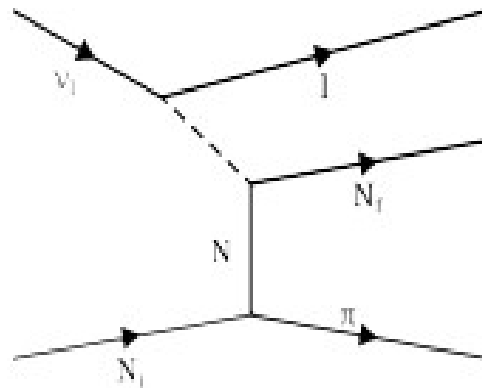
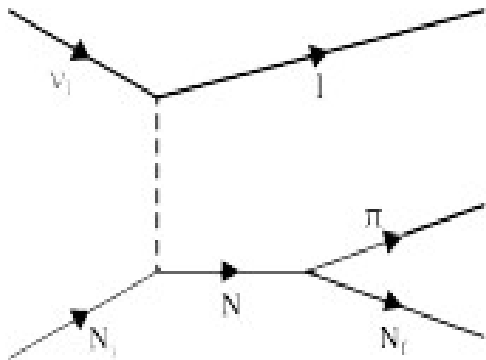
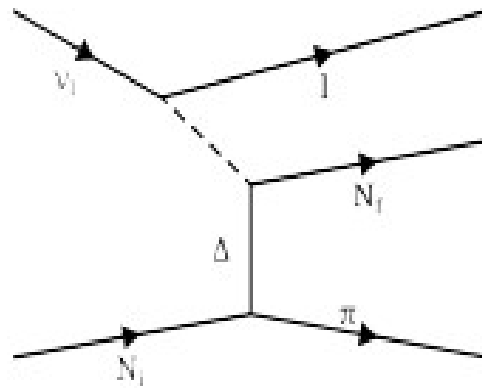
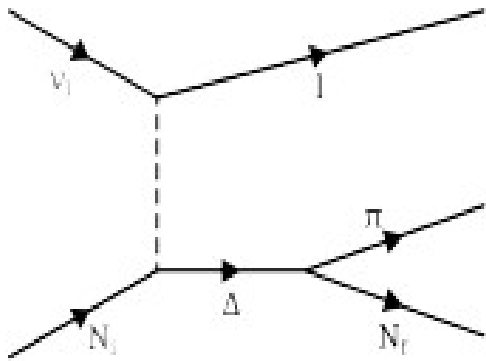


Implementation of the Alvarez-Ruso microscopic coherent model

Steve Boyd , Dan Scully , Steve Dennis

- ▶ Brief model description
- ▶ Implementation with lookup tables
- ▶ Implementation without lookup tables

Microscopic models



- Delta/Nucleon production
- Final state nucleon in the same state as initial nucleon
- A number of models on the market

- ▶ Alvarez-Ruso
- ▶ Nieves
- ▶ Hernandez

Models apply below approx. 5 GeV

None (up to now) have been included in the generators

Model

$$\frac{d^5 \sigma}{dE_l d\Omega_l d\Omega_\pi} = \frac{1}{8(2\pi)^5} \frac{|p_l||p_\pi|}{|p_\nu|} |\overline{M}|^2$$

Leptonic current is standard : $j_l^\mu = \bar{u}_l \gamma^\mu (1 - \gamma_5) u_\nu$

Hadronic current is more complicated :

$$j_{p_\pi}^\mu = \frac{f^*}{m_\pi} p_\pi^\alpha F(p_\Delta) \overline{u_{N,f}} D(p_\Delta) \Lambda_{\alpha\beta} A^{\beta\mu} u_{N,i}$$

