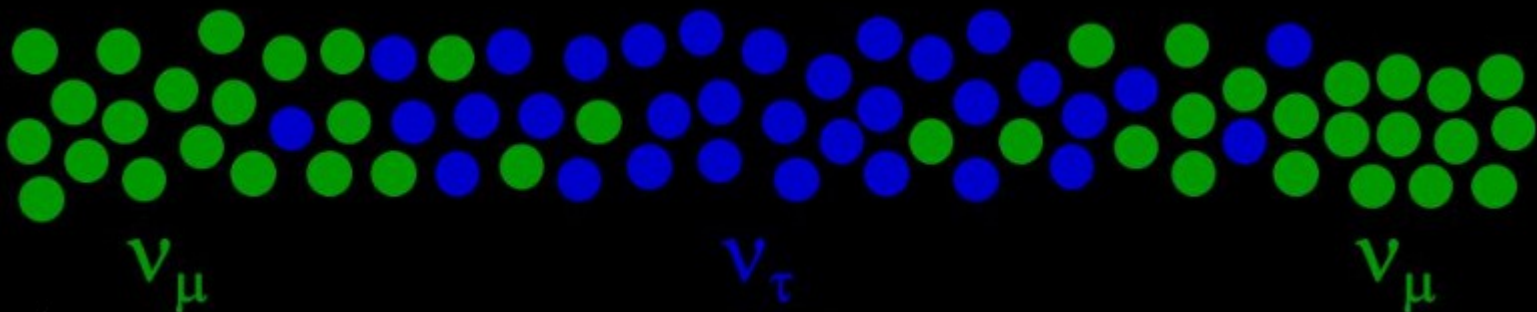


Steve Boyd

Neutrino Oscillations and the Case of the Missing Antimatter

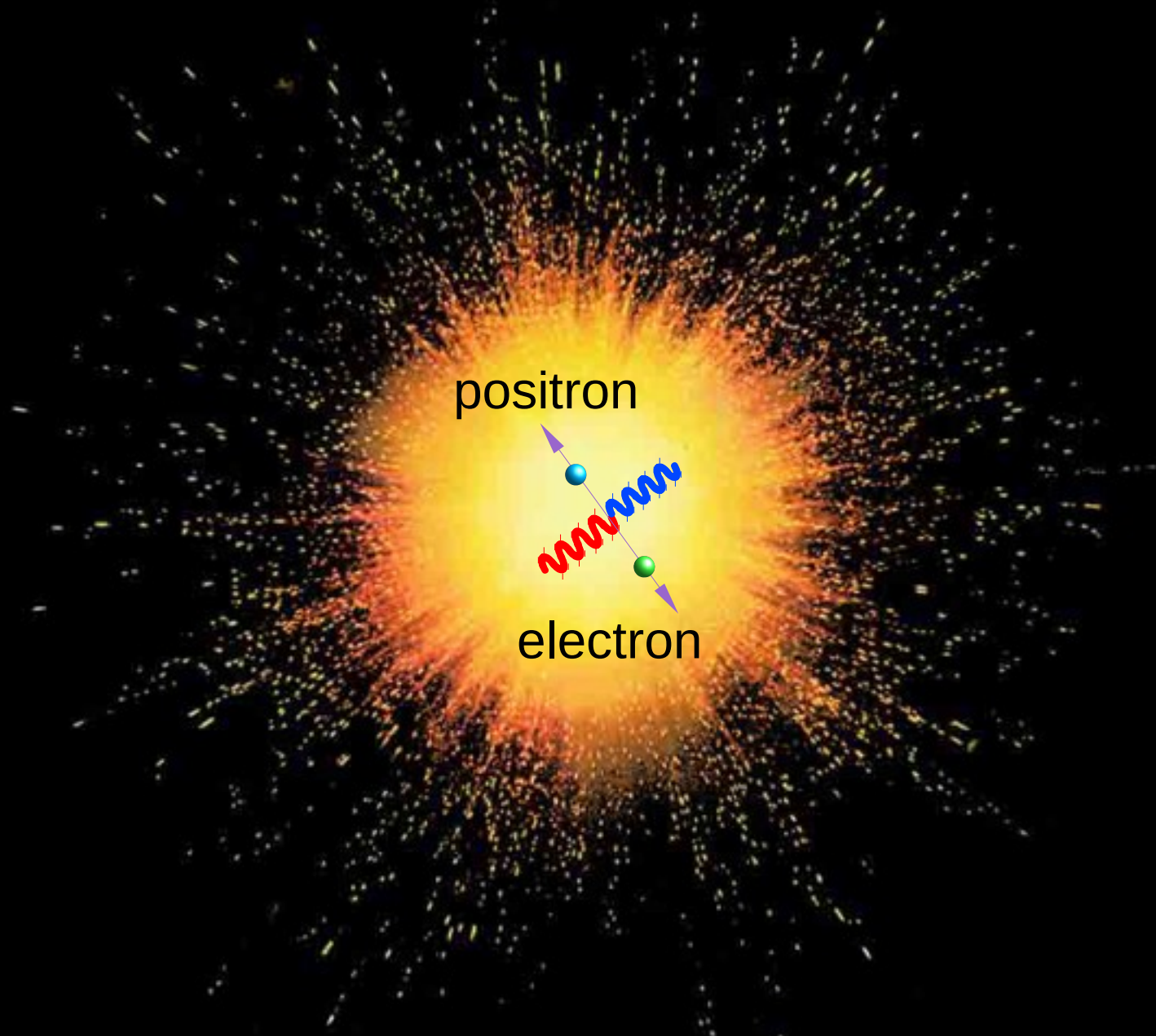




A vast field of stars, likely a star cluster or galaxy core, showing a mix of colors including blue, orange, and white. The stars are densely packed, with a concentration in the center. The background is a deep black, making the individual points of light stand out sharply.

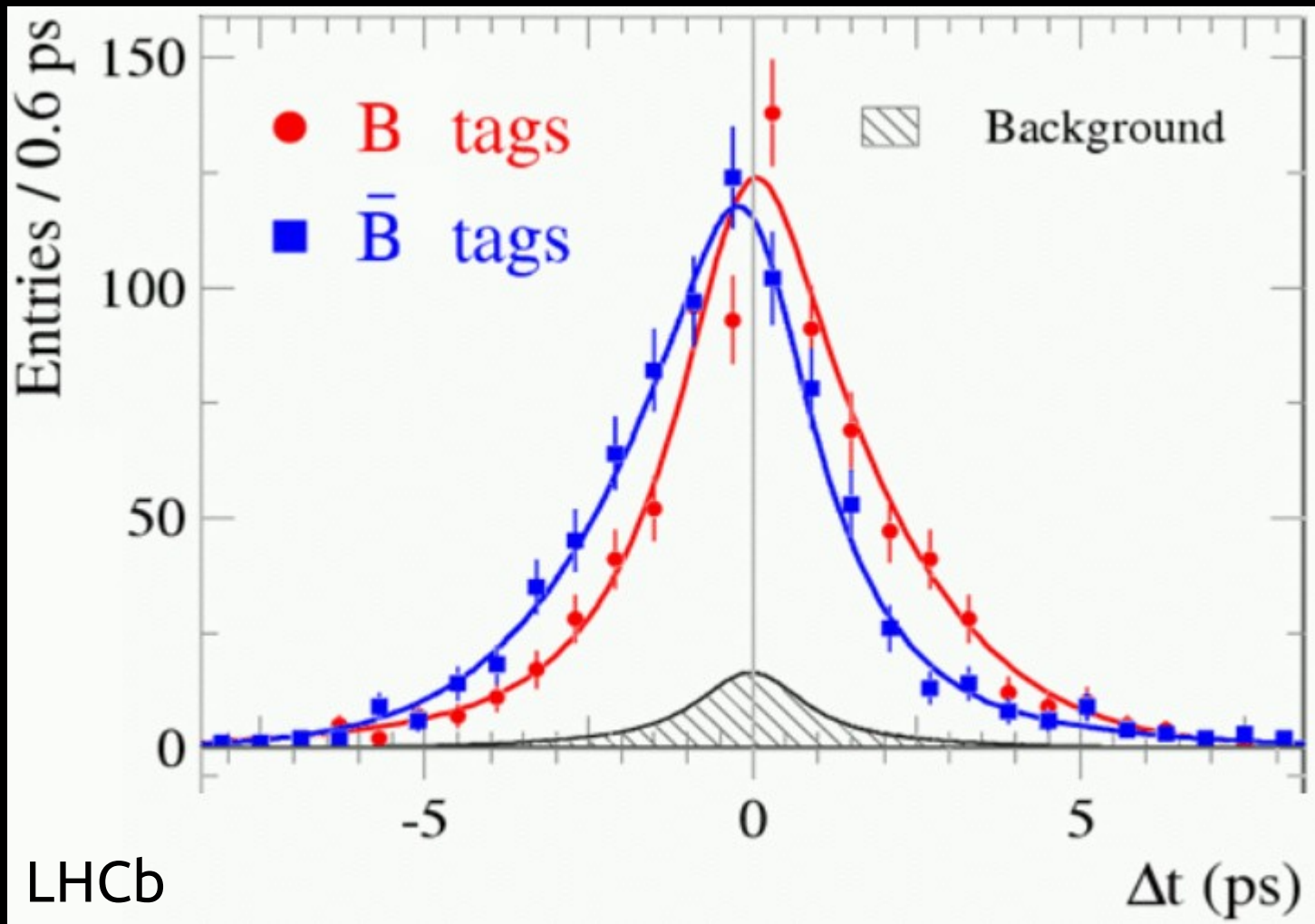
Where has all the antimatter gone?

Matter:Anti-matter was 1:1 in the early universe. Somehow most of the matter vanished with all of the antimatter.



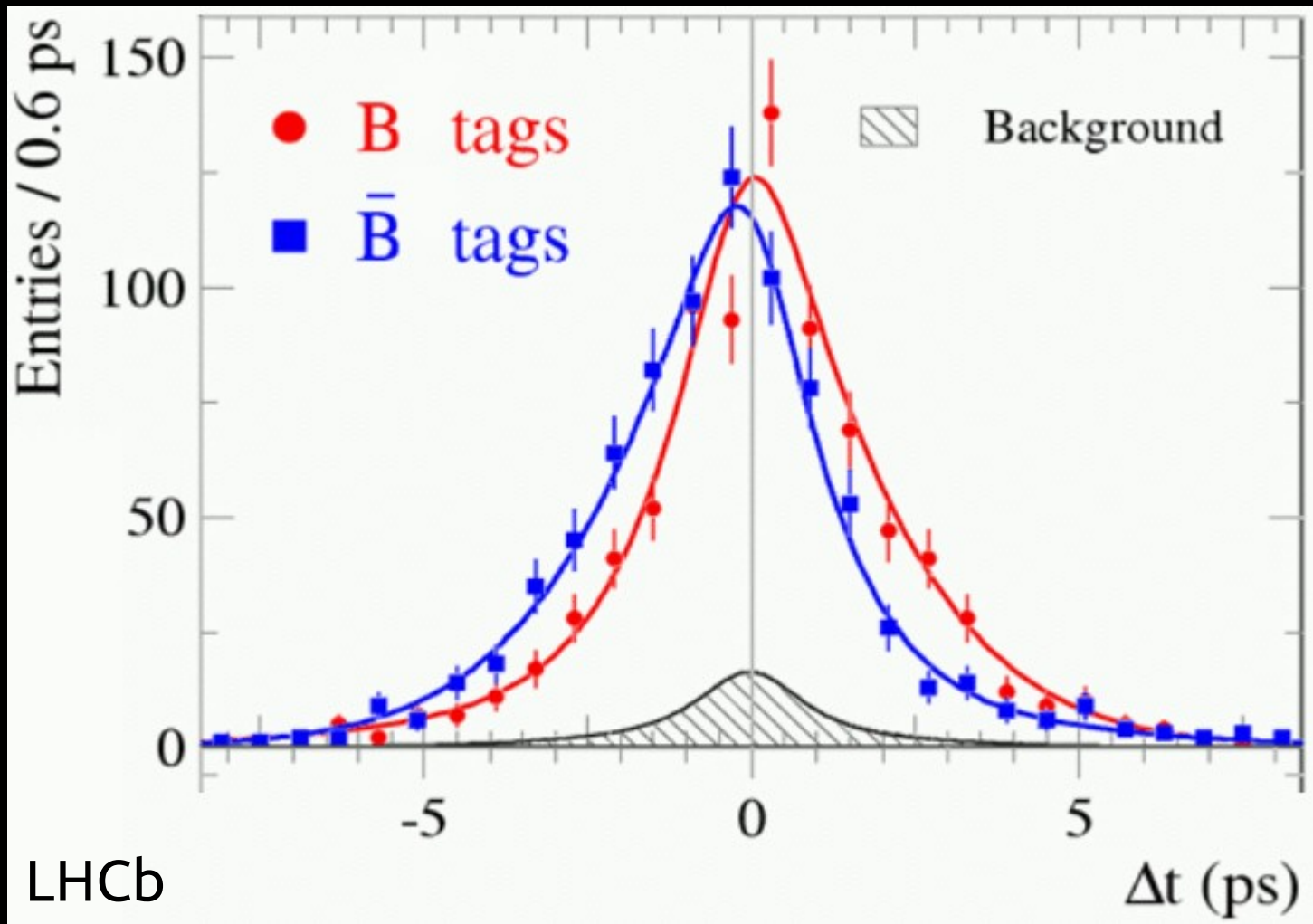
Is there a difference in the physics of matter and antimatter?

