

Belle II: flavour physics at the intensity frontier

Jim Libby (IIT Madras) University of Warwick Seminar 9th October 2018



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Overview

- Particle physics and frontiers
- Some flavour history
 - Flavour as a predictor
 - Belle
 - Complementarity with LHCb
- Belle II
 - Highlights of the instrumentation and first results
 - Some physics highlights
- Conclusion

Overview



Overview

arXiv:1808.10567 [hep-ex]

KEK Preprint 2018-27 BELLE2-PAPER-2018-001 FERMILAB-PUB-18-398-T JLAB-THY-18-2780 INT-PUB-18-047 UWThPh 2018-26

The Belle II Physics Book

E. Kou^{74,¶,†}, P. Urquijo^{142,§,†}, W. Altmannshofer^{132,¶}, F. Beaujean^{78,¶}, G. Bell^{119,¶}, M. Beneke^{111,¶}, I. I. Bigi^{145,¶}, F. Bishara^{147,16,¶}, M. Blanke^{49,50,¶}, C. Bobeth^{110,111,¶},

Probably a Y(4S) event

The standard model



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The standard model





Problems

• Empirical

- Neutrinos are massive
- Dark matter
- Dark energy!!!!
- Matter rather than antimatter
- Gravity

<u>Aesthetic</u>

- Why three of everything?
- Why eighteen parameters?
 - Many with a distinct hierarchy?
- Why do we need to know them to 18 decimal places?
- Unification