Model

$$j_{p\pi}^{\mu} = \frac{f^*}{m_{\pi}} p_{\pi}^{\alpha} F(p_{\Delta}) \overline{u}_{N,f} D(p_{\Delta}) \Lambda_{\alpha\beta} A^{\beta\mu} u_{N,i}$$

f^* : $\Delta \rightarrow \pi N$ decay constant

$P_{\{\pi,\Delta\}}$: pion, Delta momenta

$F(p_{\Delta})$: Δ - π N vertex form factor

$D(p_{\Delta})$: In-medium delta propagator

$\Lambda_{\alpha\beta}$: spin 3/2 projection operator

$A^{\beta\mu}$: N- Δ coupling

$$A^{\mu\alpha} = \left\{ \frac{C_3^V}{m_N} (g^{\mu\alpha} \gamma_5 q^{\delta} - q^{\mu} \gamma^{\alpha}) + \frac{C_4^V}{m_N^2} (g^{\mu\alpha} q \cdot P_{\Delta} - q^{\mu} P_{\Delta}^{\alpha}) + \frac{C_5^V}{m_N^2} (g^{\mu\alpha} q \cdot P_{N,i} - q^{\mu} P_{N,i}^{\alpha}) \right\} \gamma_5$$

$$+ \left\{ \frac{C_3^A}{m_N} (g^{\mu\alpha} \gamma_5 q^{\delta} - q^{\mu} \gamma^{\alpha}) + \frac{C_4^A}{m_N^2} (g^{\mu\alpha} q \cdot P_{\Delta} - q^{\mu} P_{\Delta}^{\alpha}) + C_5^A g^{\mu\alpha} + \frac{C_6^A}{m_N^2} q^{\mu} q^{\alpha} \right\}$$

Model

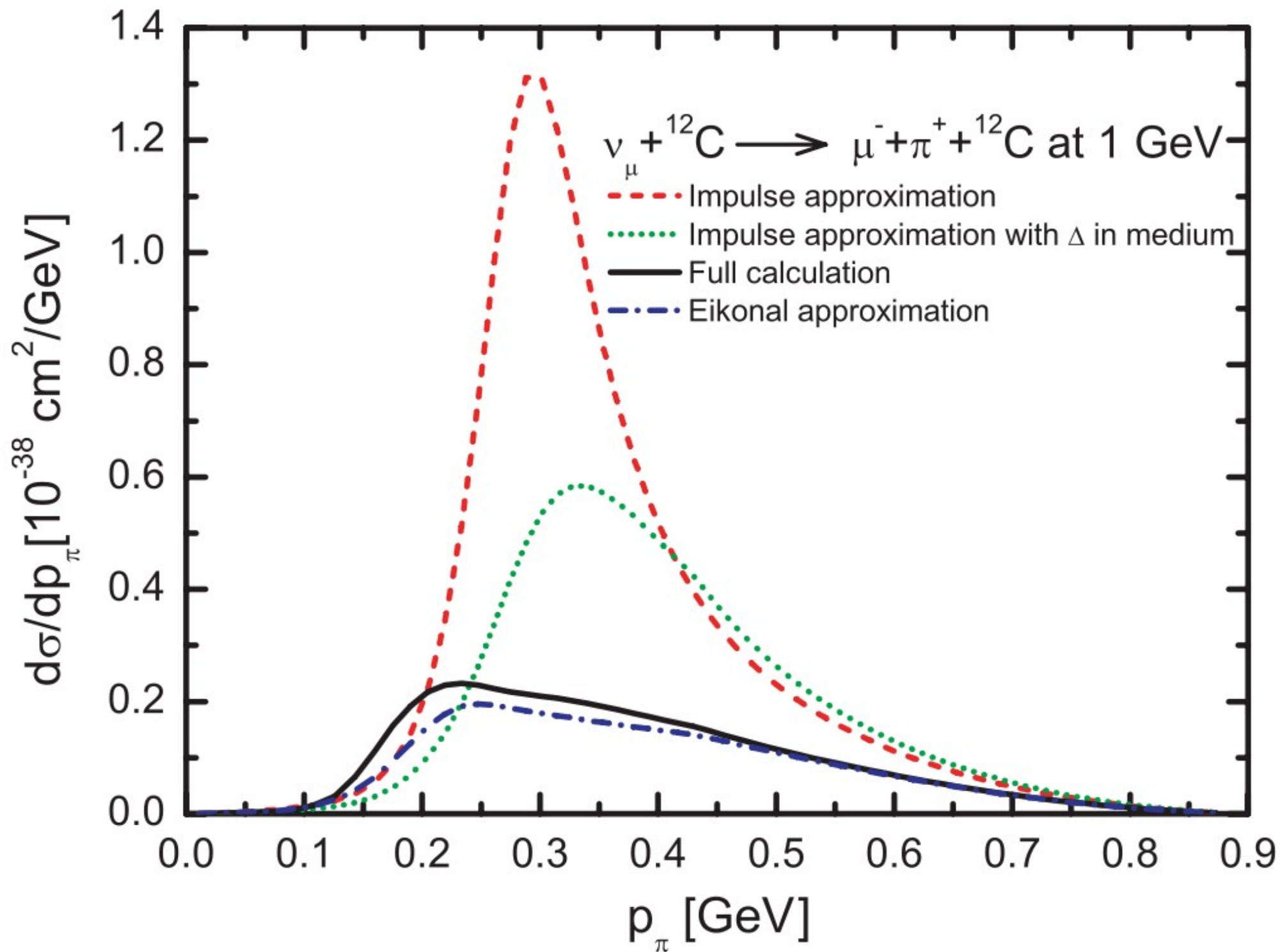
Nuclear hadronic current using a plane wave for the pion wavefunction