Matter / Antimatter Asymmetry



Quarks and anti-quarks behave differently \rightarrow CP violation
But it is a subtle effect (a few percent) – and not enough to explain the observed asymmetry in the universe.

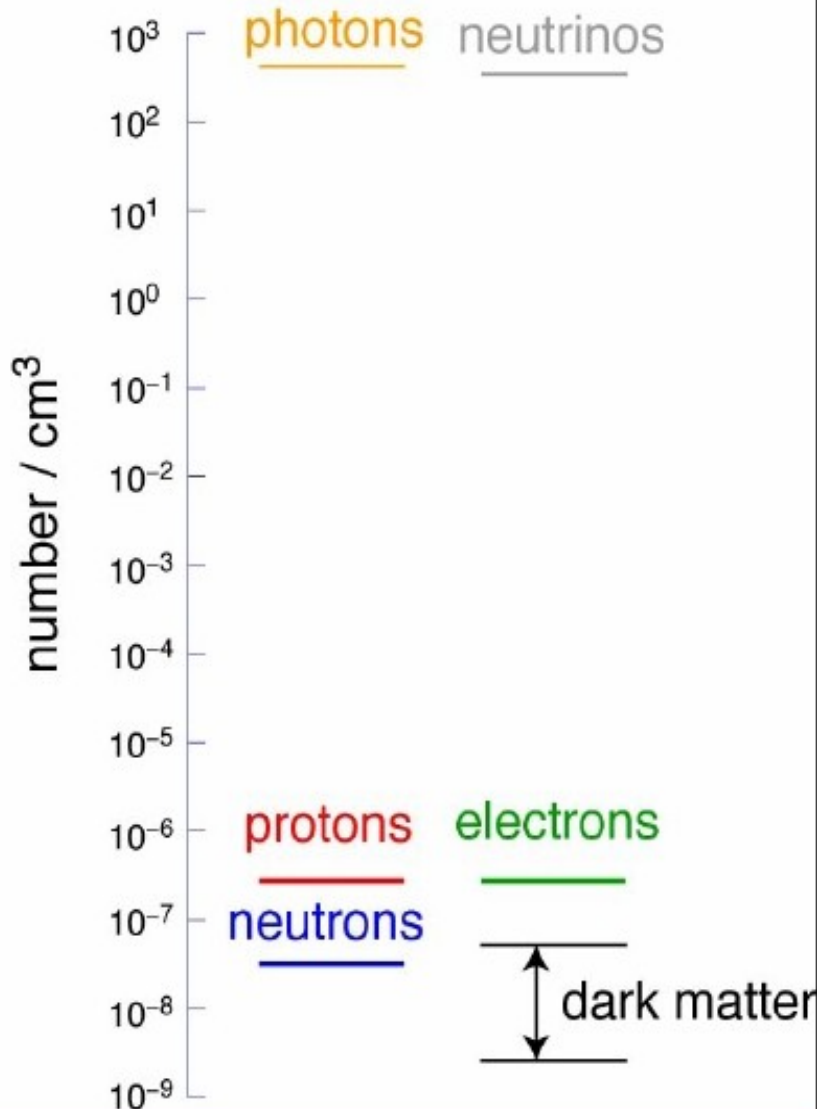
Matter / Antimatter Asymmetry



Part of the answer may lie, not with quarks, but with neutrinos

What is a neutrino?

The Particle Universe



- Spin 1/2, electrically neutral partner to a charged lepton
- Three **flavours** : ν_e, ν_μ, ν_τ
- Produced and interacts only through the weak interaction
- Almost massless
- Most common fermion in the universe
- Distinguished by very small interaction probabilities

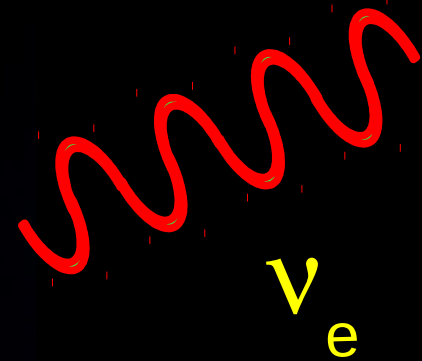
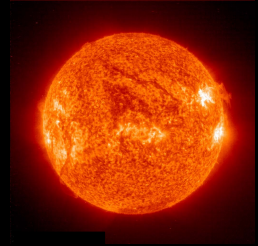
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"The chances of a neutrino actually hitting something as it travels through all this howling emptiness are roughly comparable to that of dropping a ball bearing at random from a cruising 747 and hitting, say, an egg sandwich."

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Probability $\approx 5 \times 10^{-13}$

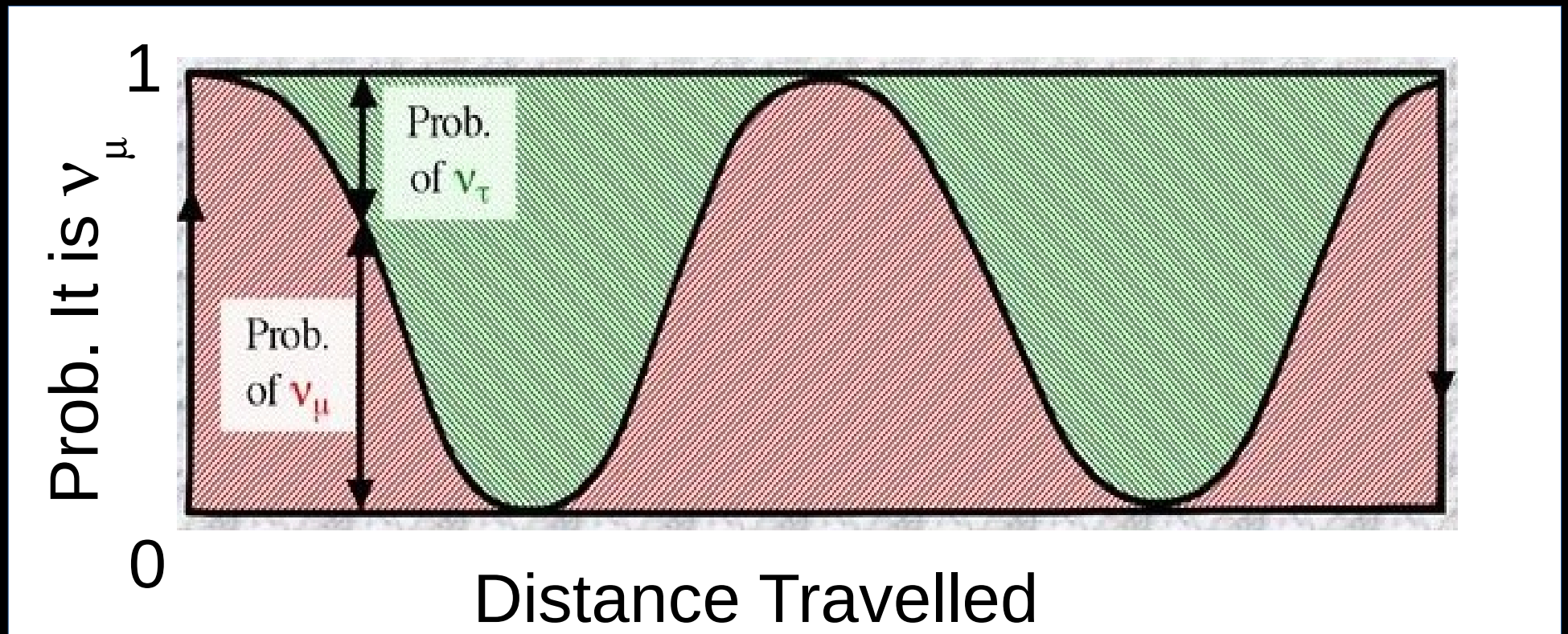


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Neutrino Flavour Oscillations

Neutrinos can change flavour as they propagate from point to point



$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2\left(2.54 \frac{L}{E} \Delta m^2\right)$$

$$\text{Mass splitting: } \Delta m^2 = m_1^2 - m_2^2$$

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They can do this because the thing we've always called the neutrino (flavour states with definite flavour) is actually made from 3 other things with definite mass (mass states).

We never know which mass state takes part in a given interaction. This uncertainty generates interference between the flavour states

Mixing parameter
we have to measure

$$\begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix}$$

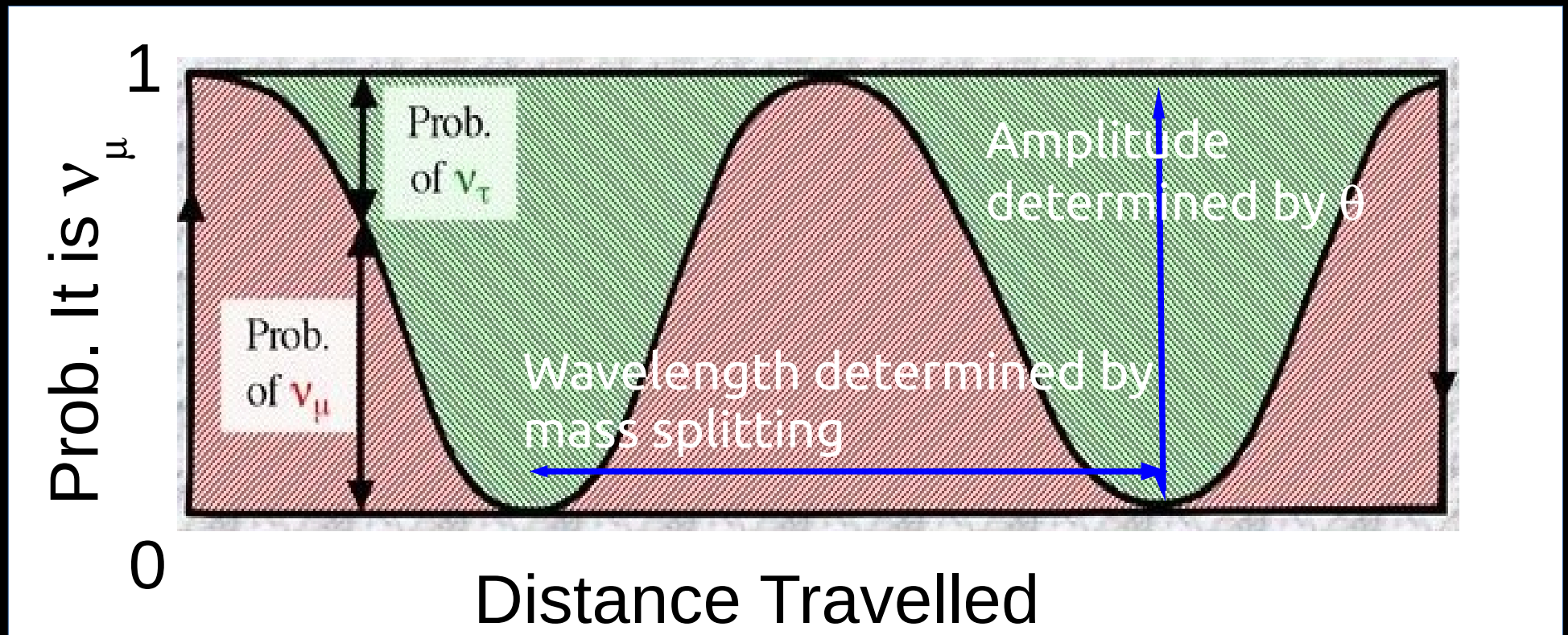
↑
Mass states

↑
Mixing matrix

↑
Flavour states

Neutrino Flavour Oscillations

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There are 3 neutrinos

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \Leftrightarrow U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix}$$

U is called the Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix
In 3-dimensions, U can have complex parameters

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$c_{ij} = \cos \theta_{ij} \quad s_{ij} = \sin \theta_{ij}$$

What we know...

