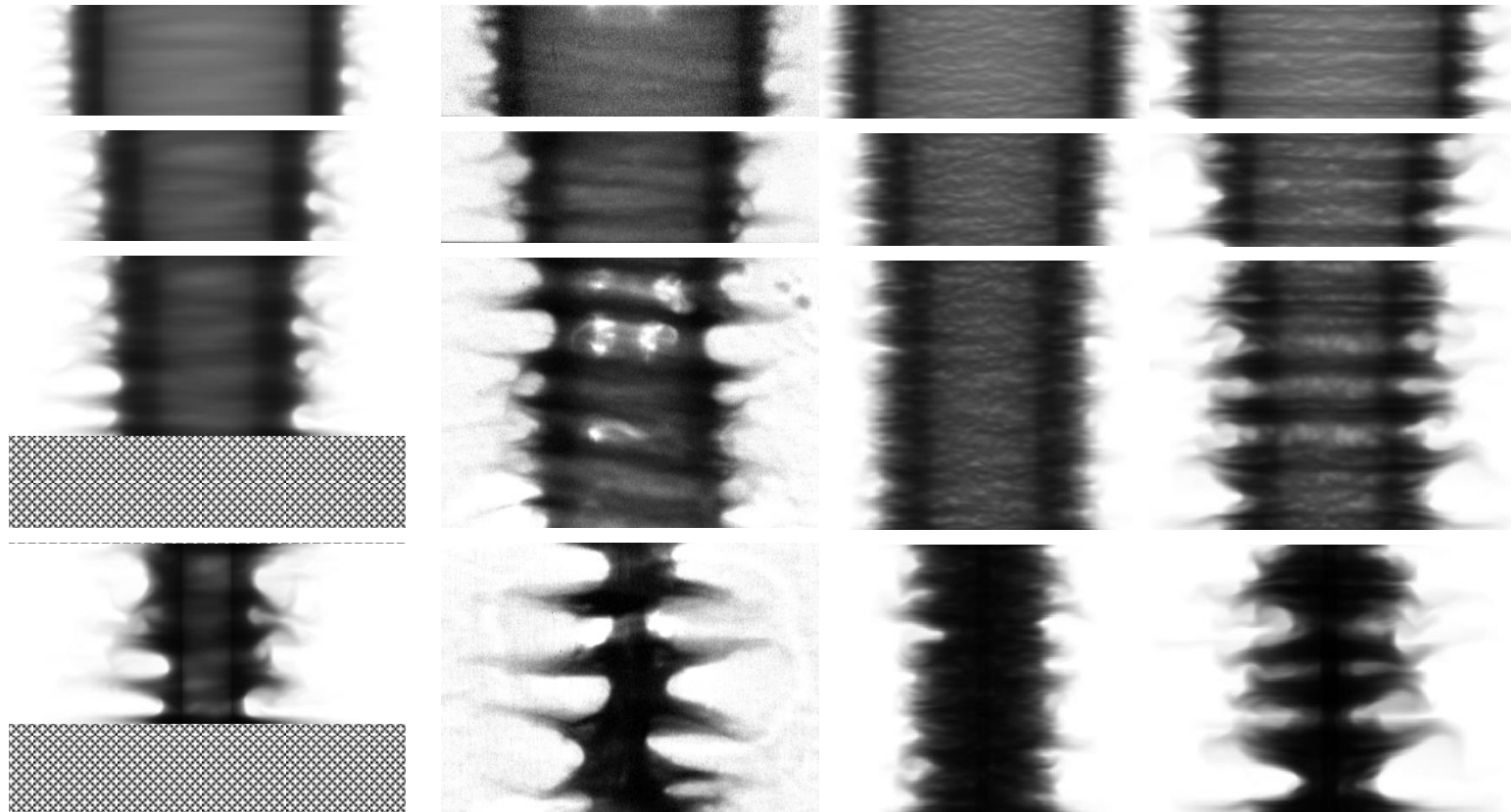
Beryllium liner, Deuterium filled

20MA current, 10 Tesla field, 2 kJ laser preheat

$2 \times 10^{12}$  neutron yield

# Material strength affects early phase instability amplitude and correlation in MagLIF liners

Synthetic radiography compared to experiment



3D Gorgon: random,  
4  $\mu\text{m}$   
inc. strength model

Experimental results  
from McBride *et al*

3D Gorgon:  
random, 20  $\mu\text{m}$

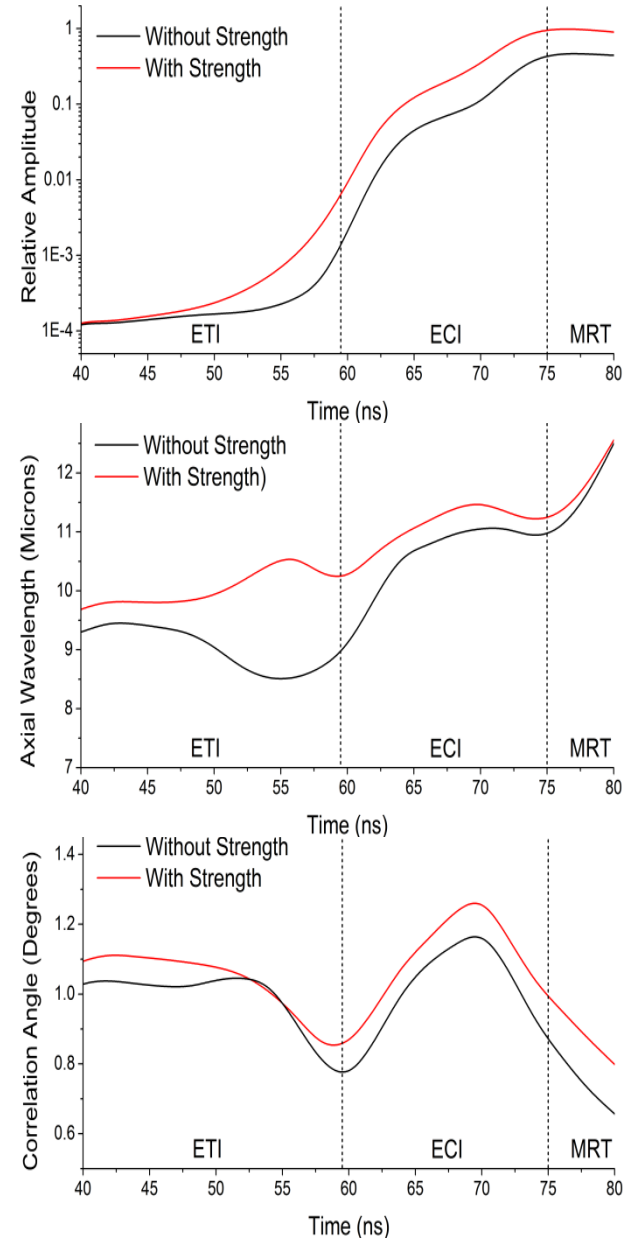
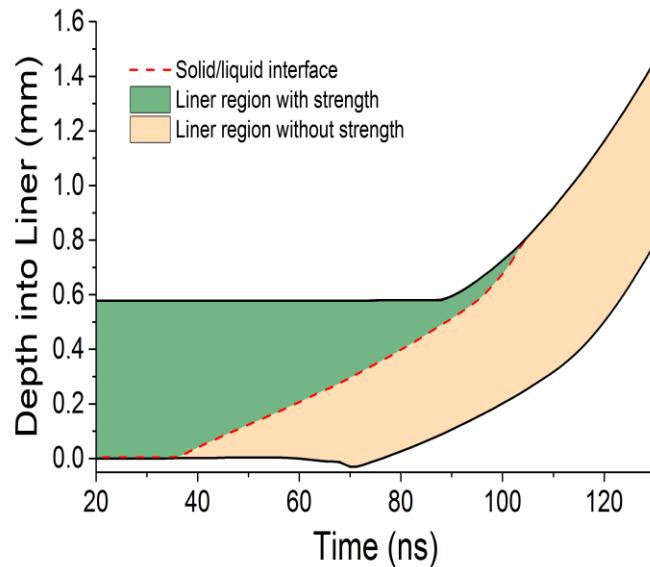
3D Gorgon:  
correlated, 20  $\mu\text{m}$

# Material strength affects early phase instability amplitude and correlation in MagLIF liners

Material strength added to Gorgon: liner loses strength by the time the ECI is seeded, has no effect in ECI or MRT phases.