$$J_{A\pi}^{\mu} = \frac{-ie}{2} \int_0^{\infty} e^{i\left(\frac{\vec{q}}{q} - P_{\pi}\right) \cdot r} \left(\rho_p(r) + \frac{\rho_n(r)}{3} \right) \frac{f^*}{m_{\pi}} P_{\pi}^{\alpha} F(P_{\Delta}) D(P_{\Delta}) \text{Tr} \left\{ \bar{u}_{N,f} \Lambda_{\alpha\beta} A^{\beta\mu} u_{N,i} \right\} d^3 r$$

More realistic description modified the pion wavefunction using the *eikonal approximation*

$$\Phi_{\text{eikonal}}^* = e^{-ip_{\pi} \cdot r} \exp \left[-i \int_r^{\infty} \frac{\Pi(r')}{2p_{\pi}} d r' \right]$$

Pion self energy



Implementation

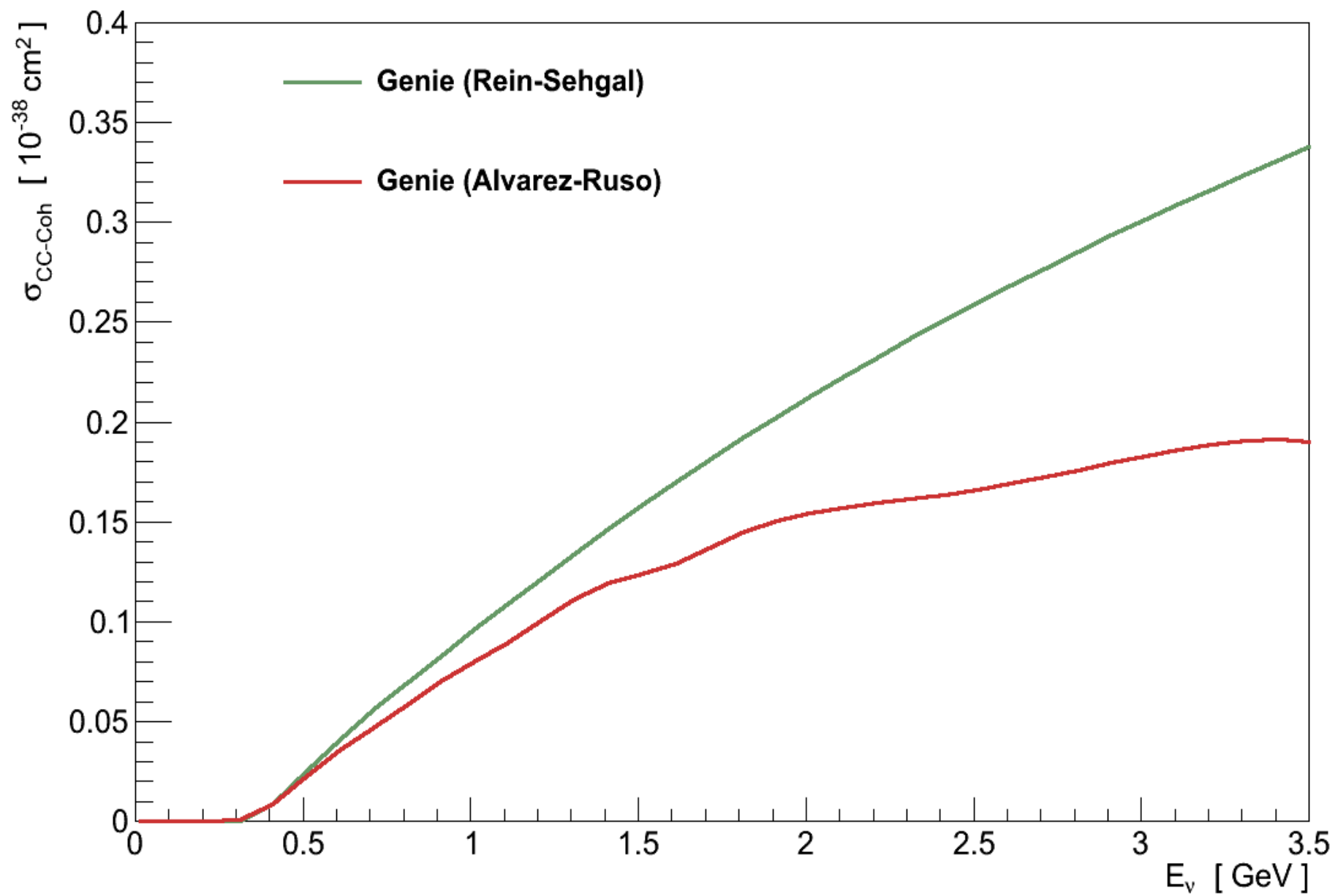
- ▶ Model was copied from Luis' FORTRAN calculation
- ▶ Standalone implementation was used to verify the translation to C++
- ▶ GENIE Nuclear utils are used
- ▶ One of the polar angles is free : so can reduce the 5-D integral to a 4-D integral

$$\frac{d^5 \sigma}{dE_l d\Omega_l d\Omega_\pi} \rightarrow 2\pi \frac{d^4 \sigma}{dE_l d\Omega_l d\theta_\pi}$$