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23 - sector

$$\mathbf{v}_\mu \rightarrow \mathbf{v}_\tau$$

$$\theta_{e\mu} = 45.0^\circ \pm 2.4^\circ$$

13-sector

$$\mathbf{v}_\mu \rightarrow \mathbf{v}_e$$

$$\theta_{13} = 9.7^\circ \pm 2.0^\circ$$

12-sector

$$\mathbf{v}_\mu \rightarrow \mathbf{v}_e$$

$$\theta_{e\mu} = 32.5^\circ \pm 2.4^\circ$$

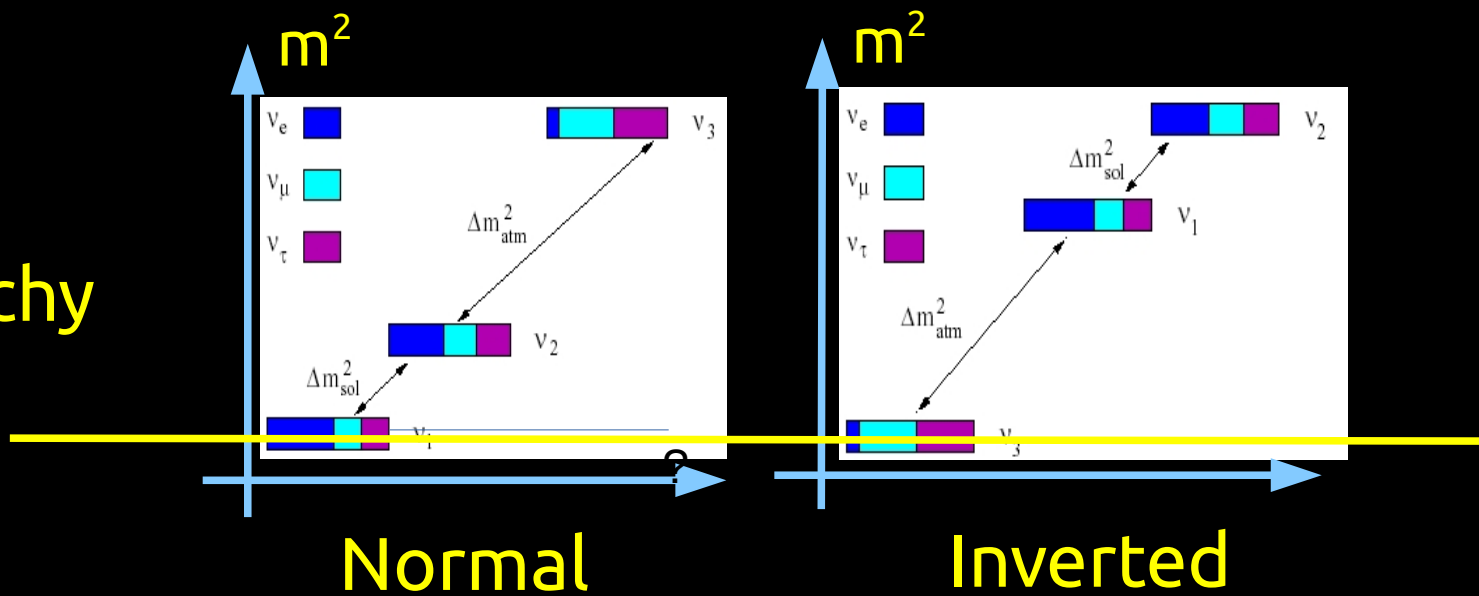
$$\Delta m_{12}^2 = 7.1 \times 10^{-5} eV^2 \quad \Delta m_{13}^2 = \Delta m_{23}^2 = |2.8 \times 10^{-3}| eV^2$$

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δ_{CP} controls the level of CP violation

Mass hierarchy



baseline mass ???

Why does it matter?

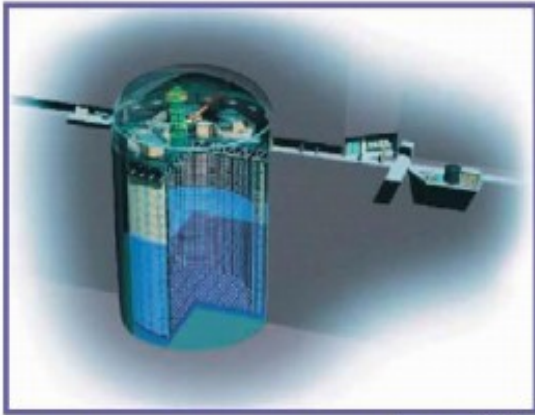
There is an idea floating about out there called Leptogenesis

It goes

1. Once upon a time in the very early universe there lived very heavy partners to our neutrino called (wait for it) “heavy neutrinos”.
2. CP violation in decays of the heavy neutrino were able to generate a lepton number asymmetry
3. Sphaleron transitions ($\bar{\nu}(\nu)$) conserve B-L, so if L is not conserved, neither is B
4. Observed baryon asymmetry is generated, at least in part, from CP violation in the leptons
5. Should look for CP violation in the neutrino sector.

T2K

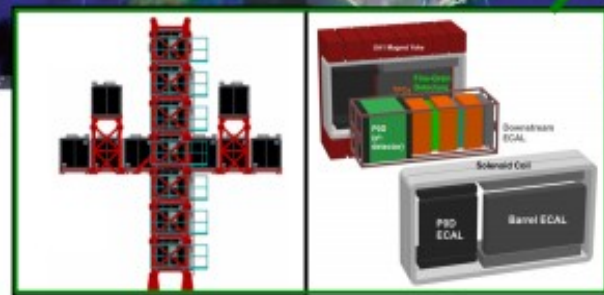
Measure flavour composition after oscillation here



Super-Kamiokande
(ICRR, Univ. Tokyo)

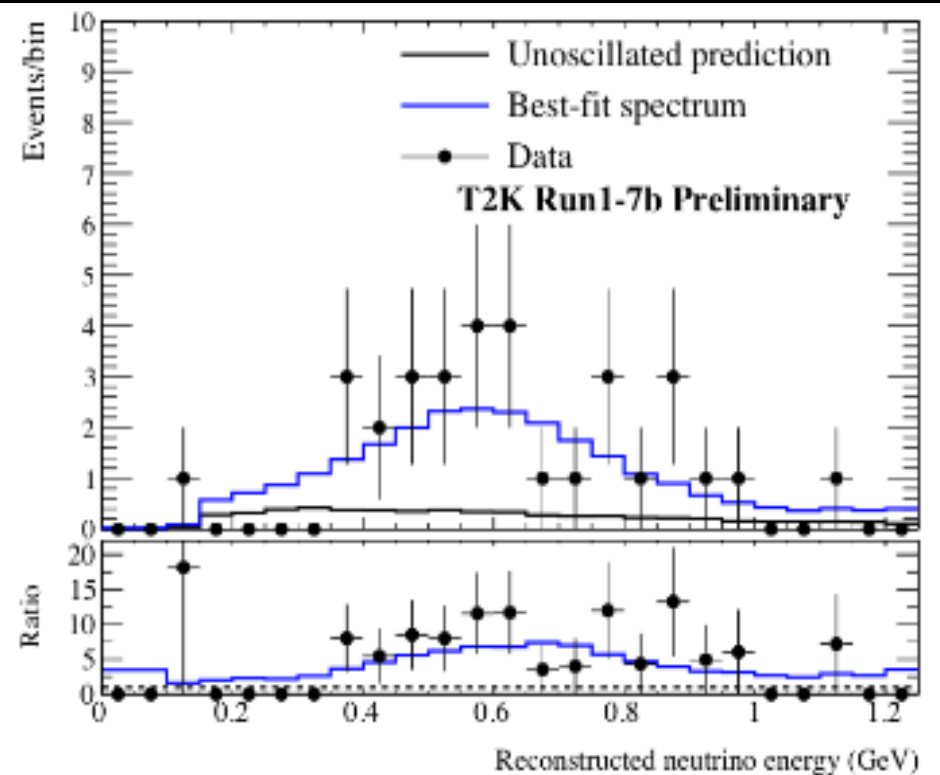
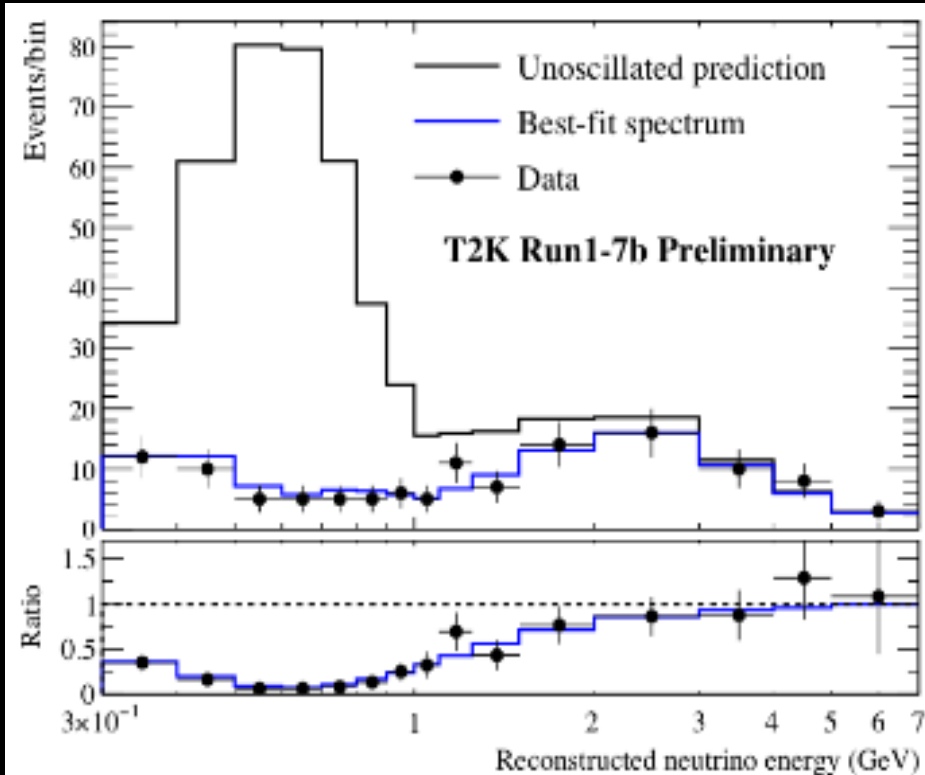


J-PARC Main Ring
(KEK-JAEA, Tokai)



Measure flavour composition before oscillation here

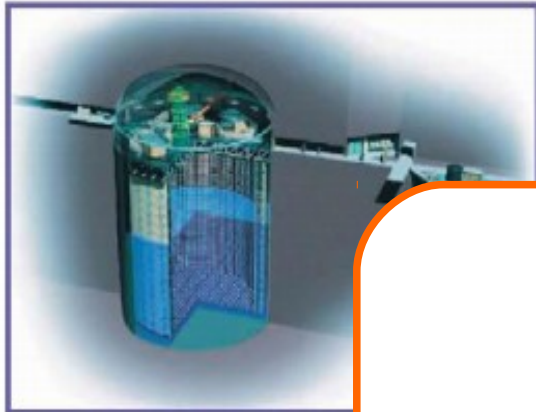
Long baseline experiments



ν_μ Disappearance Measurement

ν_e Appearance Measurement

T2K

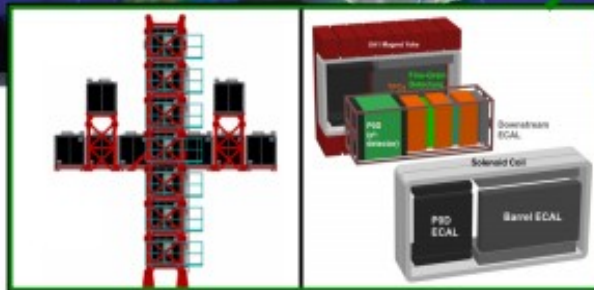


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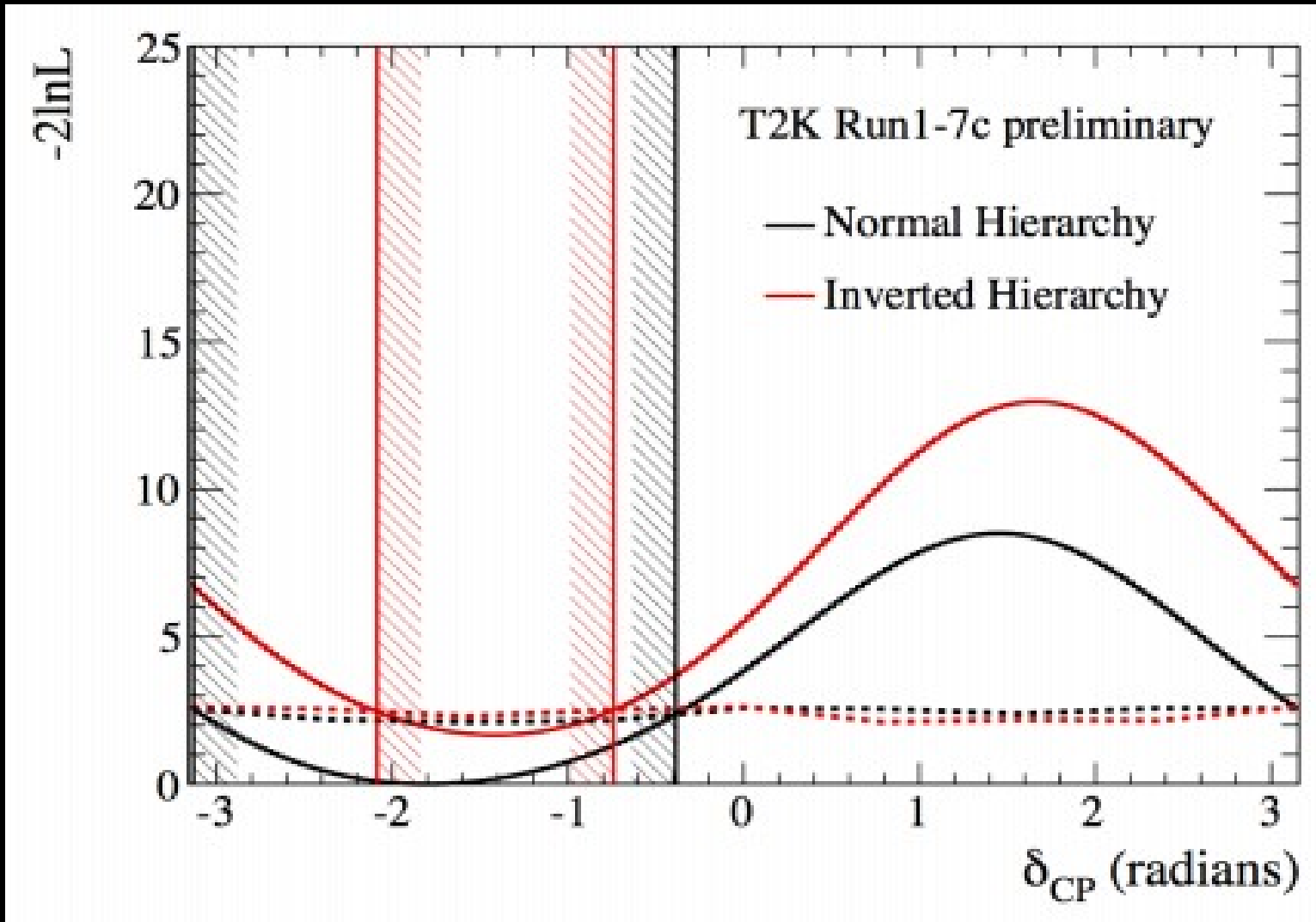