ATLAS SUSY Searches* - 95% CL Lower Limits

December 2017

	Model	e, μ, τ, γ	Jets	E ^{miss} _T	∫ <i>L dt</i> [fb	⁻¹] Mass limit	$\sqrt{s} = 7, 8$	TeV $\sqrt{s} = 13 \text{ TeV}$	Reference
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0}$ $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0}$ (compressed) $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_{1}^{0}$	0 mono-jet	2-6 jets 1-3 jets 2-6 jets	Yes Yes	36.1 36.1	 <i>q</i> <i>q</i> <	1.57 TeV	$m(\tilde{\chi}_{1}^{0}) < 200 \text{ GeV}, m(1^{\text{st}} \text{ gen.} \tilde{q}) = m(2^{\text{nd}} \text{ gen.} \tilde{q})$ $m(\tilde{q}) - m(\tilde{\chi}_{1}^{0}) < 5 \text{ GeV}$ $m(\tilde{\chi}_{2}^{0}) < 200 \text{ GeV}$	1712.02332 1711.03301 1712.02332
	$gg, g \rightarrow qq \chi_1$ $\tilde{a}\tilde{a}, \tilde{a} \rightarrow aa \tilde{X}^{\pm} \rightarrow aa W^{\pm} \tilde{X}^0$	0	2-6 jets	Yes	36.1	8 ğ	2.02 TeV	$m(\tilde{x}_{1}^{0}) < 200 \text{ GeV}$ $m(\tilde{x}_{1}^{0}) < 200 \text{ GeV}$ $m(\tilde{x}^{\pm}) = 0.5(m(\tilde{x}_{1}^{0}) + m(\tilde{v}))$	1712.02332
	$gg, g \to qqq, 1 \to qq m \times 1$ $\tilde{g}\tilde{g}, \tilde{g} \to q\tilde{q}(f)\tilde{\chi}^0,$	ee, µµ	2 jets	Yes	14.7	o ĝ	1.7 TeV	$m(\tilde{x}_{1}^{0}) < 300 \text{ GeV}$	1611.05791
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow ga(\ell\ell/\gamma v)\tilde{\chi}_1^0$	3 e, µ	4 jets	-	36.1	Ĩ	1.87 TeV	$m(\tilde{\chi}_1^0) = 0 \text{ GeV}$	1706.03731
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q q W Z \tilde{\chi}_1^0$	0	7-11 jets	Yes	36.1	Ĩ	1.8 TeV	$m(\tilde{\chi}_{1}^{0}) < 400 \text{GeV}$	1708.02794
	GMSB (Î NLSP)	1-2 <i>τ</i> + 0-1 ℓ	0-2 jets	Yes	3.2	ğ	2.0 TeV		1607.05979
	GGM (bino NLSP)	2γ	-	Yes	36.1	ĝ	2.15 Te	🗸 cτ(NLSP)<0.1 mm	ATLAS-CONF-2017-080
	GGM (higgsino-bino NLSP)	γ	2 jets	Yes	36.1	Ĩ	2.05 TeV	$m(\tilde{\chi}_{1}^{0})=1700 \text{ GeV}, c\tau(\text{NLSP})<0.1 \text{ mm}, \mu>0$	ATLAS-CONF-2017-080
	Gravitino LSP	0	mono-jet	Yes	20.3	F ^{1/2} scale 865 GeV		$m(\tilde{G})>1.8 \times 10^{-4} \text{ eV}, m(\tilde{g})=m(\tilde{q})=1.5 \text{ TeV}$	1502.01518
len. ed.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\bar{b}\tilde{\chi}_{1}^{0}$	0	3 b	Yes	36.1	Ĩ.	1.92 TeV	m($\tilde{\chi}_{1}^{0}$)<600 GeV	1711.01901
a g	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 <i>e</i> , <i>µ</i>	3 b	Yes	36.1	Ĩ	1.97 TeV	$m(\tilde{\chi}_1^0) < 200 \text{GeV}$	1711.01901
~ ~~	7 7 7 . L ²⁰	0	2.6	Vac	26.1	ĩ. 050 CoV			1708 00266
S L	$b_1 b_1, b_1 \rightarrow b \mathcal{X}_1$ $\tilde{b}_1 \tilde{b}_2, \tilde{b}_3 \rightarrow t \tilde{\mathcal{X}}^{\pm}$	2 e u (SS)	1 h	Voc	36.1	δ. 275-700 GeV		$m(\tilde{r}_1) < 420 \text{ GeV}$ $m(\tilde{r}_2^0) < 200 \text{ GeV}$ $m(\tilde{r}_2^{\pm}) = m(\tilde{r}_2^0) + 100 \text{ GeV}$	1706.03200
ctic	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b \tilde{\chi}^{\dagger}_1$	0-2 e,μ	1-2 <i>b</i>	Yes 4	4.7/13.3	<i>t</i> ₁ 117-170 GeV 200-720 GeV		$m(\tilde{\chi}_{1}^{\pm}) = 2m(\tilde{\chi}_{1}^{0}), m(\tilde{\chi}_{1}^{0}) = 55 \text{ GeV}$	1209.2102, ATLAS-CONF-2016-077
3rd gen. squ direct produ	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0 \text{ or } t \tilde{\chi}_1^0$	$0-2 e, \mu$ (0-2 jets/1-2	b Yes 2	20.3/36.1	<i>ī</i> ₁ 90-198 GeV 0.195-1.0 TeV		$m(\tilde{\chi}_{1}^{0}) = 1 \text{ GeV}$	1506.08616, 1709.04183, 1711.11520