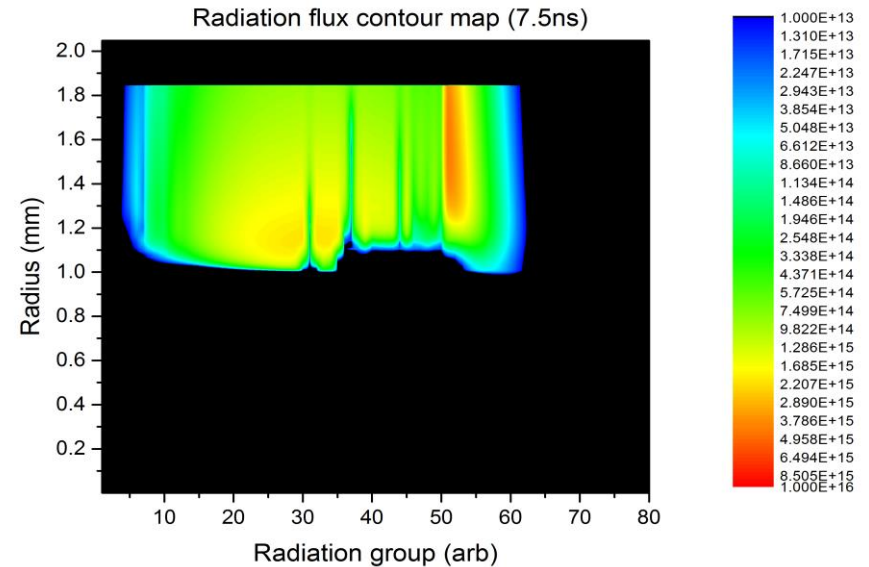
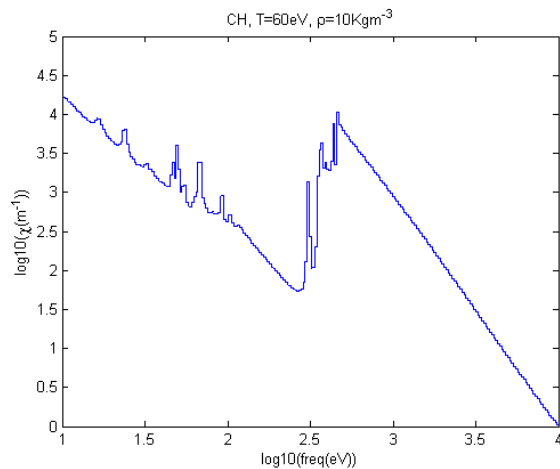
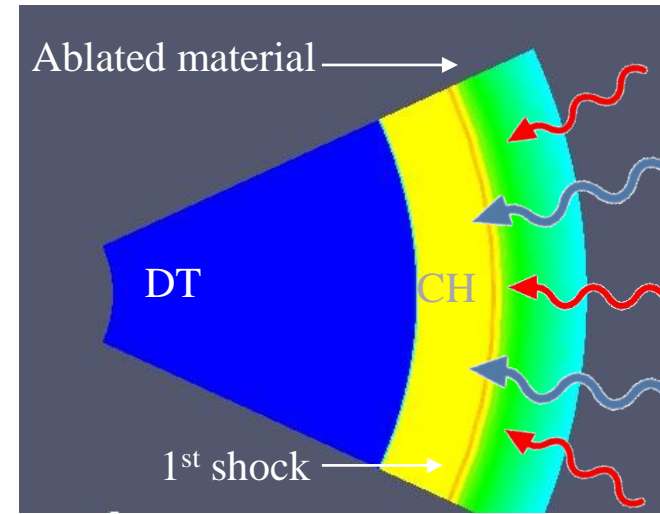
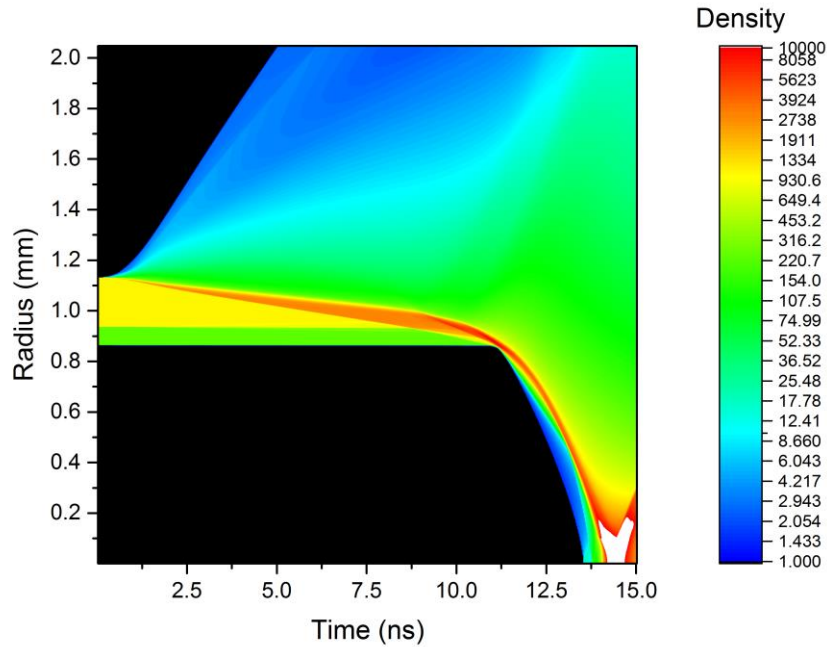
Strength results in longer ETI wavelength, higher amplitude and larger correlation angle

This difference remains through ECI and MRT phases



# X-ray ablation using a multi-group radiation diffusion approximation

see Niase et. al. Tu.Po.21

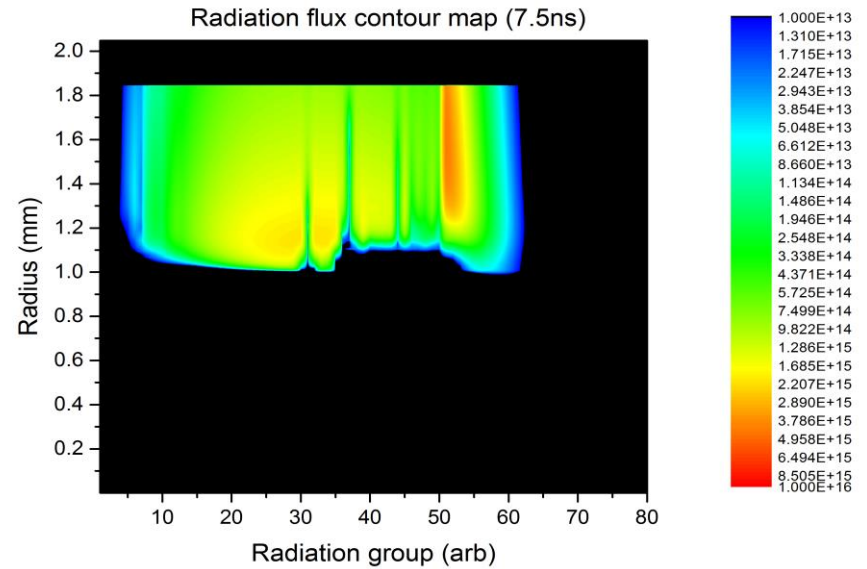
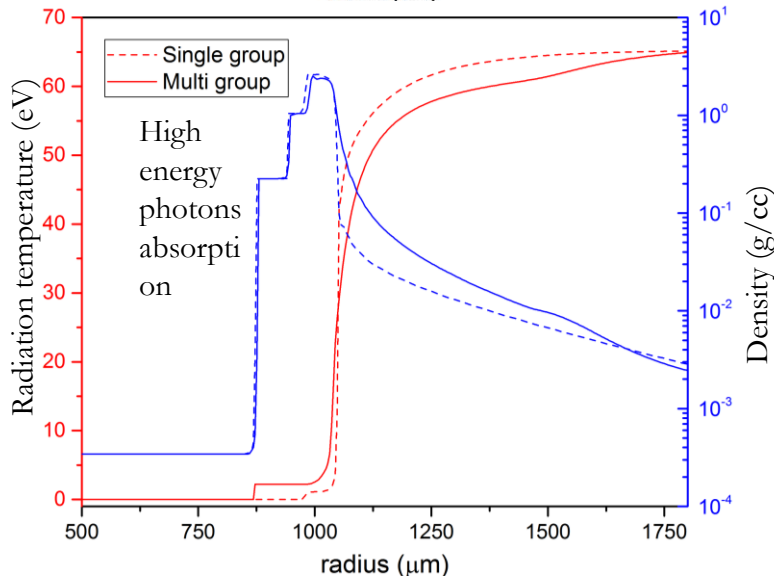
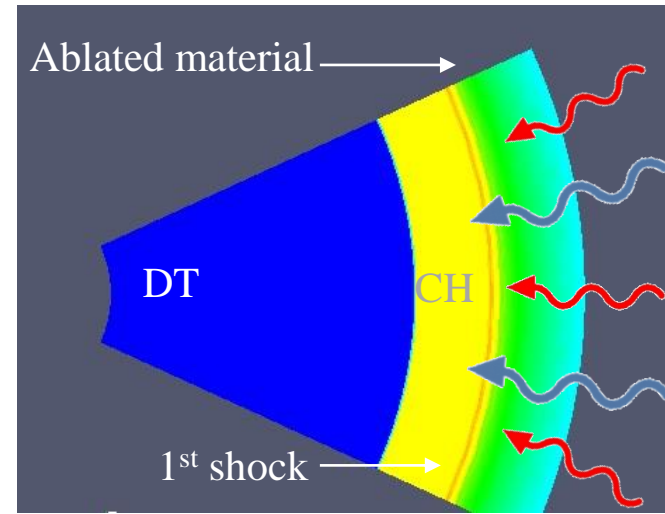
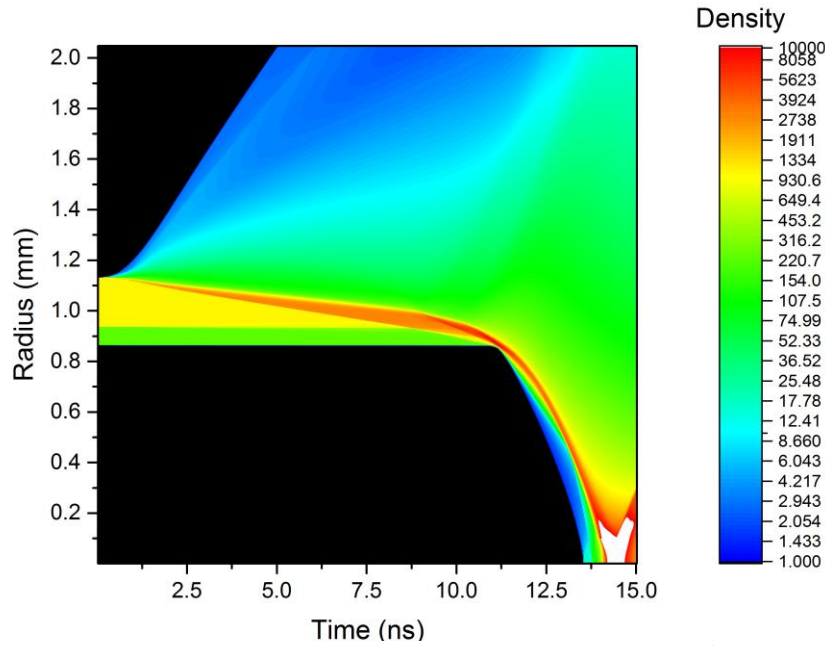