CP violation will show up if
$$P(\nu_{\mu} \rightarrow \nu_e) \neq P(\bar{\nu}_{\mu} \rightarrow \bar{\nu}_e)$$

JARC Main Ring
(JAEA, Tokai)

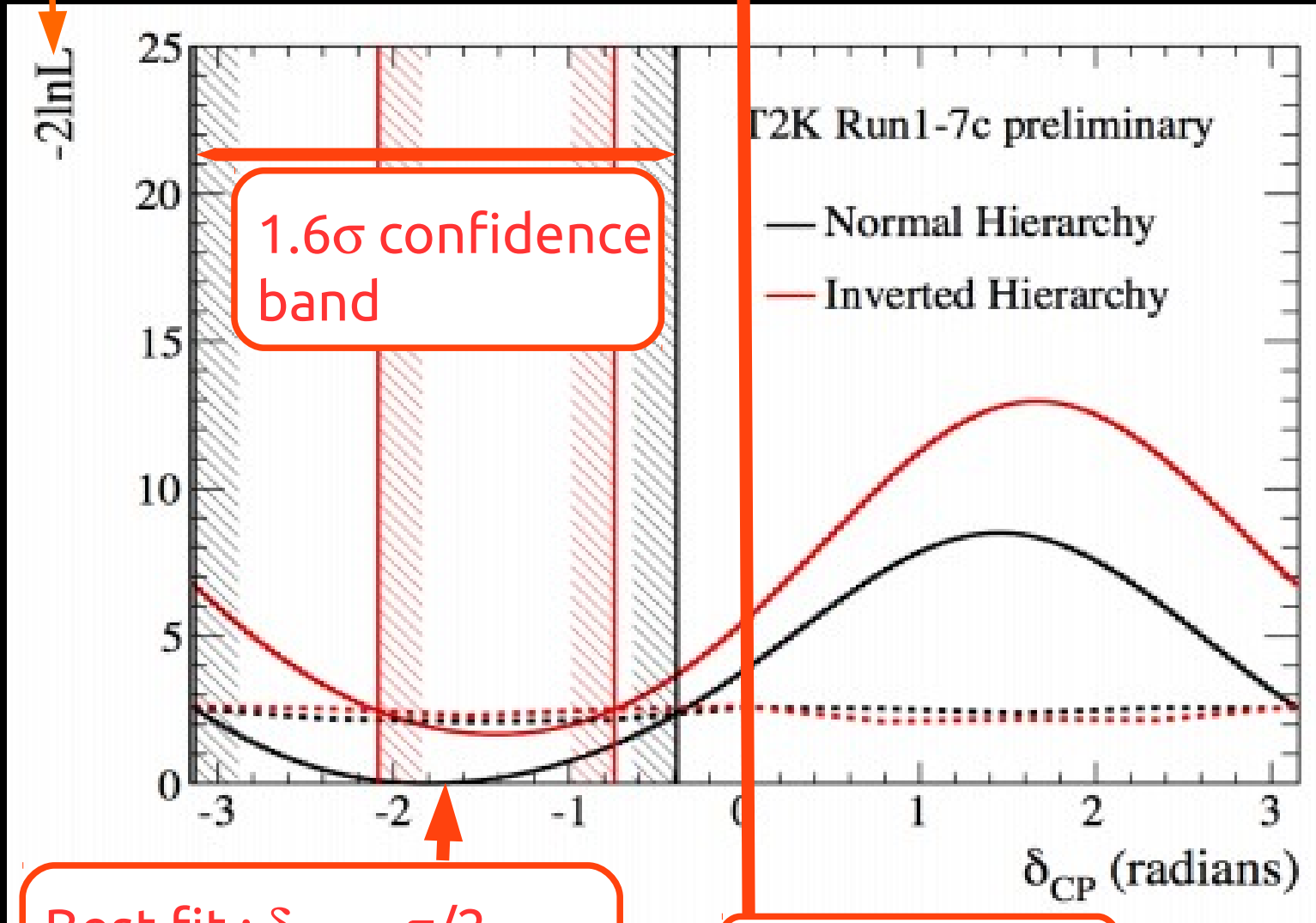


T2K Results



$$\chi^2 - \chi^2_{\min}$$

T2K Results

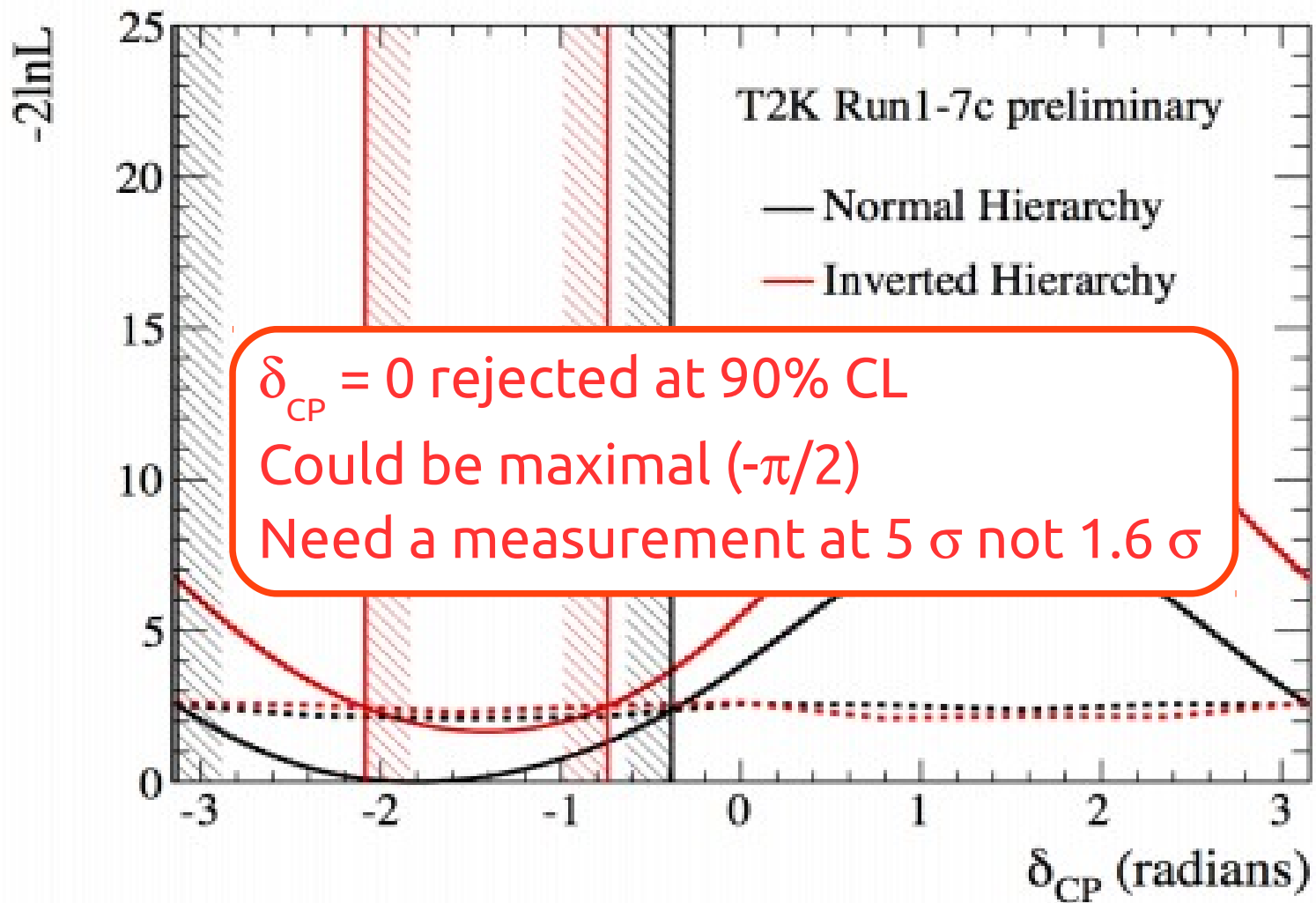


1.6 σ confidence band

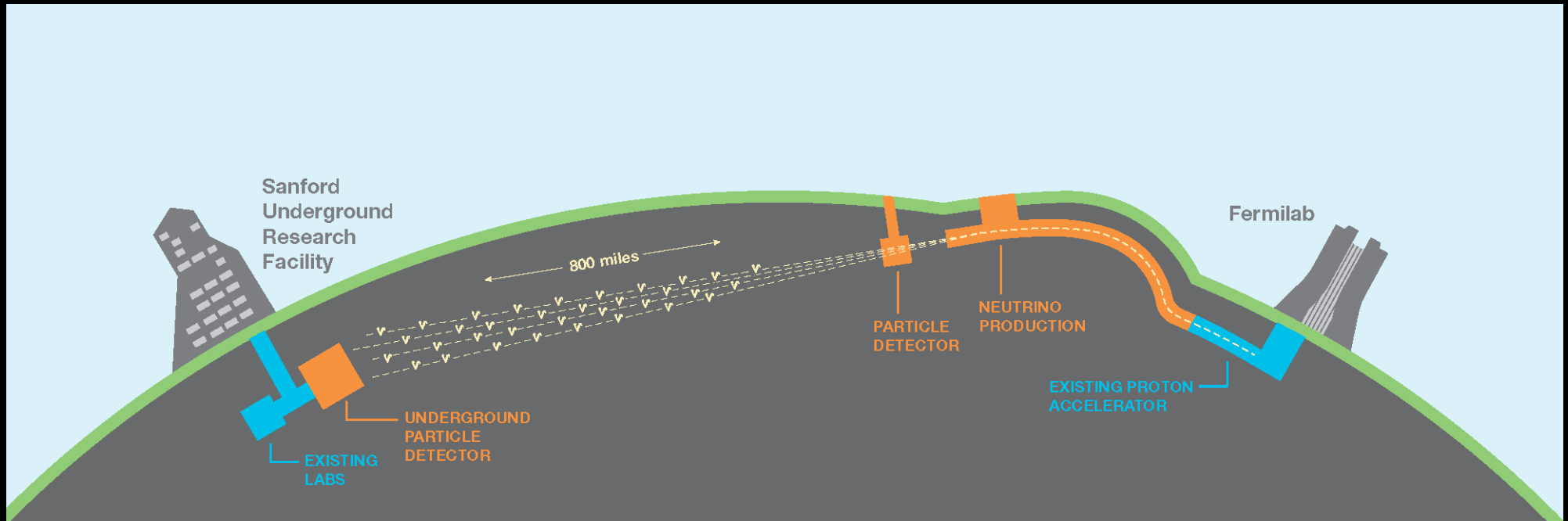
Best fit : $\delta_{CP} \sim -\pi/2$
with normal hierarchy