- ▶ Calculation currently cuts off at 5 GeV

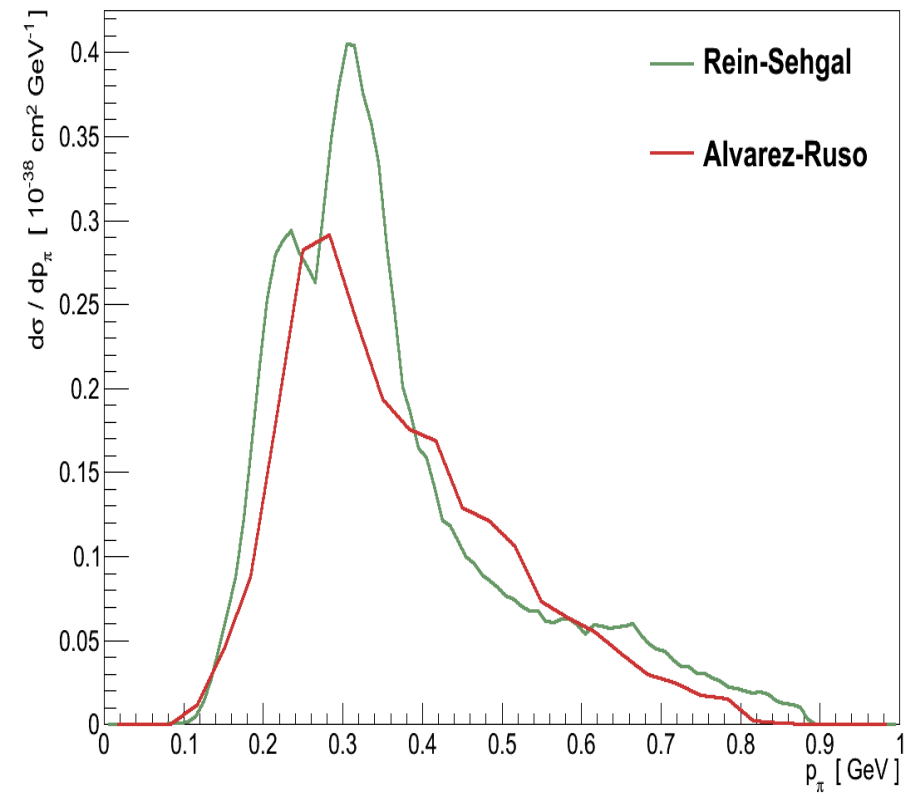
Implementation 1 : Lookup tables

- ▶ For Dan's thesis we needed the AR model to be relatively fast
- ▶ The only way to do this was to pre-calculate the pion wavefunctions and store them in look-up tables
- ▶ One look-up table per (neutrino, target) combination
 - ▶ Approximately 0.05 s per cross section evaluation
 - ▶ Takes approximately 10 hours to generate a cross-section spline
 - ▶ Size of each table is about 25 Mb. Tables are only text and this could be optimised a lot.