Tabulated CRE DCA opacities for CH from the SpK code are condensed into 80 radiation energy groups

Absorption of different photon energies at different densities smooths out the density gradient at the ablation surface

# X-ray ablation using a multi-group radiation diffusion approximation

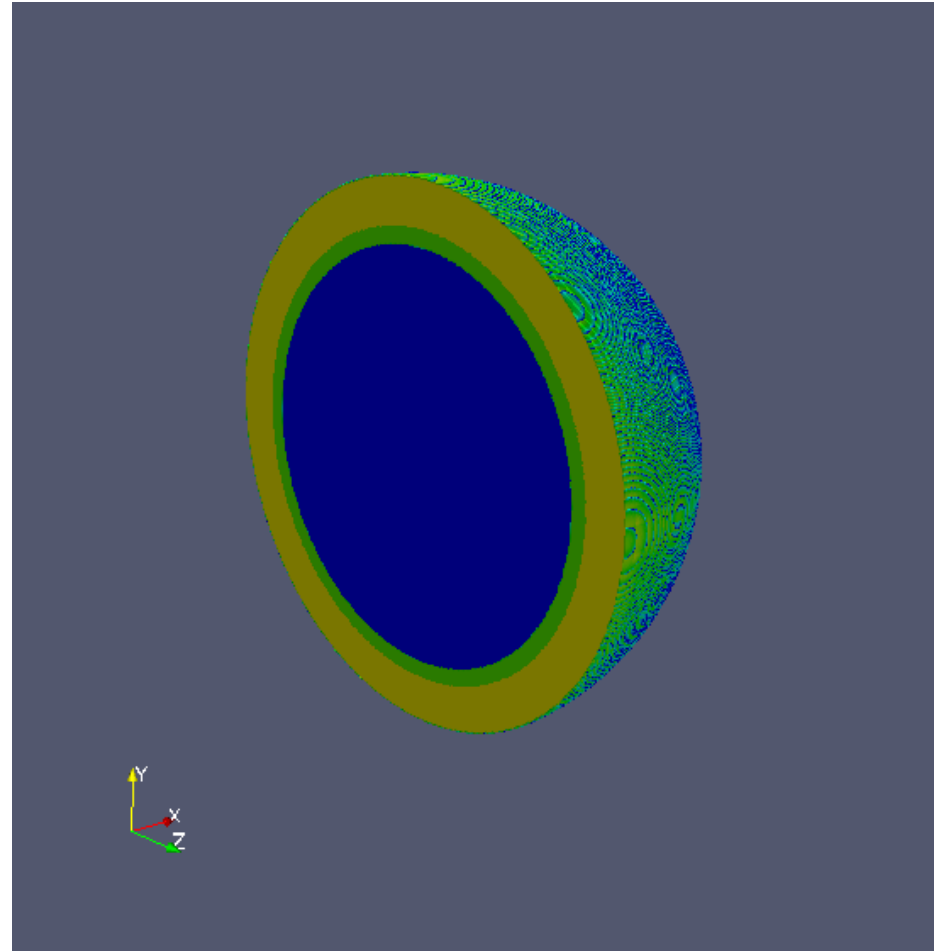
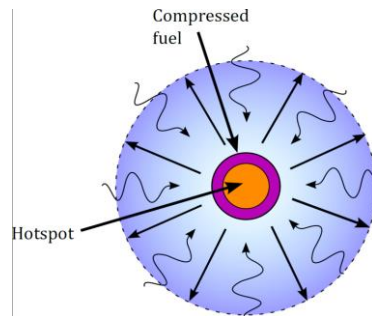
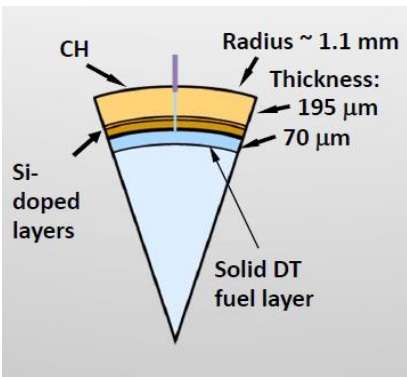
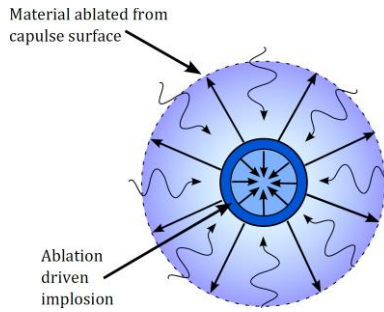
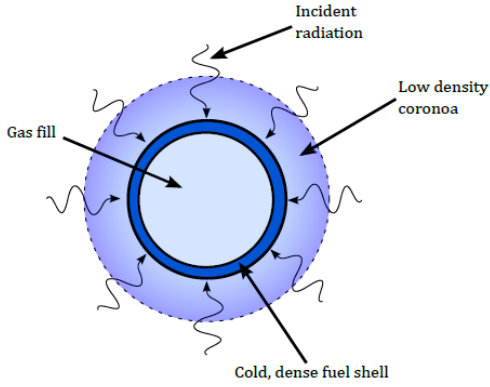
see Niase et. al. Tu.Po.21



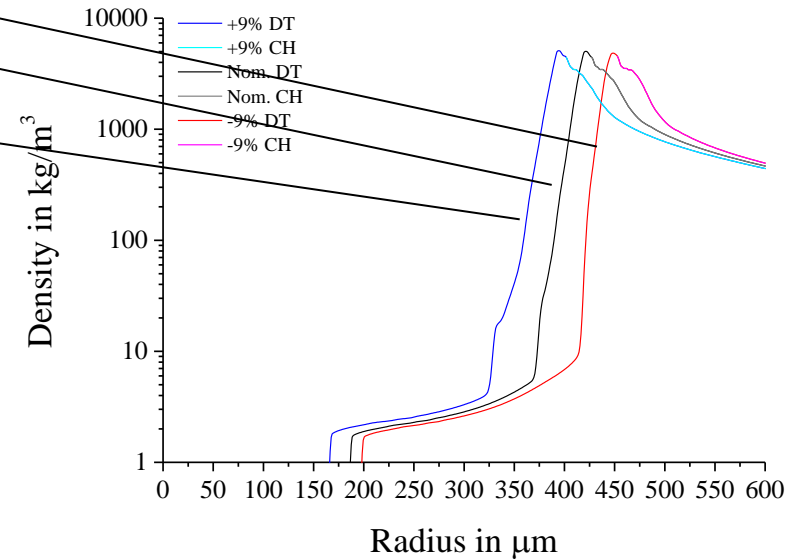
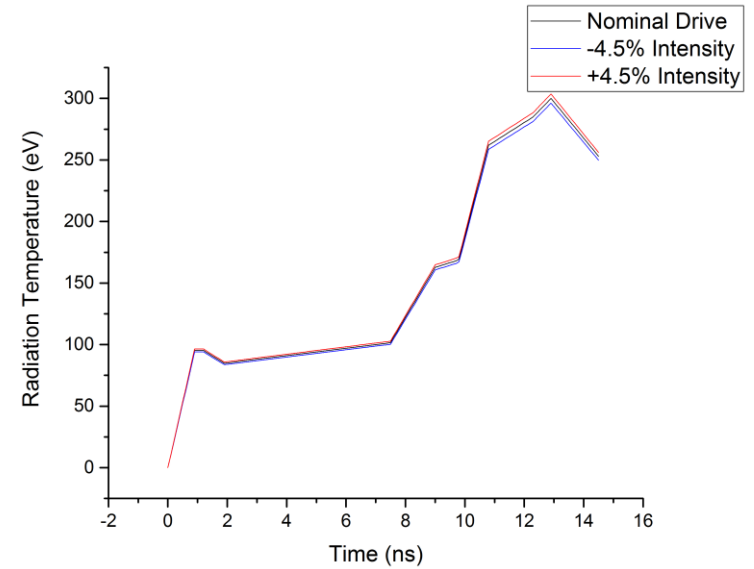
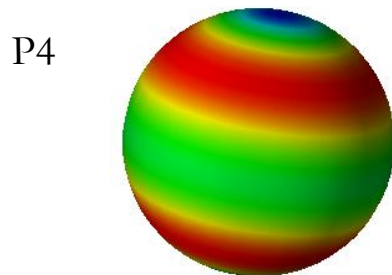
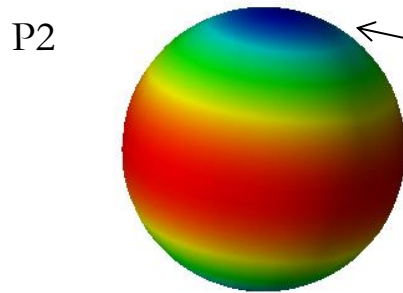
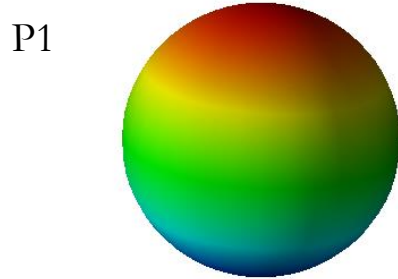
Absorption of different photon energies at different densities smooths out the density gradient at the ablation surface