CP Conserving

T2K Results



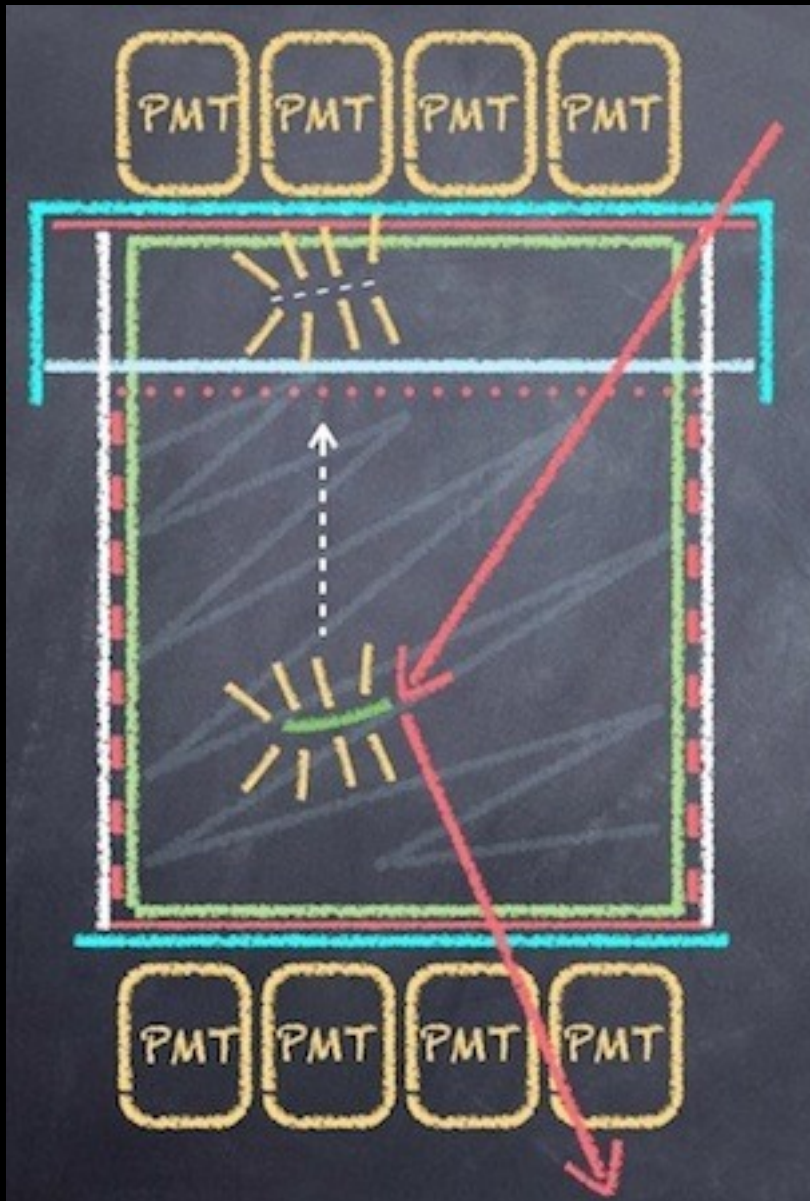
New experiments : DUNE and HyperK



DUNE (Deep Underground Neutrino Experiment)

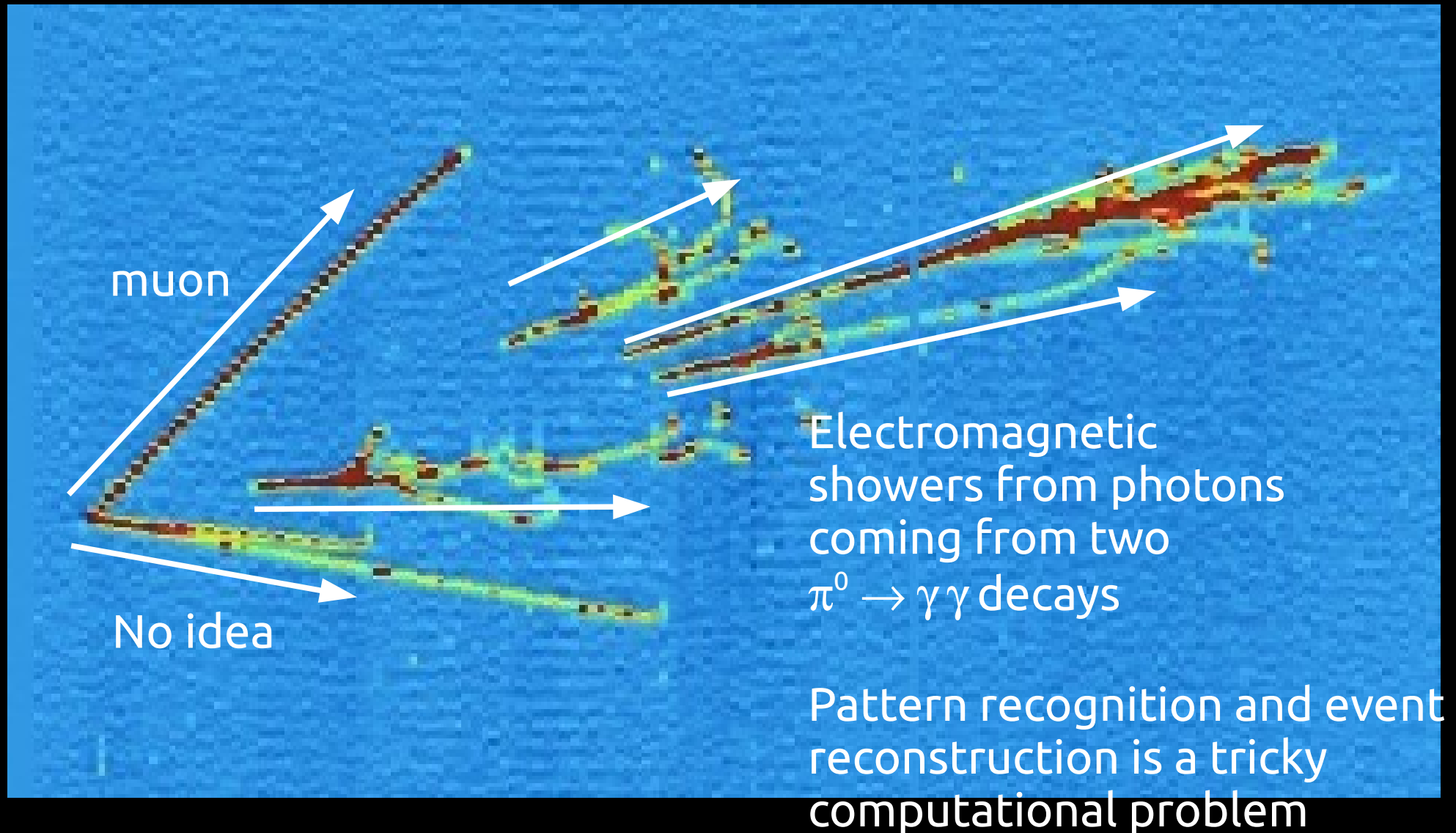
Baseline : 1300 km from Fermilab (Chicago) to Sudan (South Dakota)
Liquid Argon technology

Liquid Argon TPC

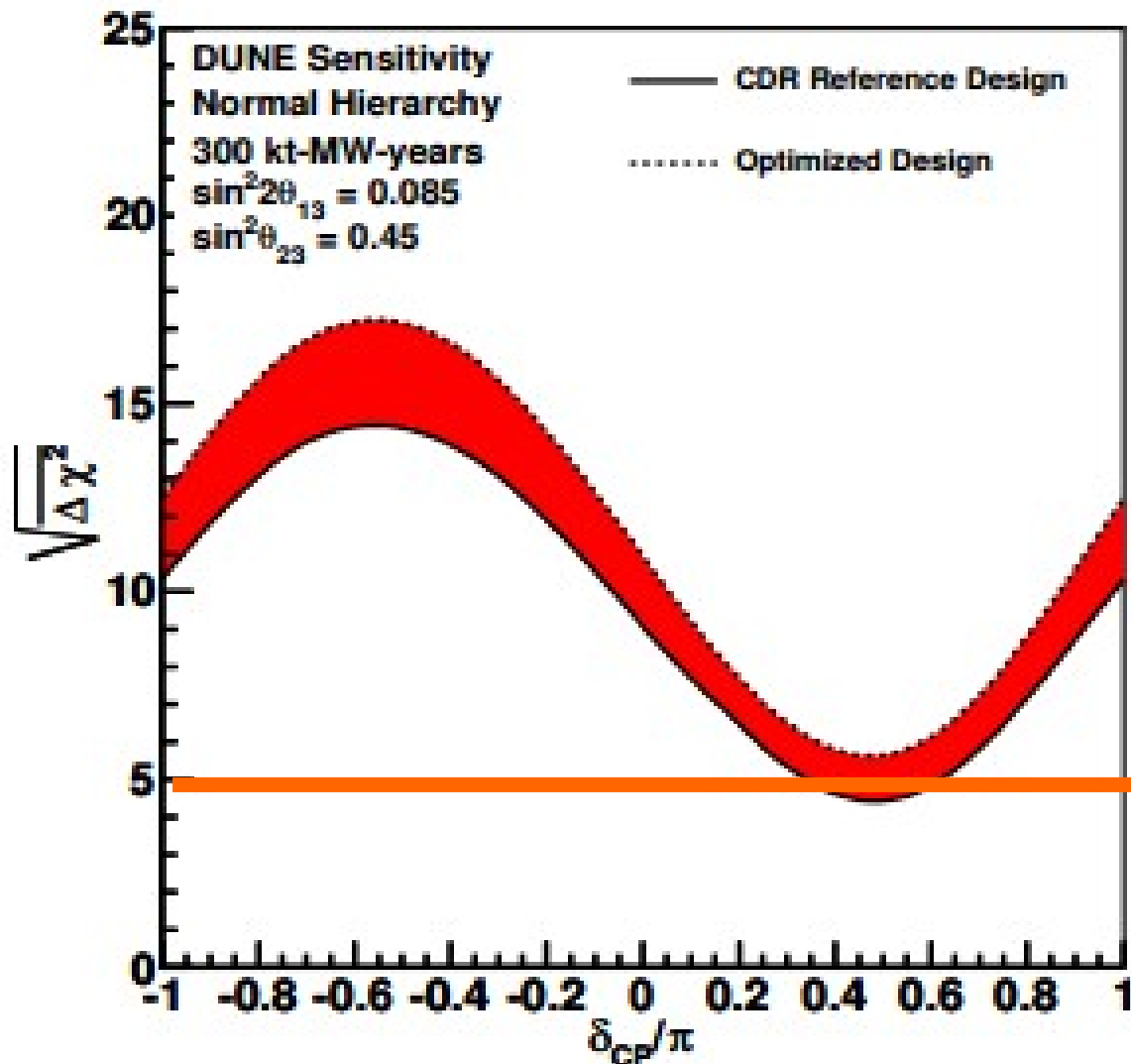


- TPC technology is not new
- TPCs usually use gas to generate ionisation electrons
- Neutrino detectors need to use liquid to get more events
- LAr is a new technology being trialled in various prototypes
- DUNE plans 4 TPC modules containing 17 kton of LAr
- Largest LAr TPCs ever built