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet	Yes	36.1	ĩ ₁ 90-430 GeV		$m(\tilde{t}_1)-m(\tilde{\chi}_1^0)=5 \text{ GeV}$	1711.03301
	$\tilde{t}_1 \tilde{t}_1$ (natural GMSB)	2 e, µ (Z)	1 <i>b</i>	Yes	20.3	Ĩ ₁ 150-600 GeV		$m(\tilde{\chi}_1^0)>150 GeV$	1403.5222
	$\tilde{t}_2 \tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 <i>b</i>	Yes	36.1	ĩ ₂ 290-790 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1706.03986
	$\tilde{t}_2 \tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$	1-2 e,μ	4 <i>b</i>	Yes	36.1	Ĩ ₂ 320-880 GeV		$m(\tilde{\chi}_1^0)=0 \text{ GeV}$	1706.03986
EW lirect	$\tilde{\ell}_{I,R}\tilde{\ell}_{I,R}, \tilde{\ell} \rightarrow \ell \tilde{\chi}_{1}^{0}$	2 e, µ	0	Yes	36.1			$m(\tilde{\chi}_{1}^{0})=0$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^+ \tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow \tilde{\ell} \nu(\ell \tilde{\nu})$	2 e, µ	0	Yes	36.1	<i>x</i> [±] ₁ 750 GeV		$m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_1^{\pm} \tilde{\chi}_1^{\mp} / \tilde{\chi}_2^0, \tilde{\chi}_1^{+} \rightarrow \tilde{\tau} \nu(\tau \tilde{\nu}), \tilde{\chi}_2^0 \rightarrow \tilde{\tau} \tau(\nu \tilde{\nu})$	2 τ	-	Yes	36.1	$\tilde{\chi}_{1}^{\pm}$ 760 GeV		$m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\tau}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$	1708.07875
	$\tilde{\chi}_{1}^{\pm} \tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L} \nu \tilde{\ell}_{L} \ell(\tilde{\nu}\nu), \ell \tilde{\nu} \tilde{\ell}_{L} \ell(\tilde{\nu}\nu)$	3 e,μ	0	Yes	36.1	$\tilde{X}_{1}^{\pm}, \tilde{X}_{2}^{0}$ 1.13 Te	$m(\tilde{\chi}_1^{\pm})=r$	$n(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_{1}^{\pm})+m(\tilde{\chi}_{1}^{0}))$	ATLAS-CONF-2017-039
	$\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} Z \tilde{\chi}_{1}^{0}$	2-3 e,μ	0-2 jets	Yes	36.1	$\tilde{x}_{1}^{\pm}, \tilde{x}_{2}^{0}$ 580 GeV		$m(\tilde{\chi}_{1}^{\pm})=m(\tilde{\chi}_{2}^{0}), m(\tilde{\chi}_{1}^{0})=0, \tilde{\ell} \text{ decoupled}$	ATLAS-CONF-2017-039
0	$\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} h \tilde{\chi}_{1}^{0}, h \rightarrow b \bar{b} / W W / \tau \tau / \gamma \gamma$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_{1}^{x}, \tilde{\chi}_{2}^{y}$ 270 GeV	-0	$m(\tilde{\chi}_1^{\pm})=m(\tilde{\chi}_2^0), m(\tilde{\chi}_1^0)=0, \tilde{\ell} \text{ decoupled}$	1501.07110
	$\tilde{\chi}_2^{\circ}\tilde{\chi}_3^{\circ}, \tilde{\chi}_{2,3}^{\circ} \rightarrow \ell_{\mathrm{R}}\ell$	4 e,μ	0	Yes	20.3	X _{2,3} 635 GeV	$m(\tilde{\chi}_2^0)=r$	$n(\tilde{\chi}_{3}^{0}), m(\tilde{\chi}_{1}^{0})=0, m(\tilde{\ell}, \tilde{\nu})=0.5(m(\tilde{\chi}_{2}^{0})+m(\tilde{\chi}_{1}^{0}))$	1405.5086
	GGM (wino NLSP) weak prod., $\chi_1^{\circ} \rightarrow \chi_2^{\circ}$	$\gamma G = 1 e, \mu + \gamma$	-	Yes	20.3	W 115-370 GeV		cr<1 mm	1507.05493
	GGM (bino NLSP) weak prod., $\chi_1 \rightarrow \chi_2$	G 2γ	-	res	36.1	W 1.06 lev		<i>c</i> τ<1 mm	ATLAS-CONF-2017-080
	Direct $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	Yes	36.1	<i>x</i> [±] 460 GeV		m($\tilde{\chi}_1^{\pm}$)-m($\tilde{\chi}_1^{0}$)~160 MeV, τ ($\tilde{\chi}_1^{\pm}$)=0.2 ns	1712.02118