Multi-group affects ablation scale length and adiabat(s)

# Full volume 3D radiation hydrodynamics simulations of NIF low-foot implosion



# Approximating low mode radiation drive asymmetry in 3D simulations of the deceleration and stagnation phases



For radiative drive asymmetry, 3D structure is constructed from a database of 1D results

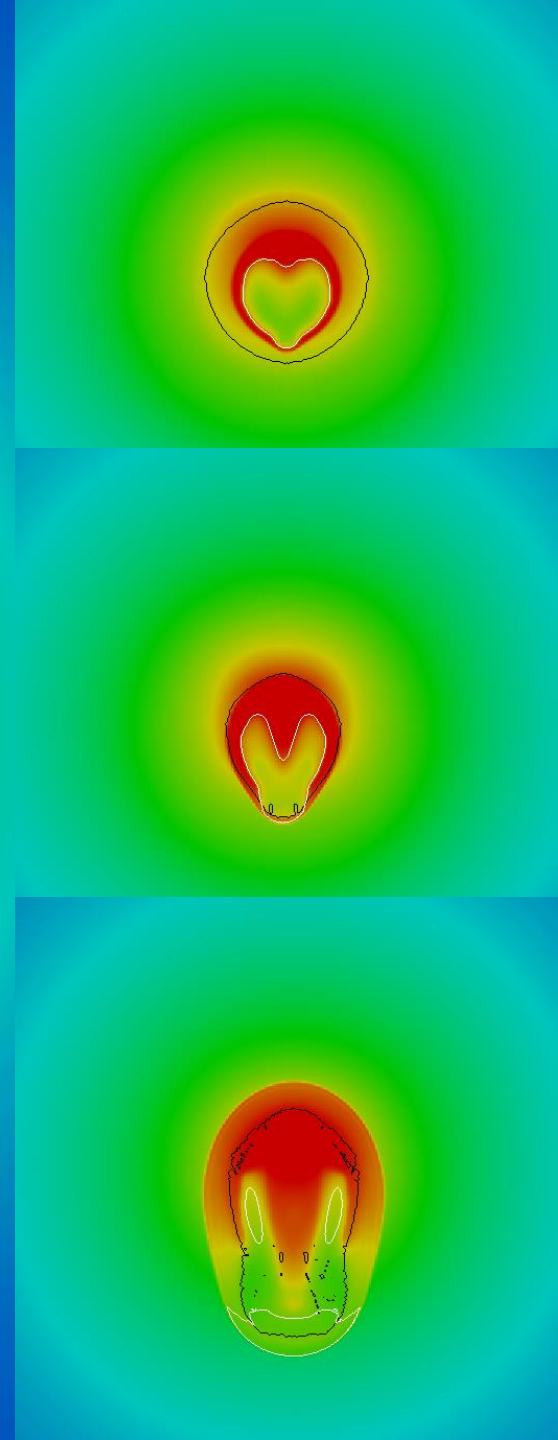
1D Chimera simulations with radiation drive multipliers at the start of the 'coast' phase

P1

Black contour – material interface  
White contour -1 keV Tion

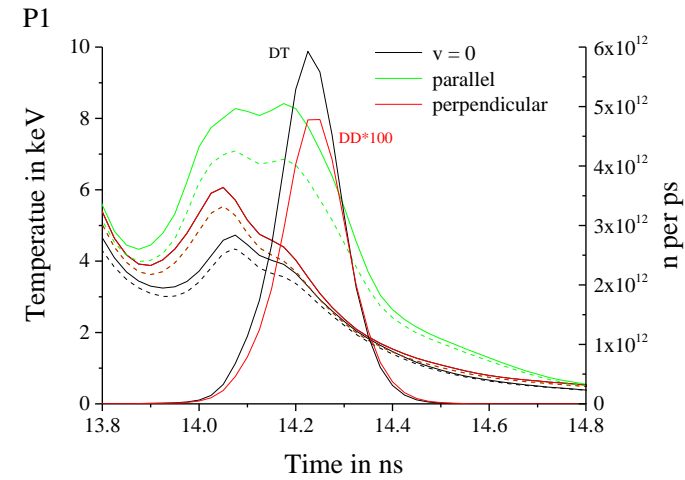
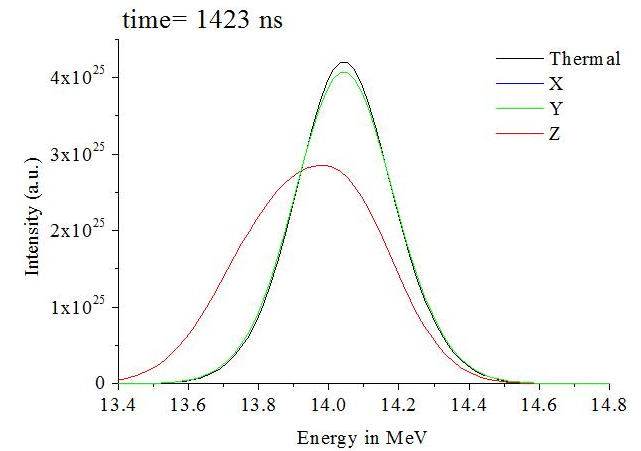
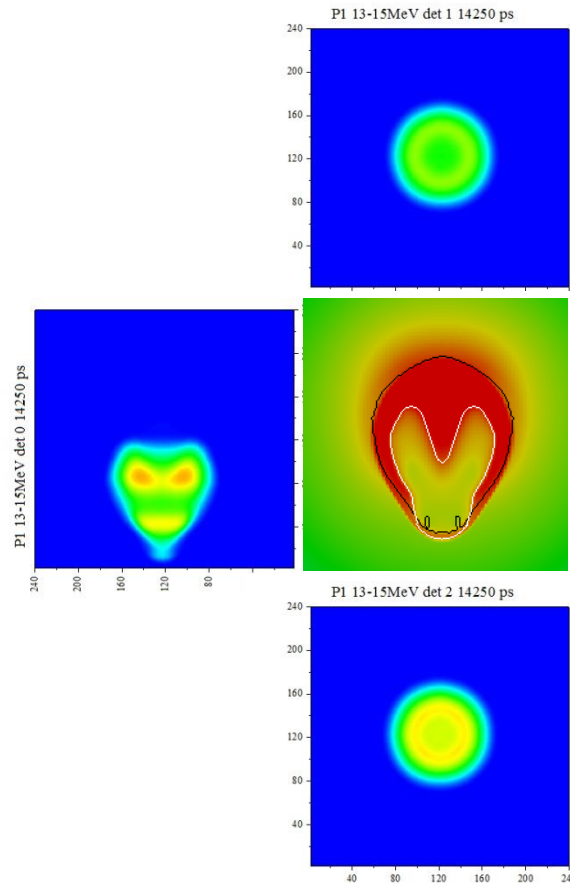