Liquid Argon Time Projection Chamber



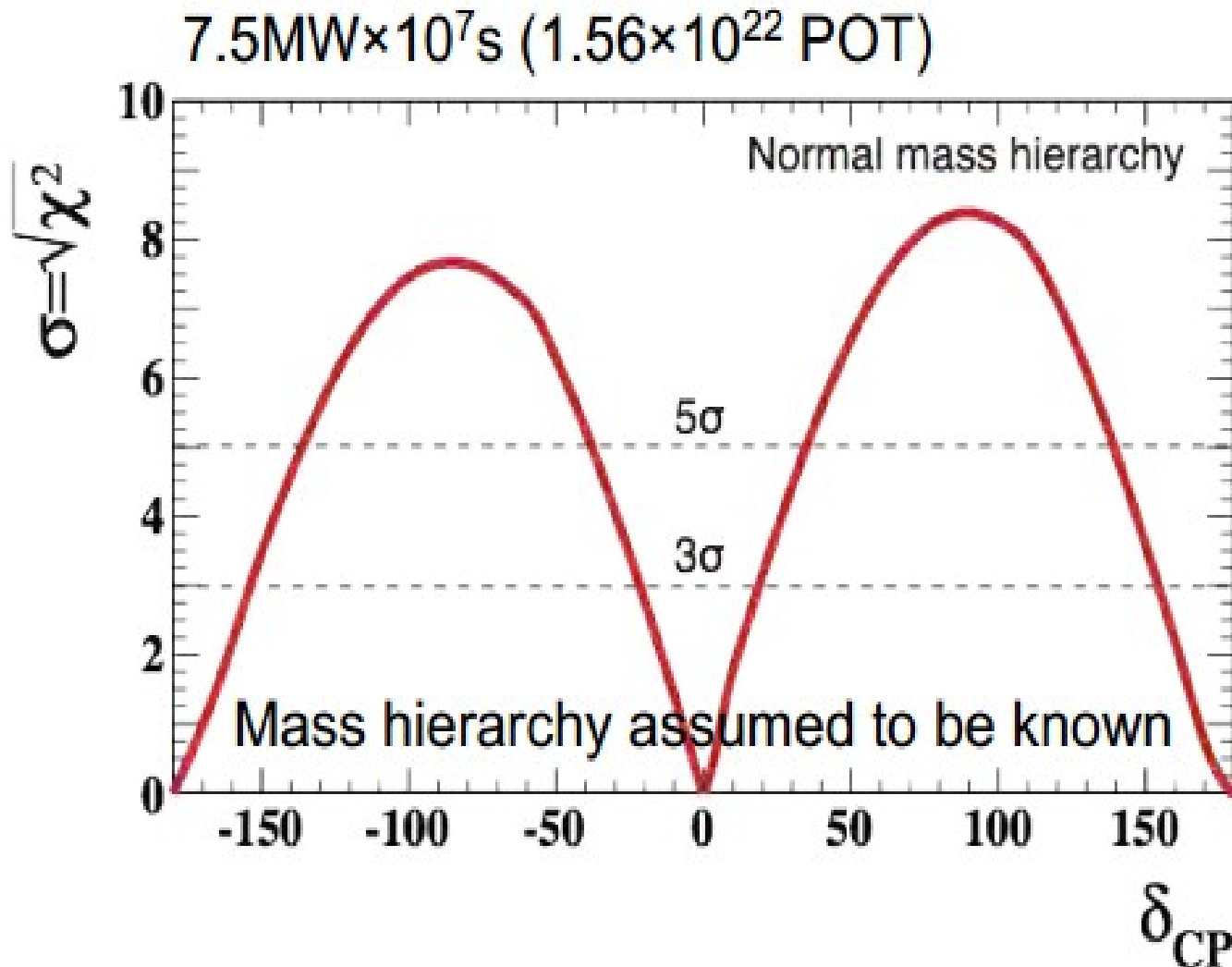
DUNE Predictions : Mass Hierarchy



■ After 4 years of operation

5 σ determination of mass hierarchy over entire range of δ_{CP}

DUNE Predictions : δ_{CP}



5 σ significance for
 $\sin \delta_{CP} \neq 0$ over
56% of δ_{CP} space

20% precision at
 $\delta_{CP} = -90$ degrees

Timeline

2018 : Technology prototyping programme at CERN

2019 : Delivery of Technical Design Report to funding agencies

2020 : Detector fabrication hubs ready and logistics program defined

2024 : Construction and installation of Far Detectors

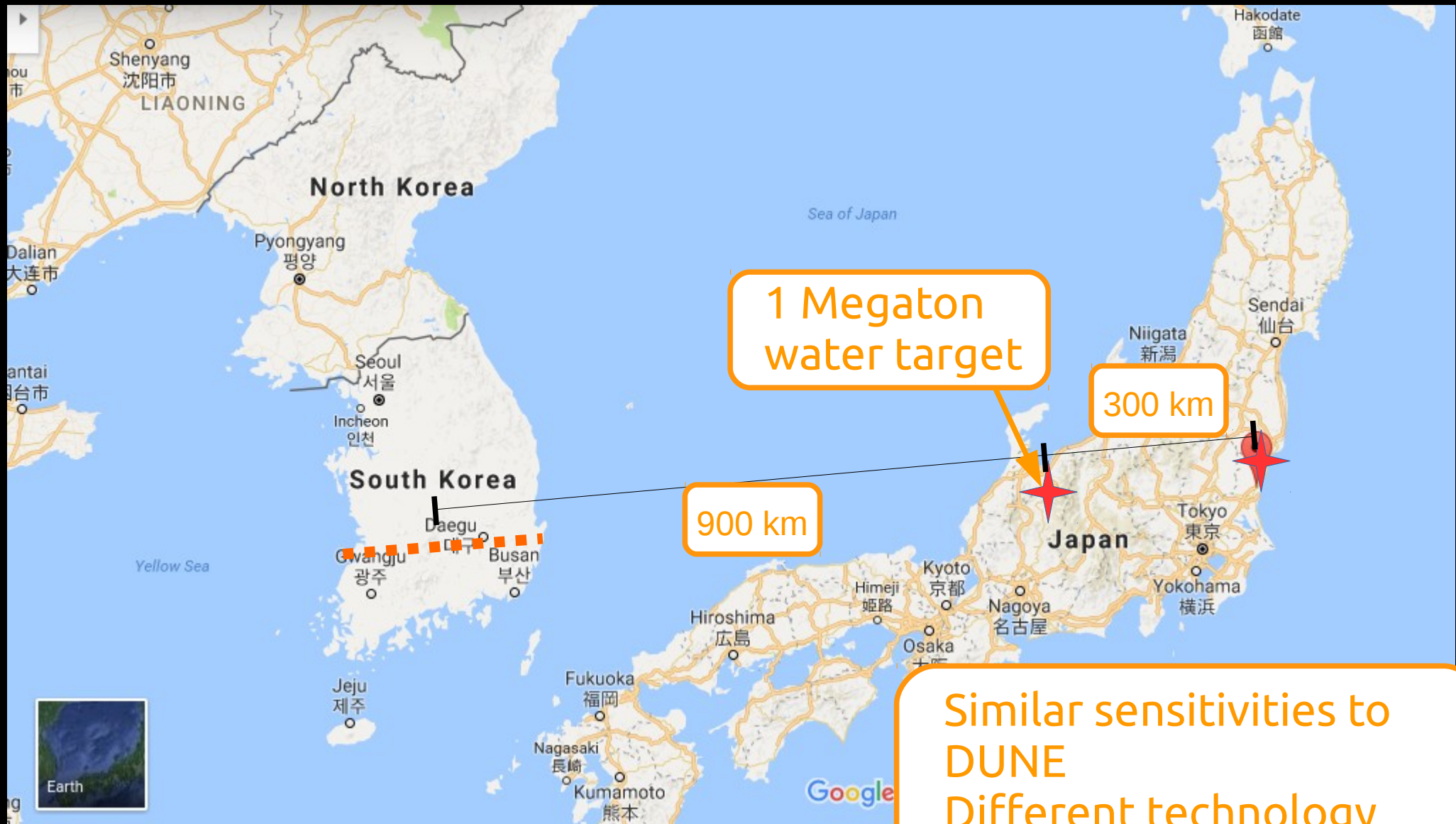
2025 : Commissioning of Far Detectors

2026 : Delivery of neutrino beam to Far site. Physics starts.

2028 : Preliminary results on mass hierarchy and δ_{CP}

2030 : Final results

Hyper-Kamiokande (The opposition)



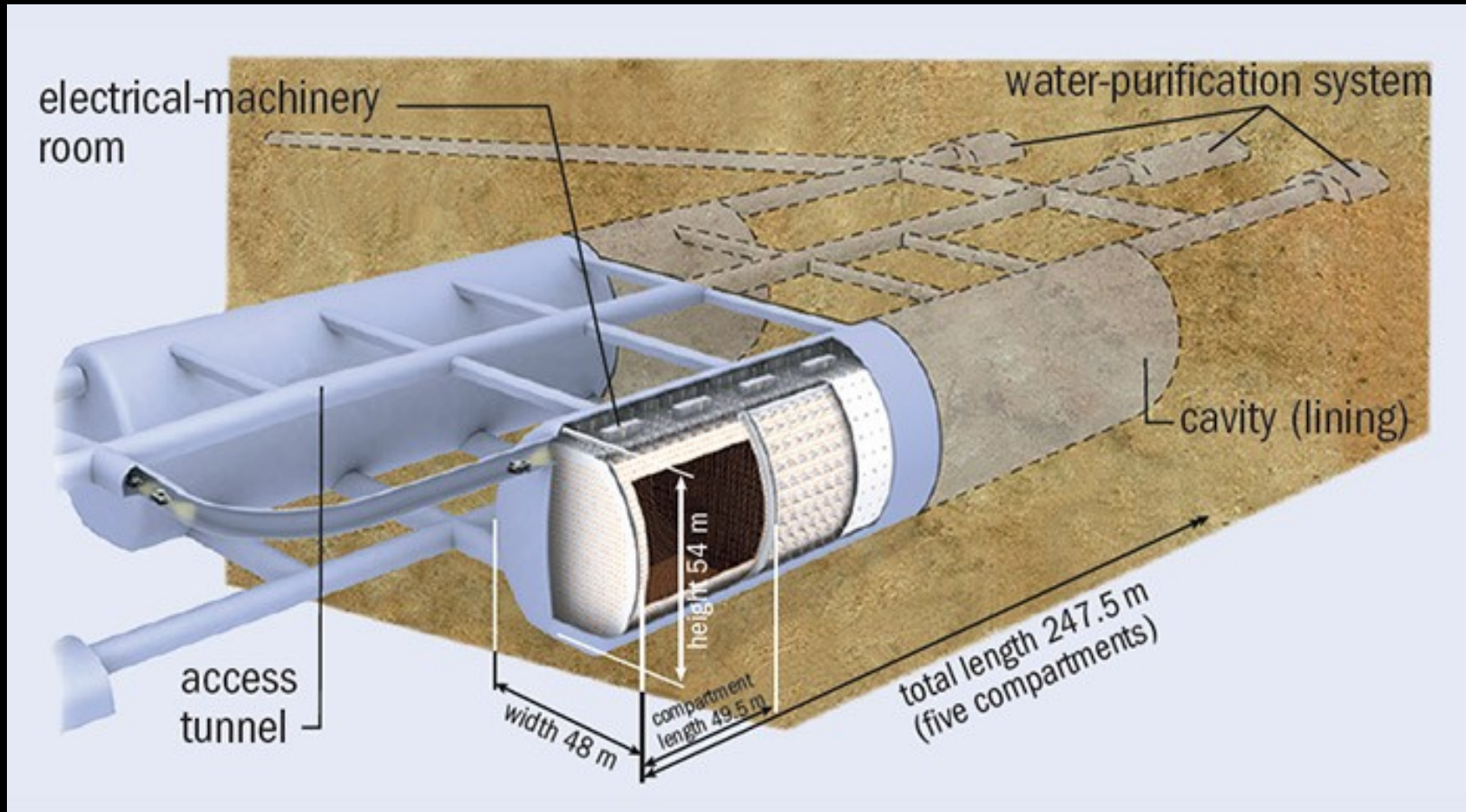
1 Megaton water target

300 km

900 km

Similar sensitivities to DUNE
Different technology
Timescale less advanced

Hyper-K Far Detector



Summary

Particle physics has studied CP violation in the quarks for years

We are only beginning to look at the leptons....and early indications are that the leptons behave significantly differently

We get all this information from long baseline experiments which are only beginning to reach the era of high-precision neutrino physics

Still 20 years of work to understand what is going on...

Assume 1 billion people eat an egg sandwich every 3 months

1.67×10^7 egg sandwiches/day

Let's say that 3 months of the year people can eat outside, and that they picnic once a month

140,000 external egg sandwiches/day

egg sandwich lifetime – 20 minutes

3000 egg sandwiches at any time

Area of egg sandwich – 15 cm x 15cm

62 m² total egg-sandwich area

Surface area of earth

500 million km²

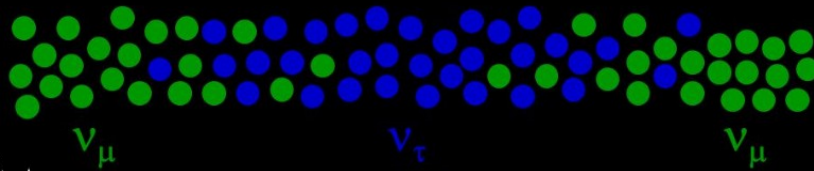
Suppose flight paths cover area of earth uniformly

Probability of egg-sandwich/
ball bearing
intersection

1×10^{-13}

Steve Boyd

Neutrino Oscillations and the Case of the Missing Antimatter

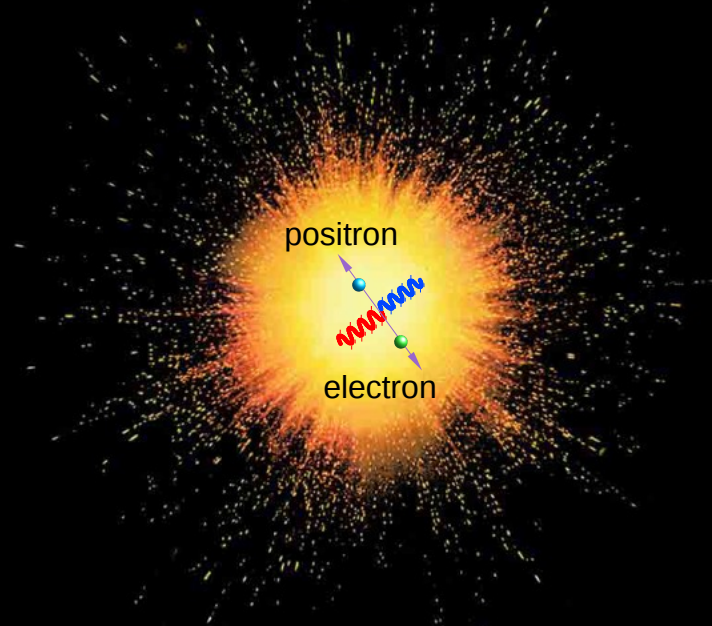






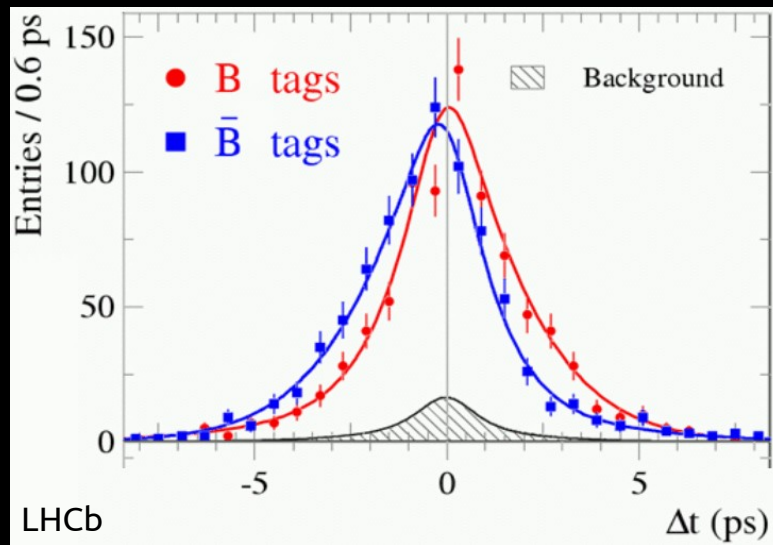
Where has all the antimatter gone?

Matter:Anti-matter was 1:1 in the early universe. Somehow most of the matter vanished with all of the antimatter.