	Direct $\tilde{\chi}_1^+ \tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^\pm$	dE/dx trk	-	Yes	18.4	<i>x</i> [±] 495 GeV		$m(\tilde{\chi}_{1}^{\pm})-m(\tilde{\chi}_{1}^{0})\sim 160 \text{ MeV}, \tau(\tilde{\chi}_{1}^{\pm})<15 \text{ ns}$	1506.05332
pe s	Stable, stopped g R-hadron	0	1-5 jets	Yes	27.9	850 GeV		$m(\tilde{\chi}_1^0)=100 \text{ GeV}, 10 \mu\text{s} < \tau(\tilde{g}) < 1000 \text{ s}$	1310.6584
ij j	Stable g R-hadron	trk	-	-	3.2	8	1.58 TeV	(⁵⁰) (00 0)) (00 0	1606.05129
Long- parti	Metastable \tilde{g} R-hadron $\tilde{z} \to a a \tilde{v}^0$	de/ax trk		Voc	32.8	8 õ	1.57 TEV	$m(\ell_1)=100 \text{ GeV}, \tau>10 \text{ ns}$	1710 04901
	GMSB stable $\tilde{\tau} \tilde{\chi}^0_1 \rightarrow \tilde{\tau}(\tilde{a} \ \tilde{u}) + \tau(a \ u)$	1-2 µ	-	-	19.1	x ⁰ 537 GeV	2.31	$10 < \tan \beta < 50$	1411.6795
	GMSB $\tilde{\chi}_{1}^{0} \rightarrow \gamma \tilde{G}$ long-lived $\tilde{\chi}_{1}^{0}$	2 2	-	Yes	20.3	x ⁰ 440 GeV		$1 \le \tau(\tilde{\chi}_1^0) \le 3$ ns. SPS8 model	1409.5542
	$\tilde{g}\tilde{g}, \tilde{\chi}^0_1 \rightarrow eev/e\mu v/\mu\mu v$	displ. ee/eµ/µ	μ -	-	20.3	<i>x</i> ₁ ⁰ 1.0 TeV		$7 < c\tau(\tilde{\chi}_1^0) < 740 \text{ mm}, m(\tilde{g}) = 1.3 \text{ TeV}$	1504.05162
	$ \text{EV} nn \rightarrow \tilde{v}_{-} + X \tilde{v}_{-} \rightarrow eu/e\tau/u\tau$	eu.et.ut			3.2	ũ.	1.9 ToV	a'=0.11_dunum=0.07	1607 09079
	Bilinear BPV CMSSM	2 e. µ (SS)	0-3 h	Yes	20.3	\tilde{q}, \tilde{g}	1.45 TeV	$m(\tilde{a})=m(\tilde{a}), c_{T,S,R}<1 \text{ mm}$	1404 2500
	$\tilde{X}_{1}^{\dagger}\tilde{X}_{1}^{-}$ $\tilde{X}_{1}^{\dagger} \rightarrow W \tilde{X}_{1}^{0}$ $\tilde{X}_{1}^{0} \rightarrow eev$ evy viv	4 e.μ	-	Yes	13.3	λ [±] 1.14 Te	eV	$m(\tilde{\chi}_{1}^{0}) > 400 \text{ GeV}$ $d_{12k} \neq 0 \ (k = 1, 2)$	ATLAS-CONF-2016-075
~	$\tilde{\chi}_{1}^{\dagger}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{\dagger} \rightarrow W \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow \tau \tau \gamma_{r}, e \tau \gamma_{\tau}$	$3e, \mu + \tau$	-	Yes	20.3	X [±] 450 GeV		$m(\tilde{\chi}_{1}^{0}) > 0.2 \times m(\tilde{\chi}_{1}^{\pm}), \lambda_{133} \neq 0$	1405.5086
d	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow ga\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow gag$	0 4-	-5 large- <i>R</i> je	ets -	36.1	Ĩ	1.875 TeV	$m(\tilde{\chi}_{1}^{0})=1075 \text{ GeV}$	SUSY-2016-22
ιL.	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qqq$	1 <i>e</i> ,μ 8	-10 jets/0-4	b -	36.1	ĝ	2.1 TeV	$m(\tilde{\chi}_{1}^{0}) = 1 \text{ TeV}, \lambda_{112} \neq 0$	1704.08493
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	1 <i>e</i> ,µ 8	-10 jets/0-4	b -	36.1	ğ	1.65 TeV	m(t̃ ₁)= 1 TeV, λ ₃₂₃ ≠0	1704.08493
	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b	-	36.7	100-470 GeV 480-610 GeV			1710.07171
	$\bar{t}_1\bar{t}_1, \bar{t}_1 \rightarrow b\ell$	2 <i>e</i> , µ	2 b	-	36.1	<i>ī</i> ₁ 0.4-	-1.45 TeV	$BR(\bar{t}_1 \rightarrow be/\mu) > 20\%$	1710.05544
Other	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_1^0$	0	2 <i>c</i>	Yes	20.3	δ 510 GeV		$m(\tilde{\chi}_1^0)$ <200 GeV	1501.01325
*Only phen simp	Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.								