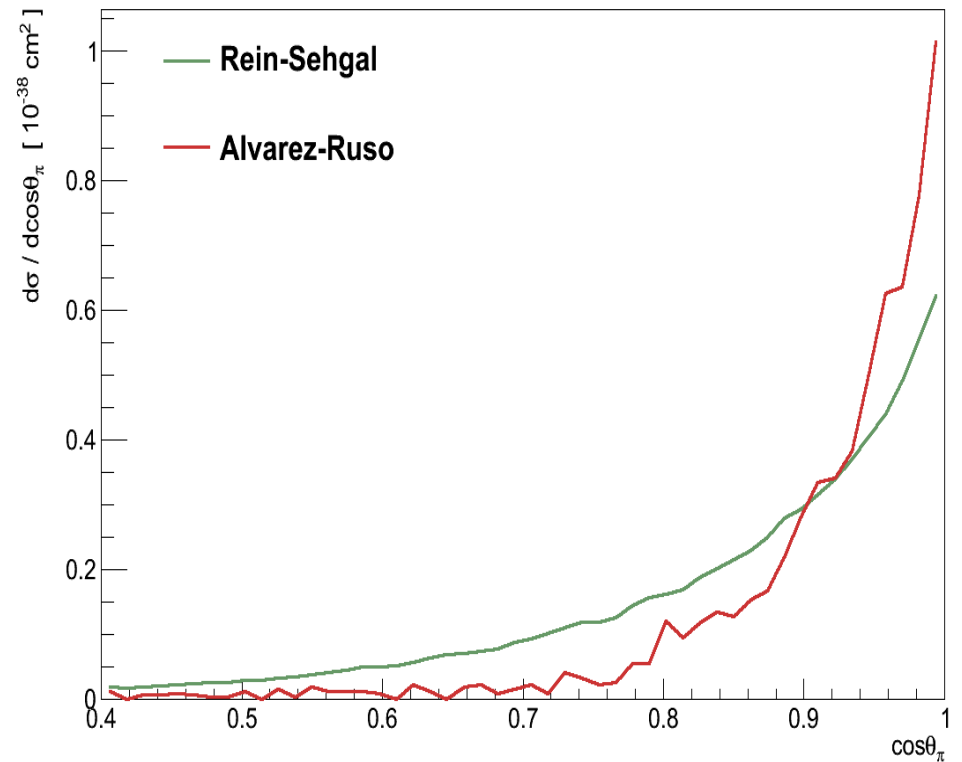


ν_μ on C14

Some kinematic distributions



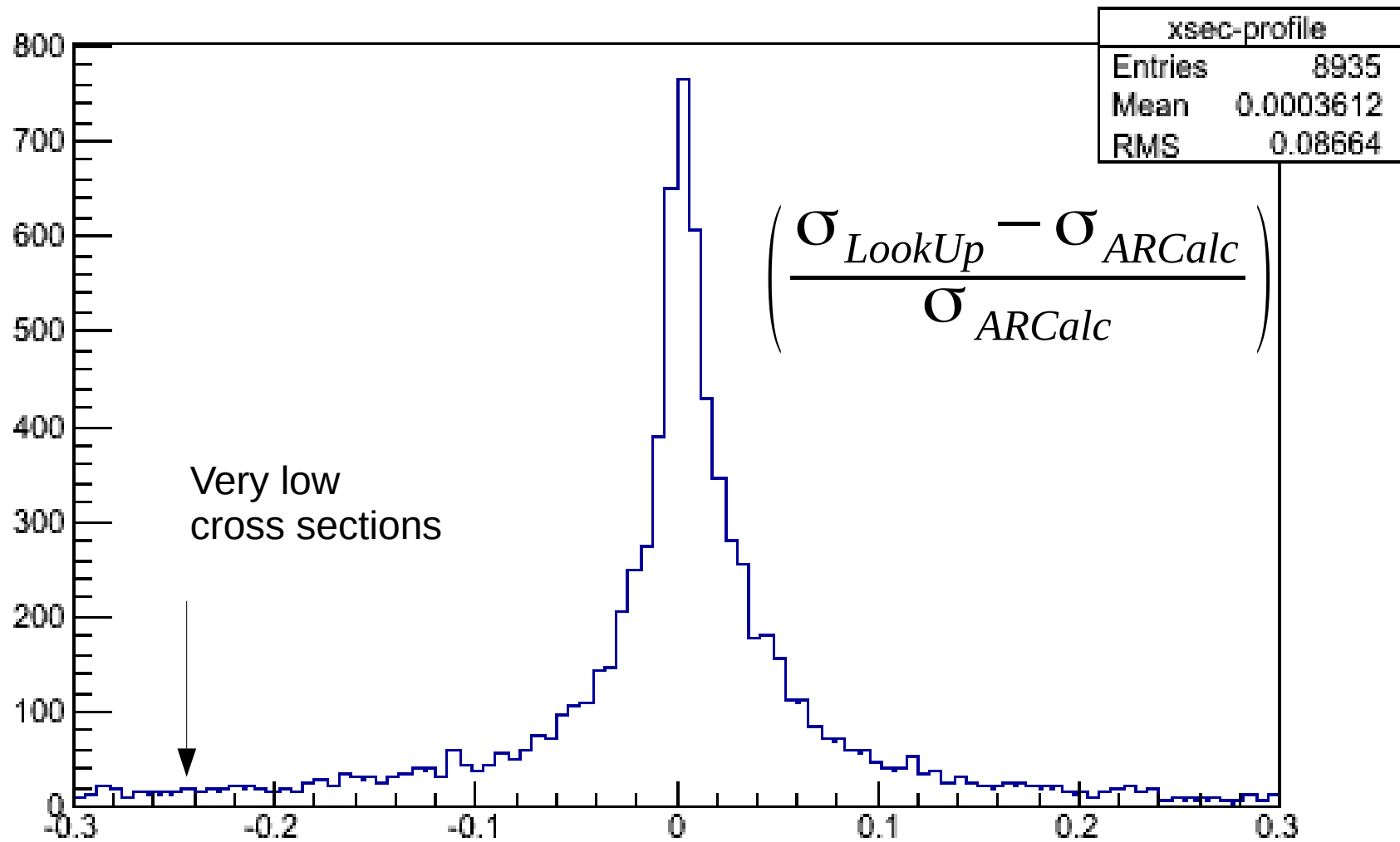
p_π



$\cos\theta_\pi$

ν_μ on C12 @ $E_\nu = 1.0 \text{ GeV}$

But is it accurate?