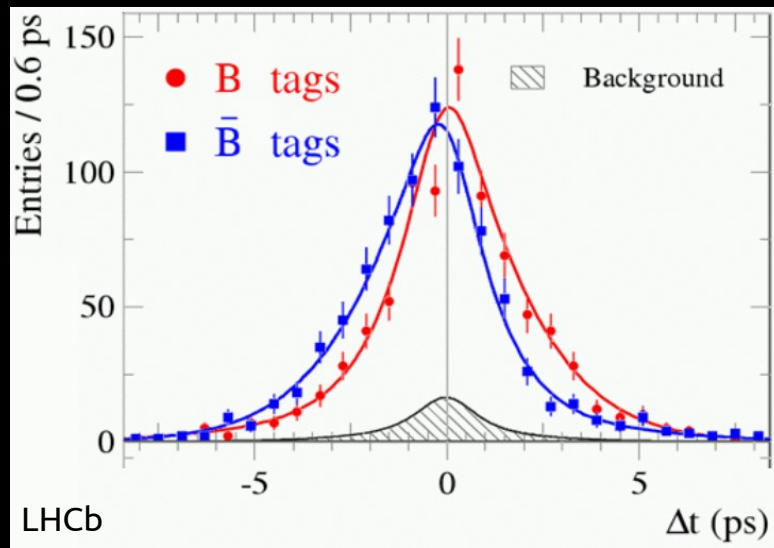
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Matter / Antimatter Asymmetry



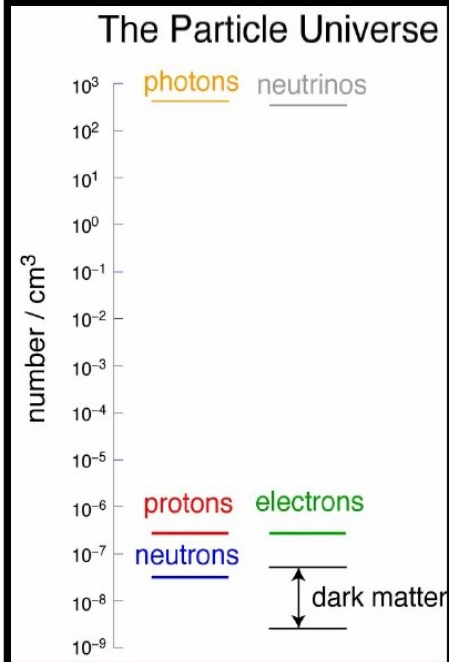
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Part of the answer may lie, not with quarks, but with neutrinos

What is a neutrino?



- Spin 1/2, electrically neutral partner to a charged lepton
- Three **flavours** : ν_e, ν_μ, ν_τ
- Produced and interacts only through the weak interaction
- Almost massless
- Most common fermion in the universe
- Distinguished by very small interaction probabilities

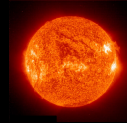
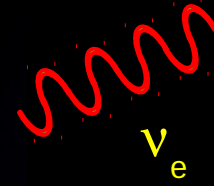
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Hans Bethe

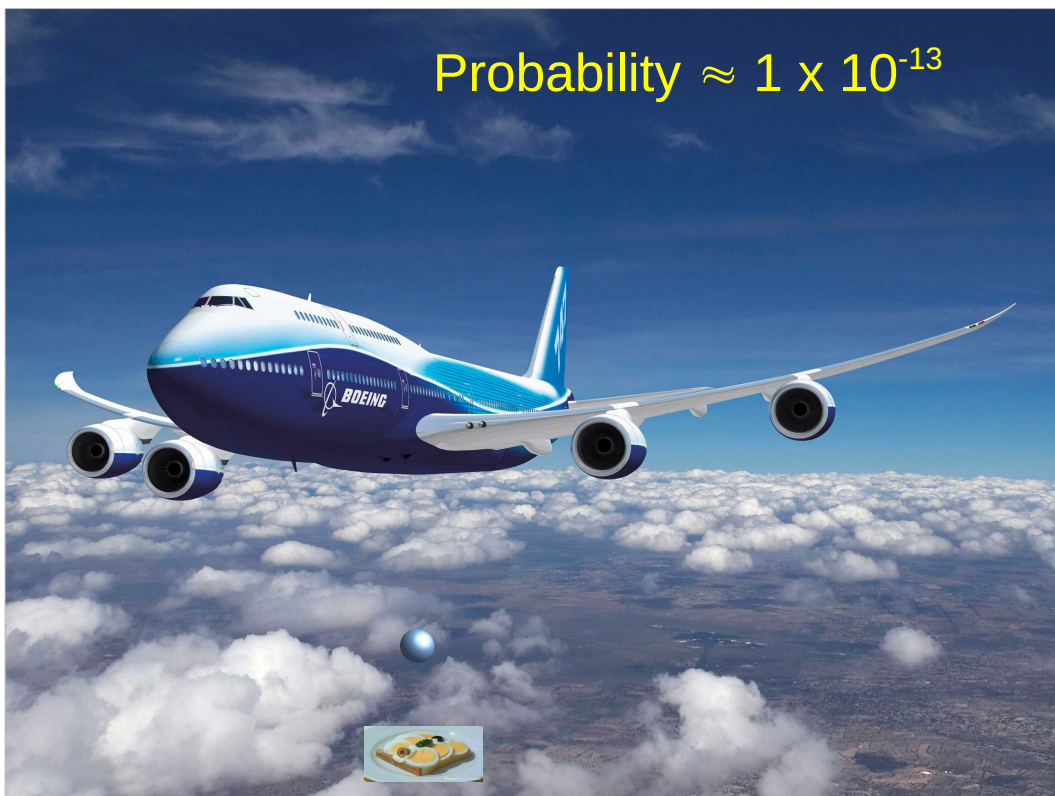
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Douglas Adams

Probability $\approx 5 \times 10^{-13}$

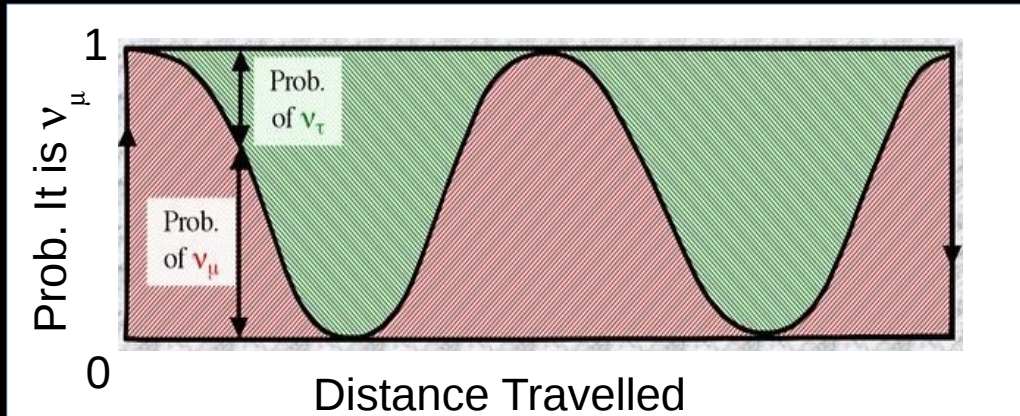


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Neutrinos can change flavour as they propagate from point to point



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Mass splitting: $\Delta m^2 = m_1^2 - m_2^2$

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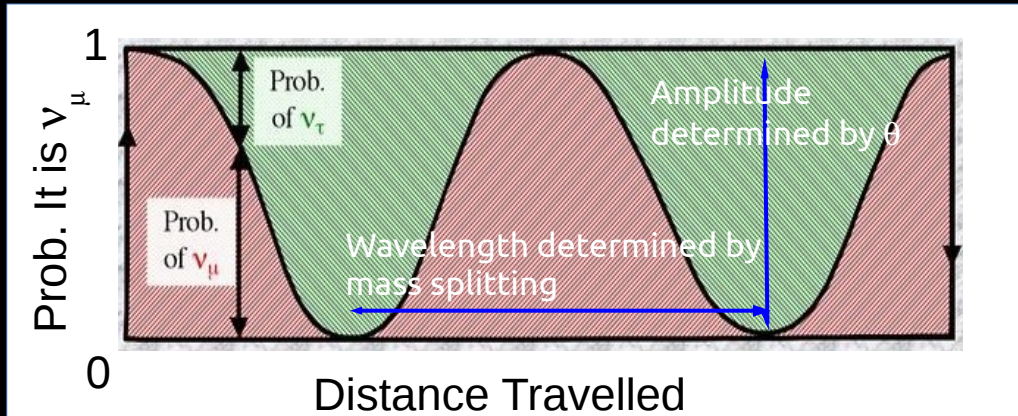
↑
Mass states

↑
Mixing matrix

↑
Flavour states

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There are 3 neutrinos

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$$c_{ij} = \cos \theta_{ij} \quad s_{ij} = \sin \theta_{ij}$$

What we know...

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13-sector

$$\mathbf{v}_\mu \rightarrow \mathbf{v}_e$$

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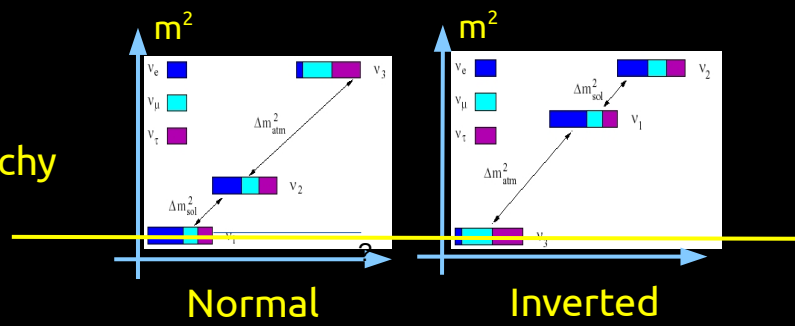
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...and what we don't

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δ_{CP} controls the level of CP violation

Mass
heirarchy



baseline
mass ???

Why does it matter?

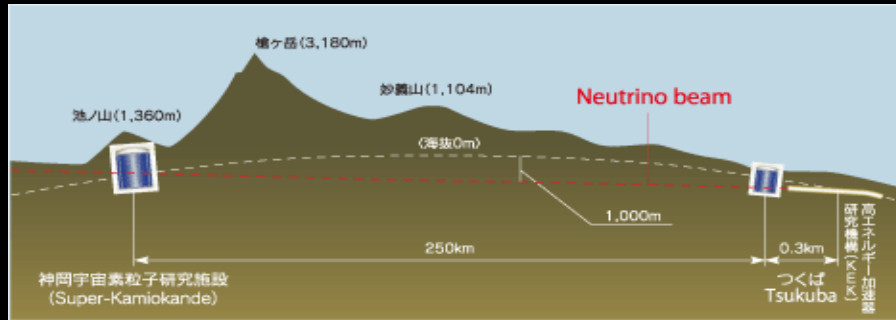
There is an idea floating about out there called Leptogenesis

It goes

1. Once upon a time in the very early universe there lived very heavy partners to our neutrino called (wait for it) "heavy neutrinos".
2. CP violation in decays of the heavy neutrino were able to generate a lepton number asymmetry
3. Sphaleron transitions ($\bar{\nu}_\mu \nu_\mu$) conserve B-L, so if L is not conserved, neither is B
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5. Should look for CP violation in the neutrino sector.

Long baseline experiments

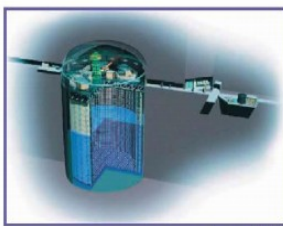
One way of studying these issues uses neutrino beams of known flavour content and measures their flavour mix after the neutrinos have travelled a long distance



Two types of measurements: *Disappearance*: $P(\nu_\mu \rightarrow \nu_\mu)$
Appearance: $P(\nu_\mu \rightarrow \nu_e)$

T2K

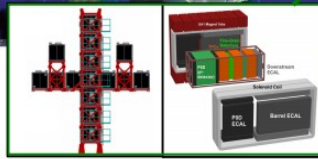
Measure flavour composition after oscillation here



Super-Kamiokande
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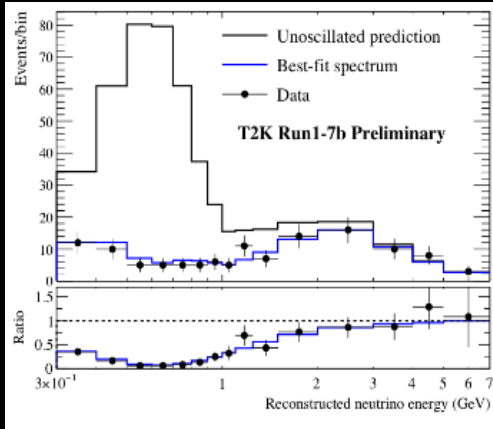


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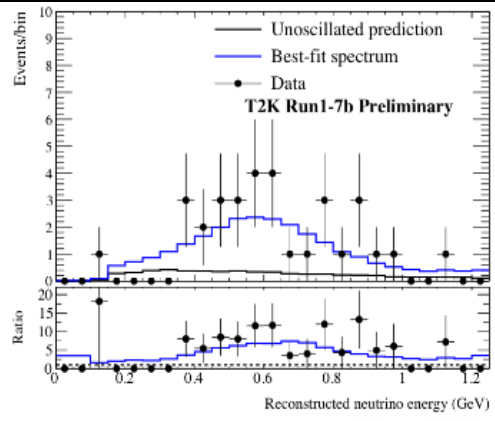