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ATLAS Preliminary $\sqrt{s} = 7, 8, 13 \text{ TeV}$

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ISTER MARGARET'S PICK DEAD POOL AGE DAVE \$50 COSBY, B \$12.01 JAK SHEEN BET \$75 78 WEASEL PICK WEST, 50 \$2.00 AGE \$110 MARC TAD WILLSON, ROCK 38 3175 \$190 RANDY TYSONTH WADE PUTIN 45 FRANKIE \$67 9150 NELSON T BOOTHE 49 62 \$120 MOLE WAYNE 56 21 JOHN \$60 JEMBO LABEOUF, S 33 \$110 \$80 RETNOLDS, R HANK BEATTY, N 29 \$190 38 MILLER, TJ REEVES 78 \$77 35 DIEFELD, R LOHAN, L JIRIK \$35 48 OSBONKNE, O \$80 MOORS 67 \$150 BYNES, A GRIGGS 29

ISTER MARGARET'S PICK DEAD POOL AGE DAVE \$50 CMSSM \$12.01 JAK SHEEN BET \$75 78 WEASEL PICK WEST, 50 \$2.00 AGE \$110 MARC TAD WILLSOW ROCK 38 3175 \$190 RANDY TYSONTH WADE PUTEN 45 FRANKIE \$67 9150 NELSON, T 49 2HDM II 62 \$120 MOLE WAYNE 56 21 JOHN \$60 JEMBO LABEOUR, S 33 \$110 \$80 RETNOLDS, R HANK BEATTY, N 29 \$190 38 MILLERTJ REEVES 78 \$77 35 DIEFELD, R JIRIK **RPV SUSY** \$35 48 OSBONKNE, O \$80 MOORS 67 \$150 BYNES, A GRIGG 29





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Problems

• Empirical

- Neutrinos are massive
- Dark matter
- Dark energy!!!!
- Matter rather than antimatter
- Gravity

<u>Aesthetic</u>

- Why three of everything?
- Why eighteen parameters?
 - Many with a distinct hierarchy?
- Why do we need to know them to 18 decimal places?
- Unification



Flavour physics – history of discovery

- Particle zoo of mesons and baryons discovered in 1950s and early 1960s lead to the quark model
 - up (u)
 - down (d)
 - strange (s)
- An allowed but rare decay such as

$$K_L^0(s\overline{d}) \to \mu^+\mu^-$$

Predicted but not seen!

$$\frac{s}{d} \ u \ W^{-} \ v \ \mu^{+}$$

Flavour physics – history of discovery

$$\frac{\sin \theta_{c}}{S} \qquad W^{-} \qquad \psi^{\mu}_{\mu^{+}} \qquad \psi^{\mu}_{\mu^{+}} \\ \frac{\cos \theta_{c}}{\cos \theta_{c}} \qquad \frac{S}{d} \qquad C \qquad W^{-} \qquad \psi^{\mu}_{\mu^{+}} \qquad \psi^{\mu}_{\mu^{+}} \qquad \frac{S}{d} \qquad W^{-} \qquad \psi^{\mu}_{\mu^{+}} \qquad \frac{W^{-}}{\mu^{+}} \qquad \frac{W^{-$$

 $-\sin\theta_c$



Glashow

liopoulos

Maiani

Phys. Rev. D 2, 1285 (1970) 2 ∞ Rate ~ O

 $m_c > m_\kappa$

Such rare virtual processes tell you about higher energy particles

ARGUS: B mixing \Rightarrow heavy top

OBSERVATION OF B⁰-B⁰ MIXING

ARGUS Collaboration



reconstructed event consisting of the decay Y

econstructed event consisting of the decay



m_t> 50 Gev



and $B_2^0 \rightarrow D_2^{*-} \mu_2^+ \nu_2$ \downarrow $D_2^{*-} \rightarrow \pi^0 D^ \downarrow$ $D^- \rightarrow K_2^+ \pi_2^- \pi_2^- .$

ARGUS: B mixing \Rightarrow heavy top

OBSERVATION OF B⁰-B⁰ MIXING

ARGUS Collaboration



CKM matrix

$$\begin{pmatrix} u & c \end{pmatrix} \begin{bmatrix} \cos \theta_C & \sin \theta_C \\ -\sin \theta_C & \cos \theta_C \end{bmatrix} \begin{pmatrix} d \\ s \end{pmatrix}$$

 Two by two mixing matrix proposed Cabibbo

CKM matrix

- Two by two mixing matrix proposed Cabibbo
 - Kobayashi-Maskawa proposed third generation to explain observed CP violation by Cronin and Fitch
- 3 × 3 unitary complex matrix
 - 4 parameters
 - 3 mixing angle and 1 phase
- Intergenerational coupling disfavoured