Implementation 2 : Using wavefunction calculation

- ▶ Moving from lookup tables to wave function calculation on an event by event basis slows down the code by about a factor of 10 : one evaluation takes about 0.5 seconds now.
- ▶ Still, for event generation this is probably endurable.
- ▶ The largest problem we have is processing time in the spline calculation.
- ▶ We've implemented this by splitting the integral. The wavefn is only dependent on E_{π} so only needs to be calculated when E_{π} changes

Implementation 2

$$\sigma(E_\nu) = 2\pi \int \frac{d\sigma}{dE_l} \left(\int \int \int \frac{d^3\sigma}{d\Omega_\mu d\theta_\pi} d\Omega_\mu d\theta_\pi \right) dE_l$$

This may have sped up the spline calculation.

I can't tell because we are still running the first spline test after 1.5 days

Incl.	Self	Called	Function	Location
99.99	0.00	3	gsl_integration_qk	libgsl.so.0.16.0
99.99	0.00	48	ROOT::Math::GSLFunctionAdapter<ROOT::Math::...	libMathMore.so.5
99.99	0.00	48	ROOT::Math::WrappedFunction<genie::utils::gsl::...	libGCrossSections
99.99	0.00	48	genie::utils::gsl::wrap::dXSec_dElep_AR::DoEval(...)	libGCrossSections
99.99	0.00	48	ROOT::Math::GSLMCIntegrator::Integral(double ...)	libMathMore.so.5
99.99	0.00	48	gsl_monte_miser_integrate	libgsl.so.0.16.0
99.98	0.00	477 651	ROOT::Math::GSLMonteFunctionAdapter<ROOT...	libMathMore.so.5
99.98	0.00	477 651	genie::utils::gsl::wrap::d3Xsec_dOmegaldThetap...	libGCrossSections
99.89	0.00	477 370	genie::AlvarezRusoCOHXSec::XSec(genie::Intera...	libGAlvarezRuso-2
99.83	0.00	477 370	alvarezruso::Coh::DXSec(double, double, double...	libGAlvarezRuso-2
99.12	35.00	477 370	alvarezruso::Coh::NuclearCurrent(TLorentzVect...	libGAlvarezRuso-2
89.58	0.00	646	gsl_monte_miser_integrate'2	libgsl.so.0.16.0
33.20	2.94	1 527 582 405	alvarezruso::Coh::DeltaCouplingInMed(TLorentz...	libGAlvarezRuso-2
12.20	2.85	7 412 010 985	TLorentzVector::TLorentzVector(TLorentzVecto...	libPhysics.so.5.34
10.83	10.83	29 801 607 225	__muldc3	libgcc_s.so.1
10.37	0.97	564 516 003	alvarezruso::Coh::DeltaWidthPauliBlocked(TLor...	libGAlvarezRuso-2
10.24	2.39	7 984 165 486	TLorentzVector::~~TLorentzVector()	libPhysics.so.5.34
9.56	9.56	2 305 111 343	0x000000000000173d0	libm-2.11.3.so
9.18	1.54	564 516 002	alvarezruso::Coh::DeltaSelfEnergyIm(double)	libGAlvarezRuso-2
7.43	7.43	16 548 144 859	TObject::~~TObject()	libCore.so.5.34
7.13	3.47	8 557 752 073	TVector3::TVector3(TVector3 const&)	libPhysics.so.5.34
6.83	6.83	15 970 240 428	TObject::TObject(TObject const&)	libCore.so.5.34
4.57	0.73	8 563 481 090	TVector3::~~TVector3()	libPhysics.so.5.34
3.22	0.79	566 425 482	alvarezruso::Coh::PiDecayVertex(TLorentzVecto...	libGAlvarezRuso-2
3.09	0.75	564 516 003	alvarezruso::Coh::DeltaWidthFree(TLorentzVect...	libGAlvarezRuso-2
3.05	0.56	764 175 202	cexp	libm-2.11.3.so
2.97	1.83	1 527 582 404	alvarezruso::Coh::NucleonPropagator(TLorentz...	libGAlvarezRuso-2
2.19	0.95	1 527 582 405	csqrt	libm-2.11.3.so
1.78	1.78	3 085 884 808	__divdc3	libgcc_s.so.1
1.76	1.76	4 583 515 214	__fpclassify	libm-2.11.3.so
1.59	1.59	769 951 226	0x000000000000190c0	libm-2.11.3.so
1.23	1.18	1 129 032 005	alvarezruso::Coh::PionMomentumCM(TLorentzV...	libGAlvarezRuso-2
0.89	0.23	571 677 395	TLorentzVector::TLorentzVector(TVector3 const...	libPhysics.so.5.34
0.86	0.78	19 572 149	alvarezruso::integrationTools::RG202D(double, ...)	libGAlvarezRuso-2
0.63	0.00	48	alvarezruso::Coh::SolveWavefunctions()	libGAlvarezRuso-2
0.63	0.01	384 000	alvarezruso::EikonalSolution::Element(double, d...	libGAlvarezRuso-2
0.48	0.48	2 822 580 010	alvarezruso::Coh::DeltaSelfEnergyConstant(dou...	libGAlvarezRuso-2
0.42	0.42	9 931 947 102	alvarezruso::Constants::NucleonMass()	libGAlvarezRuso-2
0.37	0.01	7 684 800	alvarezruso::Nucl::DensitiesR(double, double&, ...)	libGAlvarezRuso-2
0.37	0.01	15 369 600	genie::utils::nuclear::Density(double, int, double)	libGUtils-2.8.0.so
0.36	0.02	15 369 600	genie::utils::nuclear::DensityGaus(double, doubl...	libGUtils-2.8.0.so
0.28	0.01	19 669 491	genie::Messenger::operator()(char const*)	libGMessenger-2.