14250 ps





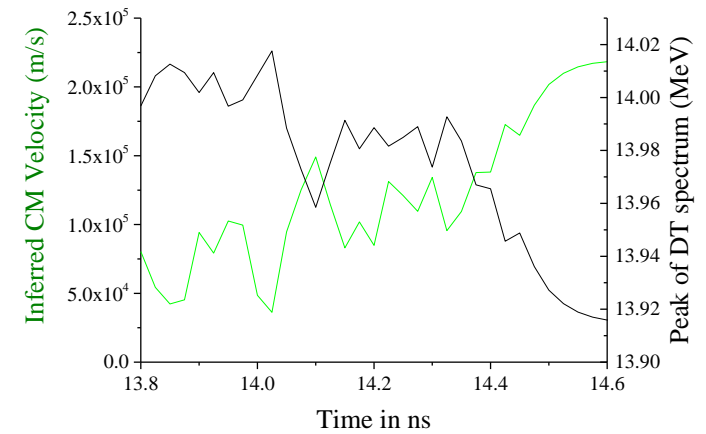
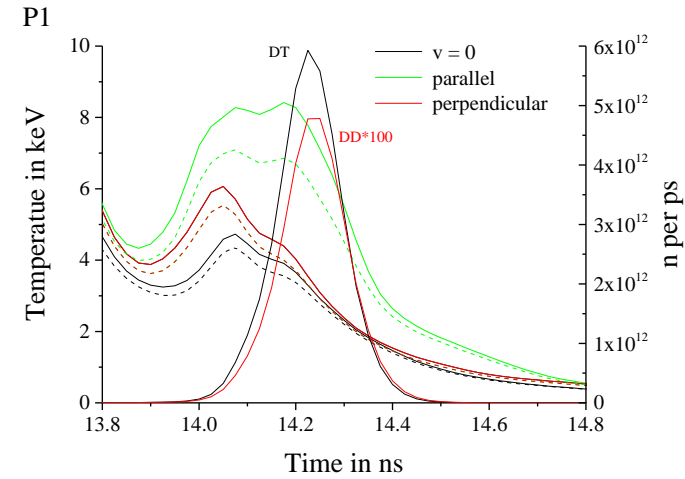
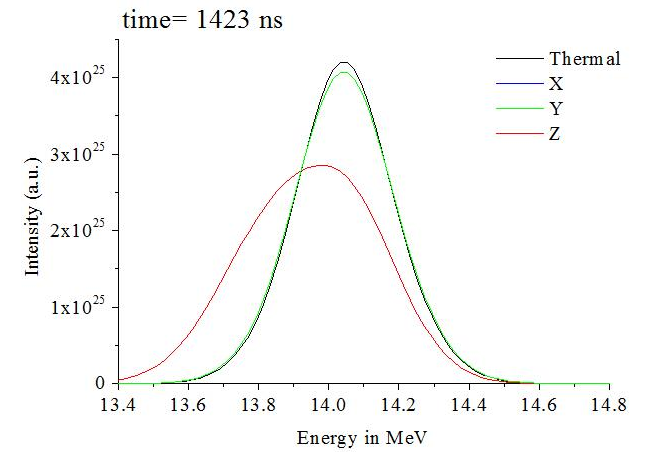
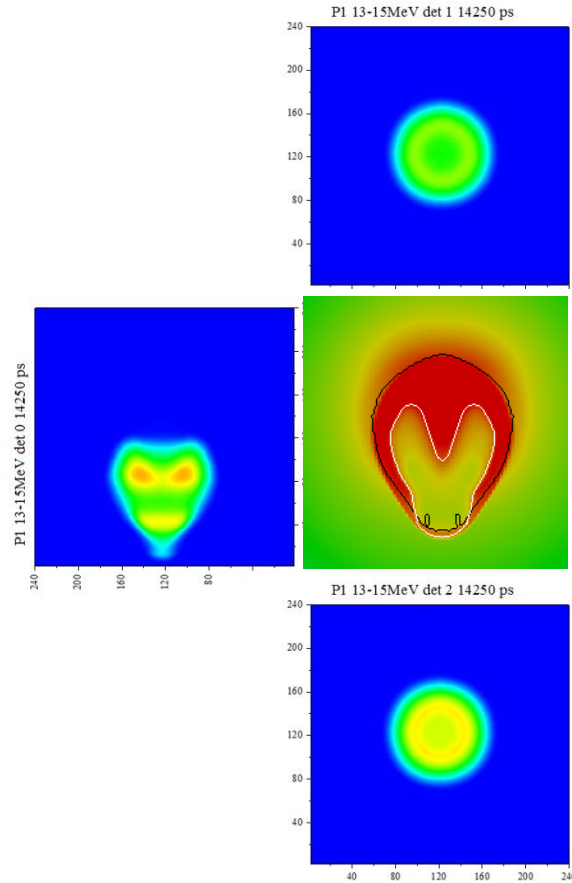
# P1 Temperatures inferred from primary neutron spectra



| Direction | T <sub>i</sub> (DT) |
|-----------|---------------------|
| v=0       | 2.973               |
| +x        | 3.191               |
| +z        | 6.731               |
| -z        | 6.971               |

T<sub>DT</sub> and T<sub>DD</sub> are anisotropic  
 T<sub>DT</sub> > T<sub>DD</sub> in parallel direction

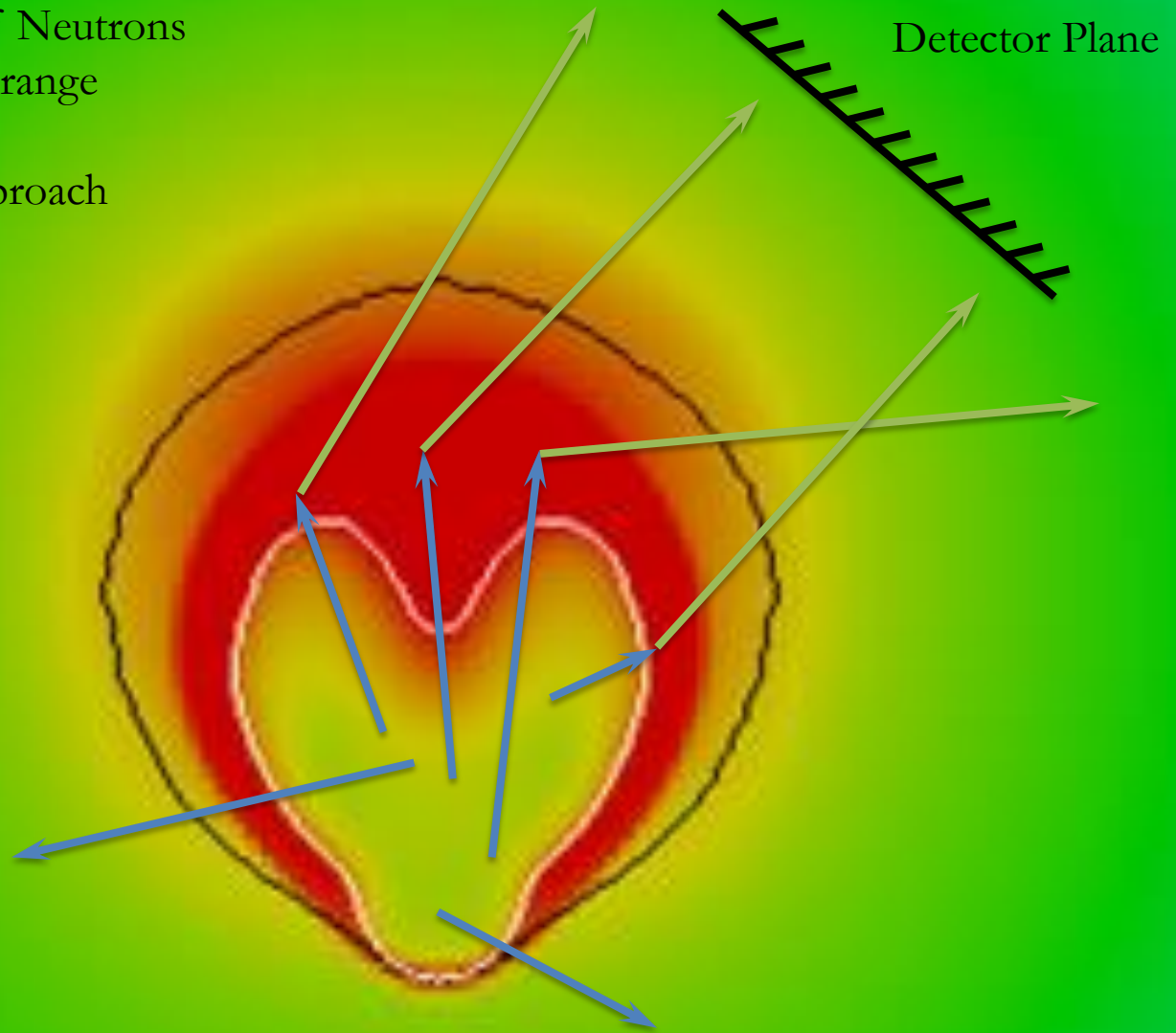
# P1 Bulk velocity inferred from primary neutron spectra