Visualising CP violation:
the unitarity triangle
1)
$$\begin{pmatrix} 1-\lambda^2/2 \\ -\lambda \\ \lambda^3 [1-(\rho-i\eta)] \end{pmatrix}$$
 $\begin{pmatrix} \lambda \\ 1-\lambda^2/2 \\ -A\lambda^2 \end{pmatrix}$ $\begin{pmatrix} A\lambda^3 (\rho-i\eta) \\ A\lambda^2 \\ 1 \end{pmatrix} + O(\lambda^4)$
2) Exploit unitarity (1st and 3rd col.) $V_{ud}V_{ub}^* + V_{cd}V_{cb}^* + V_{td}V_{tb}^* = 0$
3) $V_{ud}V_{ub}^*$ $\begin{pmatrix} V_{td}V_{tb}^* \end{pmatrix}$ $\phi_1 = \beta = \arg\left(-\frac{V_{cd}V_{cb}^*}{V_{td}V_{tb}^*}\right)$

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 $\simeq \arg\left(\frac{1}{1-\rho-i\eta}\right)$

 $V_{cd}V_{cb}^*$

 γ

β

Belle

- Operation from 1999 to 2010
- $e^+e^- \rightarrow \Upsilon(4S) \rightarrow B\overline{B}$ for CKM measurements
- Asymmetric energy to allow time-dependent measurements
- Coherent production of $B^0 \overline{B^0}$
- Low multiplicity
- Detectors with good tracking, PID and calorimetry
 - plus hermeticity for full event reconstruction/tagging





In SM $S_f = \sin 2\beta$ and $C_f = 0$ when no CPV in f

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Time-dependent CPV violation



Over constraint



Tree level only



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From Abi Soffer: HEPMAD 1800 1st dark searches Nobel prize to KM / 1600 Decisive confirmation of CKM picture Observation of direct 1400 CP violation in B $\rightarrow \pi^+\pi^-$ Excess in 1200 $\bar{B} \rightarrow D^{(*)} \tau \bar{\nu}$ Observation of

Integrated Luminosity in fb⁻¹ Time reversal asymmetry Observation of 1000 $b \rightarrow d\gamma$ Evidence for CP violation in D⁰ mixing B-meson system 800 Evidence for B→τv Observation of 600 B → K(*)II Evidence for direct 400 CP violation in B \rightarrow K⁺ π ⁻ Measurements of mixing-induced 200 CP violation in $B \rightarrow \phi K_s, \eta' K_s, ...$ Exotic hadrons С 1999 2000 2001 2002 2003 2004 2005 2006 2007 2008 2009 2010 Year >100 unique CPV results ~350 papers published after shutdown, 21 in 2018

Belle achievements

Belle II: can never have too much of a good thing (x 50 Belle)

• But isn't LHCb doing this already?

Property	LHCb	Belle II	
$\sigma_{b\bar{b}}$ (nb)	~150,000	~1	
$\int L dt$ (fb ⁻¹) by ~2024	~25	~50,000	
Background level	Very high	Low	
Typical efficiency	Low	High	
π^0 , K_S reconstruction	Inefficient	Efficient	
Initial state	Not well known	Well known	
Decay-time resolution	Excellent	Very good	
Collision spot size	Large	Tiny	
Heavy bottom hadrons	<i>B_s, B_c, b</i> -baryons	Partly B _s	
au physics capability	Limited	Excellent	
B-flavor tagging efficiency	3.5 - 6%	36%	

"Moore's" Law of Luminosity



The path to higher luminosity



Brute force: Increase beam currents by a factor of 5-10! Increase the beam-beam parameter by a factor of a few (crab cavities). Too hard, too expensive (power, melt beam pipes)

The path to higher luminosity



(1) Smaller β_{y}^{*} (20 x)

(2) Increase beam currents (~2-3x)





SUPERKEKB



Schedule and status





Phase 2 goals:

- · Progress toward high luminosity
- Progress toward stable operation

Achievements:

- $L = 5.5 \times 10^{33} \,\mathrm{cm}^{-2} \mathrm{s}^{-1}$
- Collected ~0.5 fb⁻¹ for commissioning & calibration

Ohnishi-san eeFACT, HKUST Super KEKB performance



 $\sigma = 4.5 \text{ mm}$



σ = 550 µm

measurement at Belle II



Ohnishi-san eeFACT, HKUST Super KEKB performance



Belle II Collaboration





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Belle II - TOP

Simulation of a 2 GeV pion and kaon interacting in a quartz bar.