Code Profile with cachegrind

← Hadronic current (35%)

← gcc multiplication (11%)

← exponentiation (10%)

} Assorted ROOT overhead (18%)

} Other math functions (5%)

Why is the hadronic current taking so much time?

Pros & Cons

Of lookup table approach:

- PRO:**
- ▶ (Relatively) fast
 - ▶ (Relatively) accurate
- CON:**
- ▶ Lookup-table needed for each neutrino/target combination
 - ▶ Limited by cut-off in E_{π}

Pros & Cons

Of full calculation:

- PRO:**
- ▶ Exact at each point in phase space
 - ▶ Valid over any neutrino energy range
- CON:**
- ▶ Slow to generate events and
 - ▶ Glacially slow to generate splines (not necessarily a problem if this is not done often)
 - ▶ Valid over any neutrino range

Summary & Questions

- ▶ Alvarez-Ruso microscopic model has been implemented in GENIE
- ▶ and even used in Dan's analysis
- ▶ Full calculation at each cross section evaluation is slow which could inhibit spline calculation
- ▶ We are in the process of trying to optimise it.
- ▶ Alternate lookup-table is in place. This is faster, but less accurate

Questions

- ▶ Do we want to keep both calculation methods? Which should be used by default?
- ▶ Need to finish optimising the full calculation before we can tell if it is really feasible
- ▶ How do we deal with the high energy behaviour of the model? The pion wavefunction starts oscillating rapidly at high E_π – the integration of the hadronic current will need to be done carefully
- ▶ Do we try to merge AR model and Rein-Seghal. If so how, and over what energy range?