P1

Down Scatter of Neutrons  
into 10-12 MeV range

Monte Carlo approach

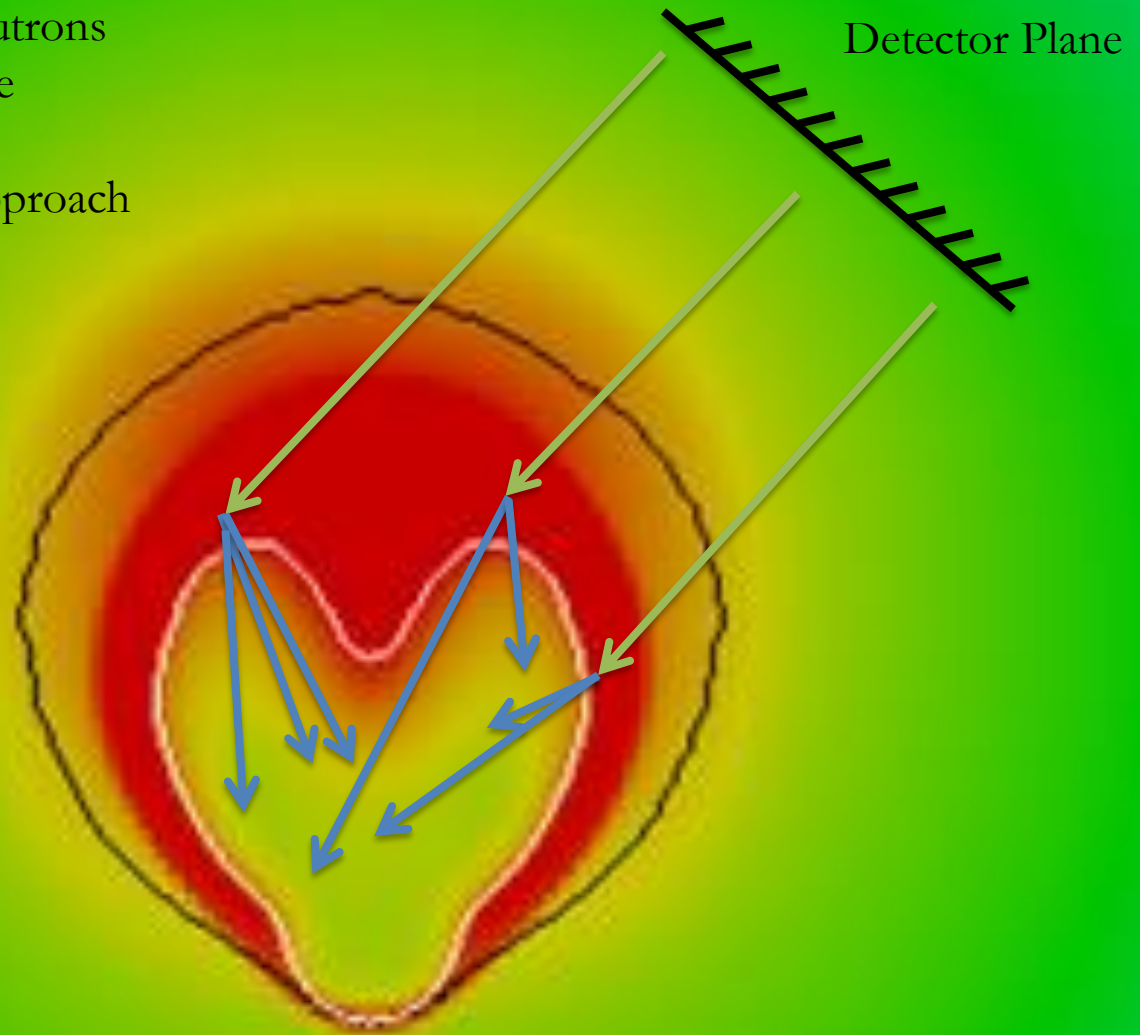


All neutron sources and scattering  
events are modelled – computationally  
expensive, numerically noisy

P1

Down Scatter of Neutrons  
into 10-12 MeV range

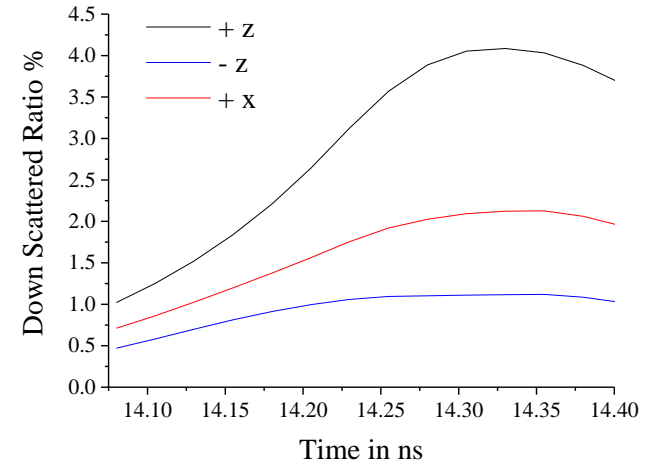
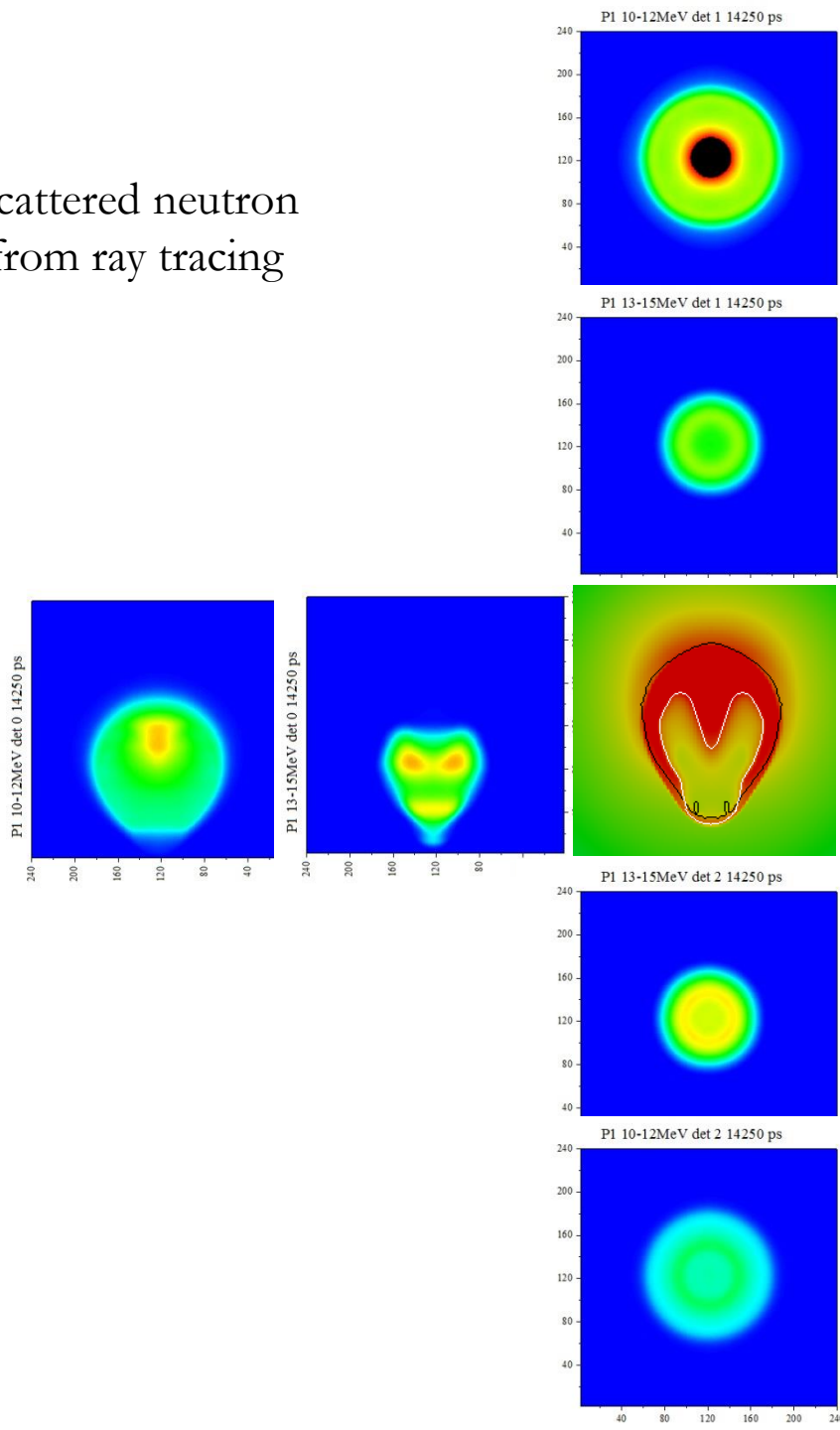
Inverse ray tracing approach



Rays perpendicular to detector plane intersect a number of scattering sites, paired with broad range of possible source sites which emit spectra drawn from database of  $10^6$  source spectra for different  $T_{ion}$ , velocity and angle.