Belle II - TOP

Channel Vs. time for 3GeV pions/kaons with beam test serup



At 3 GeV <u>Timing at the ~100 ps level is</u> needed to separate pion and Kaon





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At 3 GeV Timing at the ~100 ps level is needed to separate pion and Kaon

ToP signature of kaon identified kinematically via $D^{*+} \rightarrow D^0 \pi_*^+; D^0 \rightarrow K^- \pi^+$

is visibly more consistent with being a kaon than a pion or proton





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SVD performance



Belle a factor two worse than Belle II

Stand alone SVD track finding efficiency good for K_s finding (30% over Belle) and slow π from D*





Charm meson reconstruction



Many charm decays seen including CP eigenstates used in CP violation measurements – reasonable agreement with expectations



- We are on the Y(4S) resonance and recording B anti-B pairs with ~99% efficiency.
- Not so obvious: When we change accelerator optics, we remain on Y(4S).



B meson reconstruction









 $B \rightarrow D^{(*)}h \ (h=\pi,\rho) \text{ and } B \rightarrow J/\psi K^{(*)}$ Reconstructed > 200 B events Jniversity of Warwick Seminar 9/10/18

A FEW PHYSICS PROSPECTS

 $B \rightarrow K^*(892)l^+l^-$

- This is a rare flavour changing neutral current process
- The four-body final state allows differential distributions to be probed
 - Large new physics contributions possible as they appear via interference c.f. forward-backward asymmetries in e⁺e⁻
- Also variation with the invariant mass of the l⁺l⁻ system q²



$B \rightarrow K^*(892)l^+l^-$ nomenclature



- Goal is to measure this 4D differential distribution and extract the coefficients from data to compare to the SM predictions
- Much work on defining observables with minimal theoretical uncertainties
- Let us focus on S₅ which get normalized as $P_5' = \frac{S_5}{\sqrt{F_L(1-F_L)}}$ to minimize form factor uncertainties

LHCb

Exp.: R. Aaij et al., JHEP 02 (2016) 104 Theory: S. Descotes-Genon et al., JHEP 12 (2014) 125



3.7 σ disagreement with Standard Model

Other analyses of the data also show inconsistency i.e. RH currents at large q² A. Karan et al. arXiv:1603.04355 [hep-ph]



Belle

PRL118, 111801 (2017)



Lepton Universality Violation (LUV)



2-3 standard deviations for H = K and K*



Belle II predictions



Semi-tauonic decays

• Tree level in the SM but allows lepton universality tests



 Measure ratios to reduce theoretical and experimental uncertainties

$$R(D) = \frac{\Gamma(\overline{B} \to D\tau \nu)}{\Gamma(\overline{B} \to D\ell\nu)} \qquad R(D^*) = \frac{\Gamma(\overline{B} \to D^*\tau\nu)}{\Gamma(\overline{B} \to D^*\ell\nu)}$$

 Babar reported an anomalous result PRL 109, 101802 (2012) much activity since



Belle results



- Tag signal by fully reconstructing or identifying a semileptonic (SL) decay of the other B
- Then use residual energy in ECL, missing mass, multivariates and/or lepton momentum to separate signal
- Example: Phys. Rev. D 94, 072007 (2016) – SL tag



Phys. Rev. Lett. 115, 111803 (2015)



LHCb also in the game using their vertexing prowess



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More modes for tagging Full Event Interpretation





Belle II predictions projection for 50 ab⁻¹



Many other measurements

- CKM metrology
 - ϕ_3/γ 1.5 degrees
 - Same from LHCb
 - V_{ub}-1.2%
- Other rare decays
 - B→τν 1.5-2.0%
 - B→µν 5%
 - $B \rightarrow X_s |^+|^- R_x 3-5\%$
 - $b \rightarrow s \tau \tau$, $b \rightarrow s \nu \nu$ and LFV versions
- CPV gluonic penguins
 - $B \rightarrow \eta' K_{s}^{0} \sin 2\phi_{1}$ to 0.02
- LFV $\tau \rightarrow \mu \gamma$ 10⁻⁹ limit at 90% C.L
- + charm, XYZ spectroscopy, dark photon





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Conclusion

- Particle physics is tackling its problems on three complementary frontiers
 - 1. Energy
 - 2. Cosmic
 - 3. Intensity
- Flavour physics has played a significant role in the development of the Standard Model
- **Belle II** is a project that will continue flavour physics at the intensity frontier until the middle of the next decade along with LHCb
 - First collisions this year much more to come..