P1

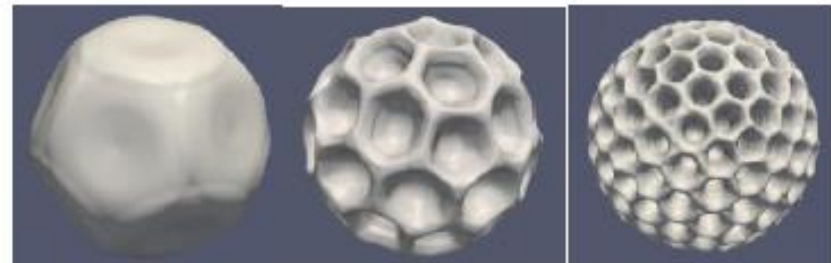
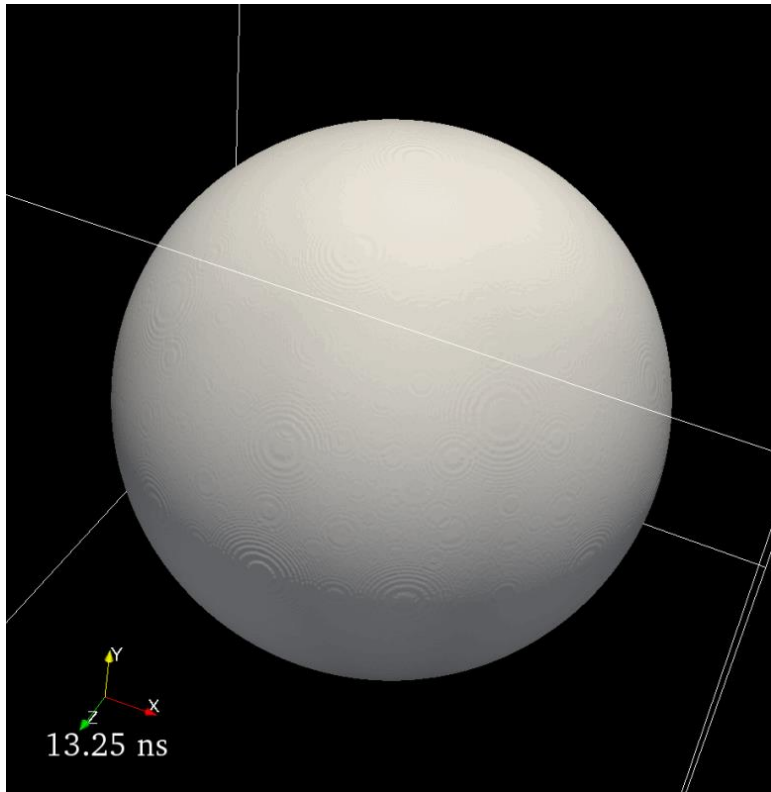
Down scattered neutron images from ray tracing



DSR from ray tracing

## Multi-mode

Attempt to generate more isotropic velocity variance by inducing quasi-turbulent flow



Geodesic  $N=12$ ,  $N=42$  &  $N=162$

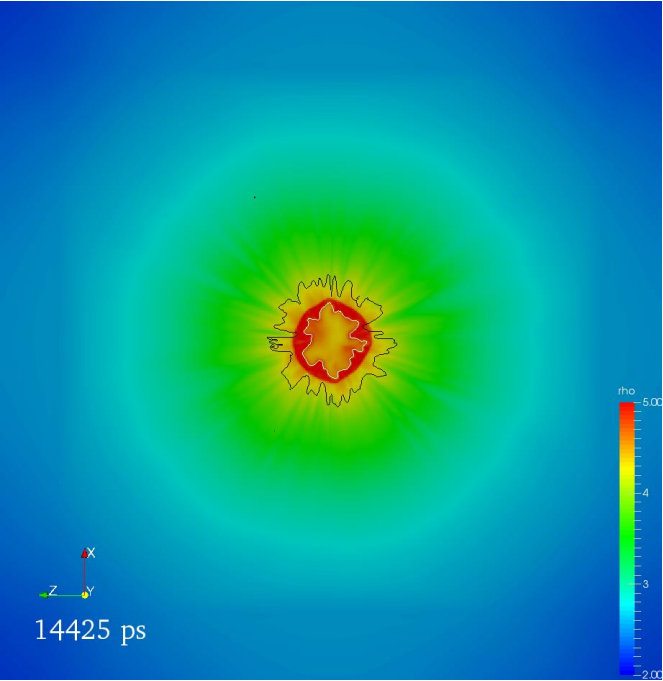
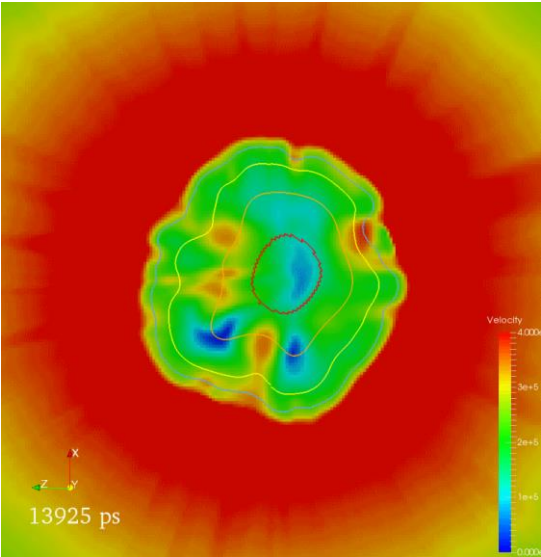
single-mode perturbations

S. Taylor & J. Chittenden *Phys. Plasmas* **21** (2014)

D. Layzer - *Astrophysical J.* **122** 1 (1955)

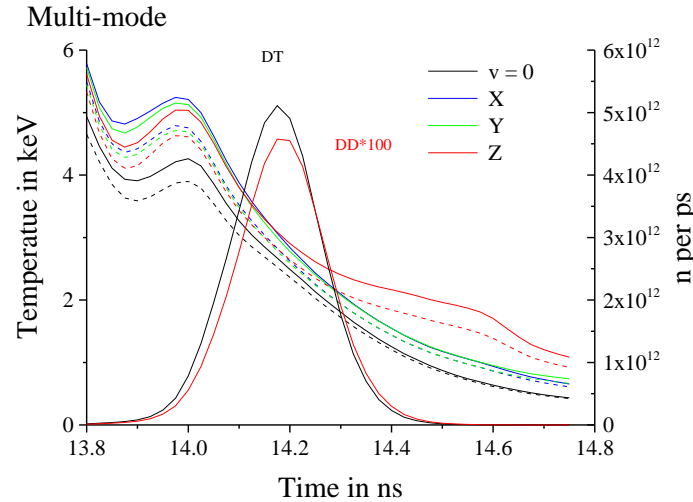
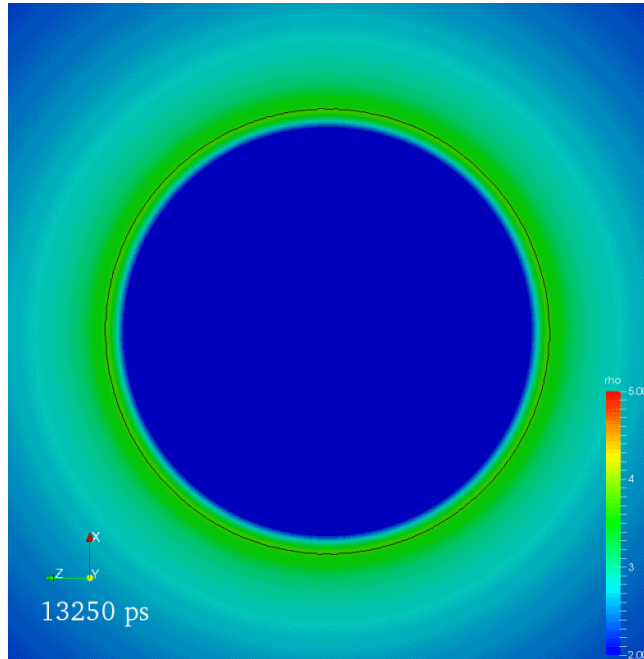
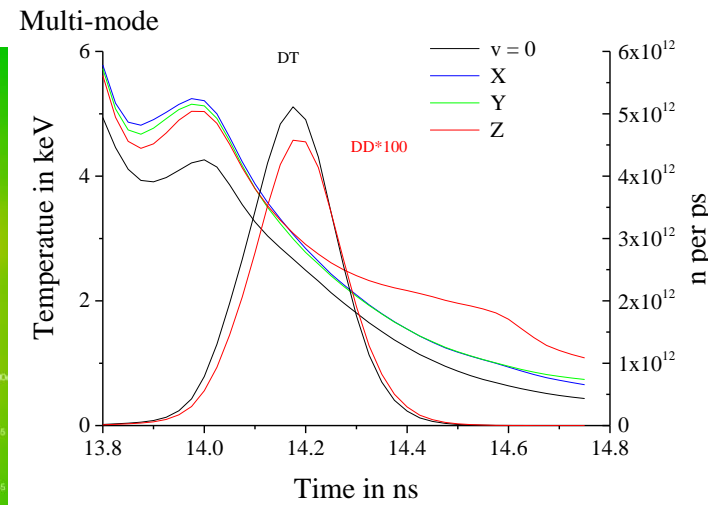
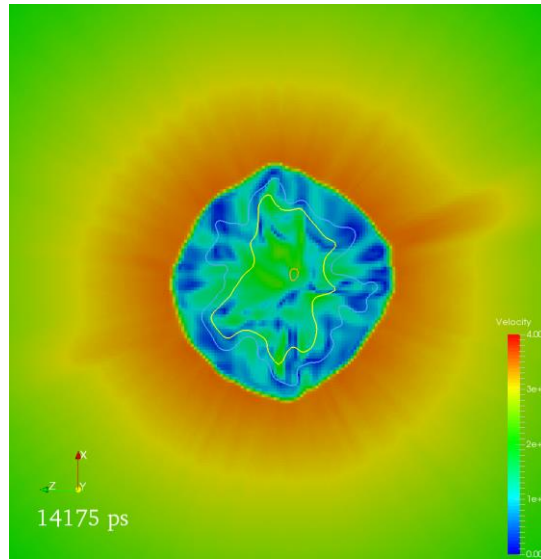
# Multi-mode

Attempt to generate more isotropic velocity variance by inducing quasi-turbulent flow



# Multi-mode

Attempt to generate more isotropic velocity variance by inducing quasi-turbulent flow

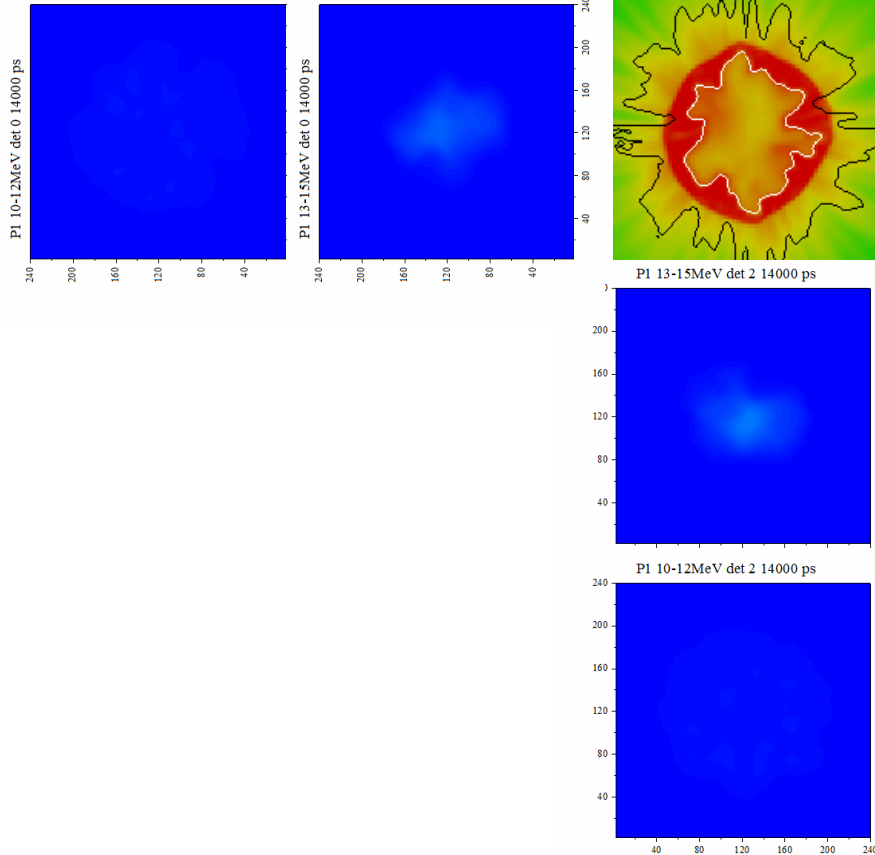


| Direction | Ti (DT) | Ti (DD) |
|-----------|---------|---------|
| X         | 2.918   | 2.598   |
| Y         | 2.865   | 2.569   |
| Z         | 3.073   | 2.701   |

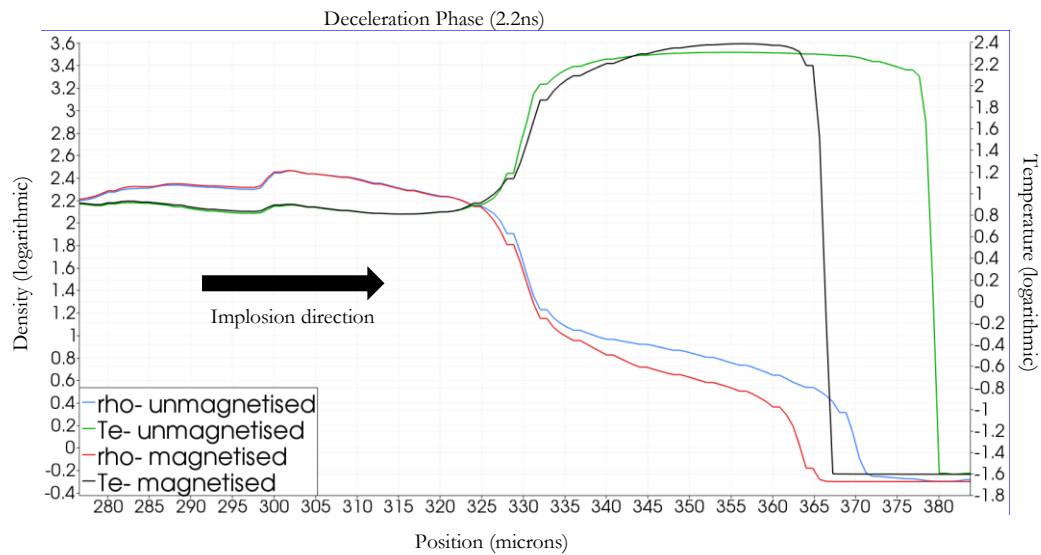
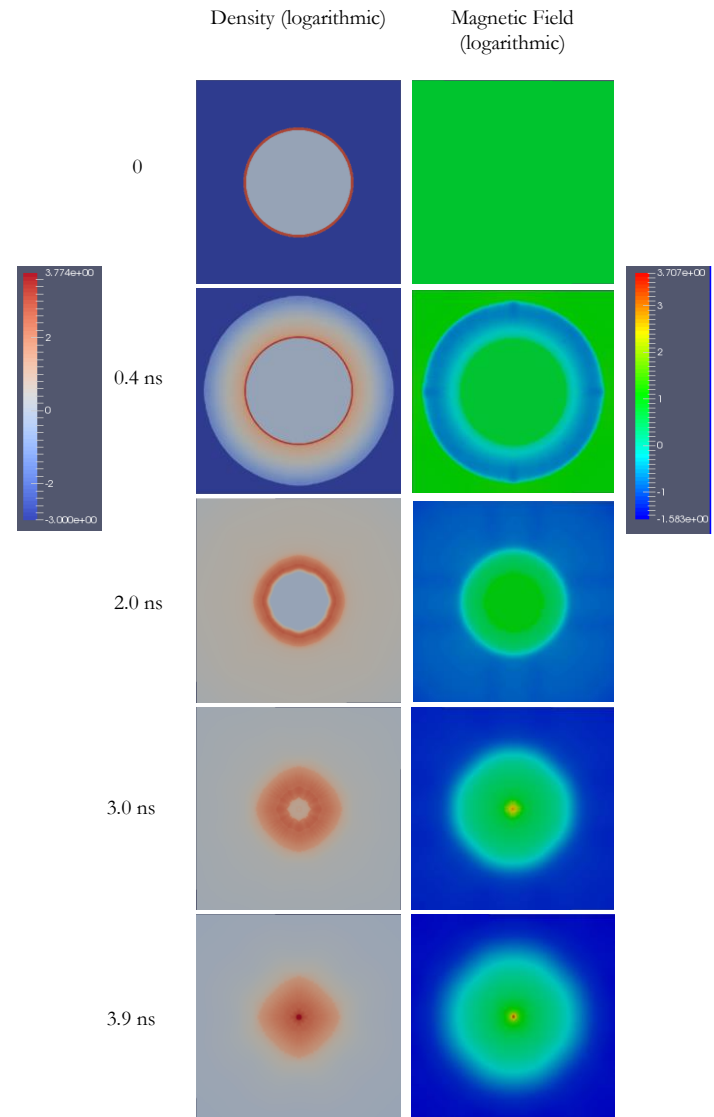
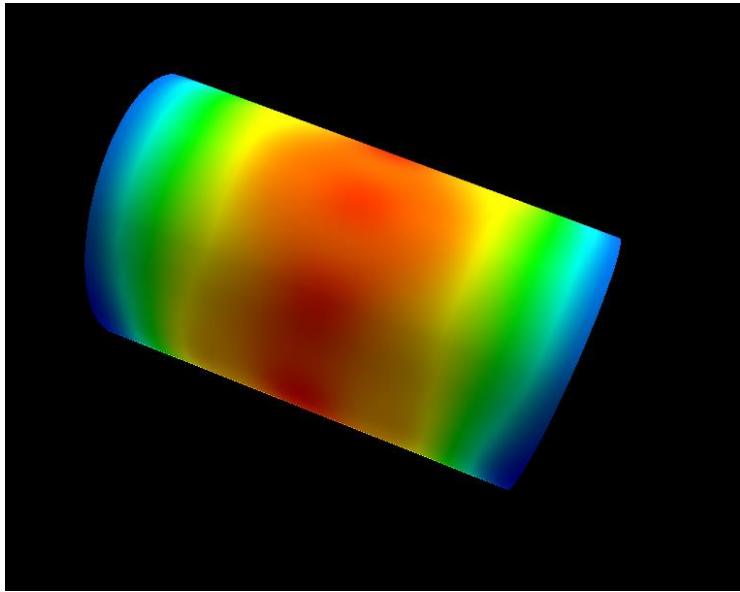


# Multi-mode

Time dependent primary and down scattered images, show a fairly consistent structure despite plasma motion



# Direct drive magnetised cylinder on Omega (J. Knauer et. al.)



## High Energy Density Physics modelling at Imperial College.

J.P. Chittenden, B.D. Appelbe, N.P.L. Niasse, J. Pecover, K. McGlinchey,  
C. Walsh, D. Botero-Garcia, J. Tong, A. Seaton, F. Manke

Centre for Inertial Fusion Studies, Imperial College, U.K.  
j.chittenden@imperial.ac.uk