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Long baseline experiments

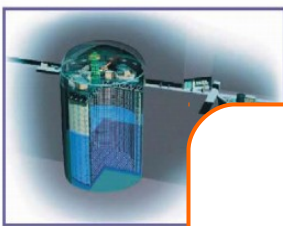


ν_{μ} Disappearance Measurement



ν_e Appearance Measurement

T2K

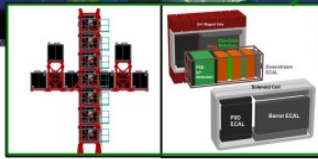


Super-Kamiokande
(ICRR, Univ. of Tokyo)

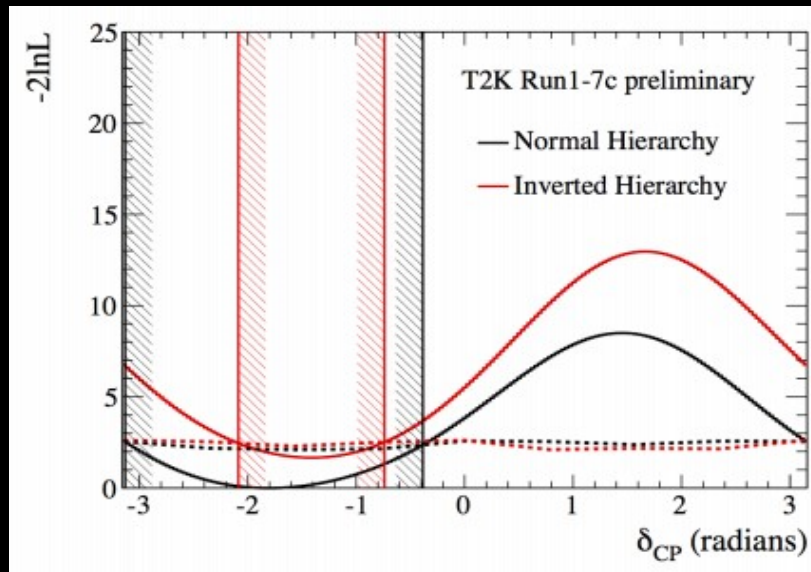


CP violation will show up if
 $P(\nu_\mu \rightarrow \nu_e) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$

J-PARC Main Ring
(JAEA, Tokai)

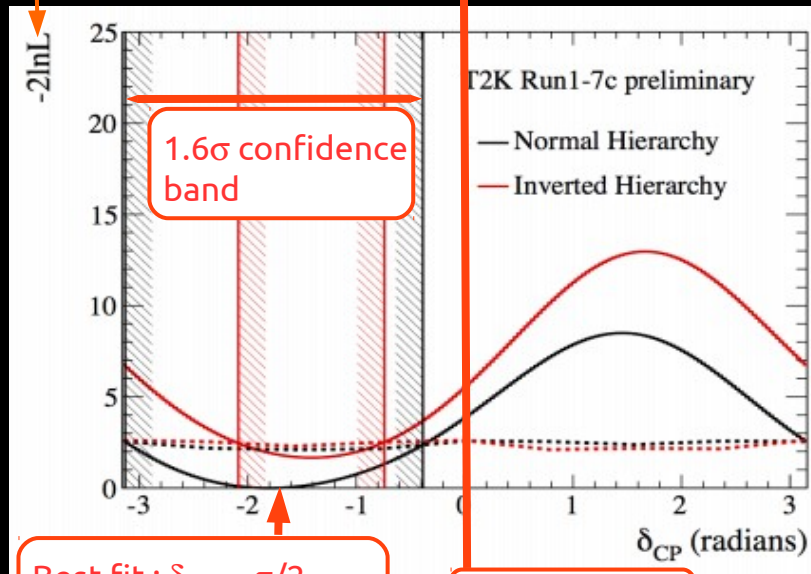


T2K Results



T2K Results

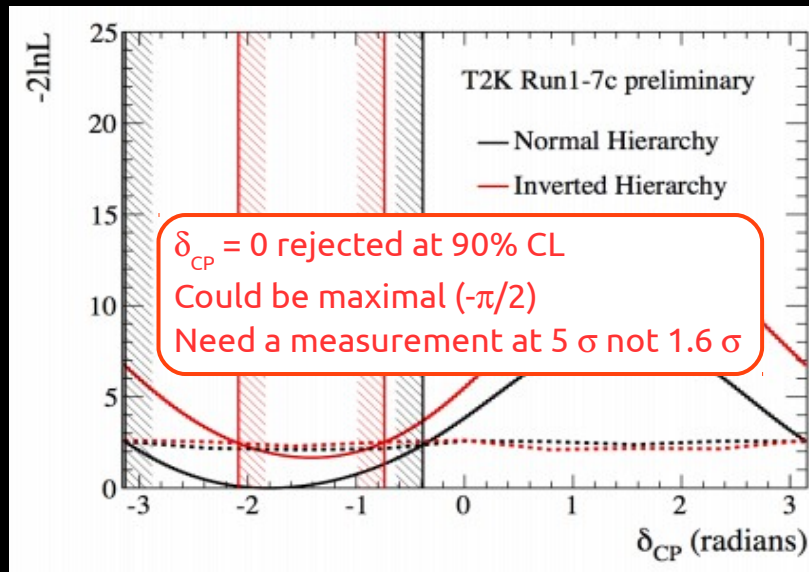
$$\chi^2 - \chi^2_{\min}$$



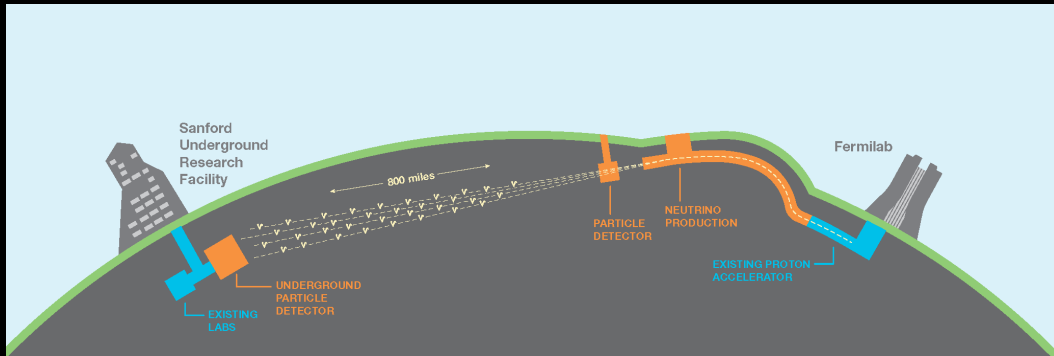
Best fit: $\delta_{CP} \sim -\pi/2$
with normal hierarchy

CP Conserving

T2K Results



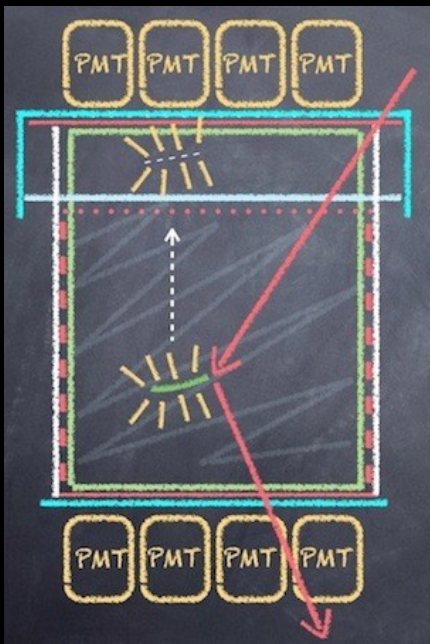
New experiments : DUNE and HyperK



DUNE (Deep Underground Neutrino Experiment)

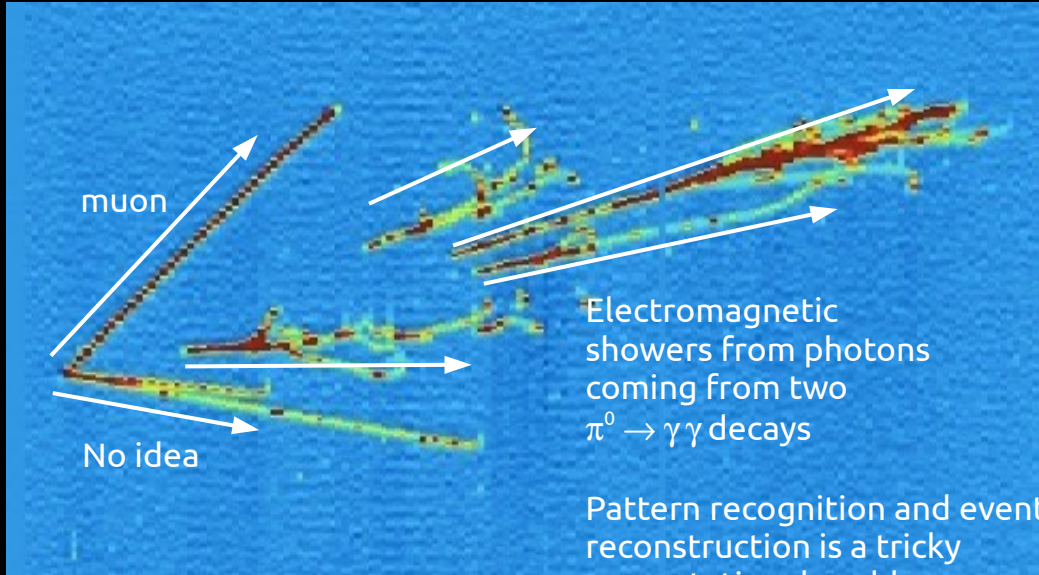
**Baseline : 1300 km from Fermilab (Chicago) to Soudan (South Dakota)
Liquid Argon technology**

Liquid Argon TPC

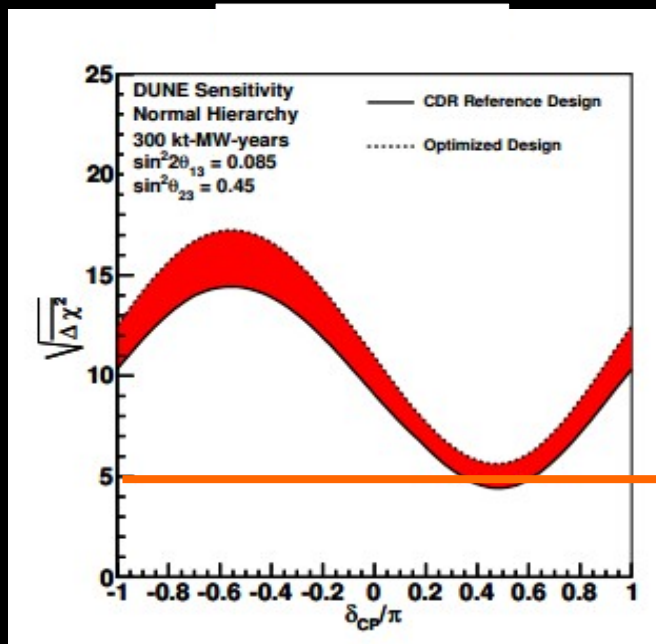


- TPC technology is not new
- TPCs usually use gas to generate ionisation electrons
- Neutrino detectors need to use liquid to get more events
- LAr is a new technology being trialled in various prototypes
- DUNE plans 4 TPC modules containing 17 kton of LAr
- Largest LAr TPCs ever built

Liquid Argon Time Projection Chamber



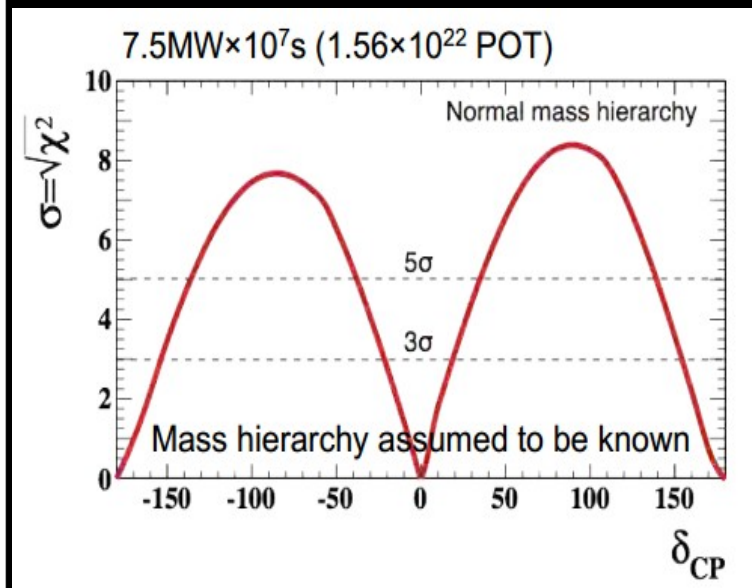
DUNE Predictions : Mass Hierarchy



After 4 years of operation

5 σ determination of mass hierarchy over entire range of δ_{CP}

DUNE Predictions : δ_{CP}



5 σ significance for
 $\sin \delta_{CP} \neq 0$ over
56% of δ_{CP} space

20% precision at
 $\delta_{CP} = -90$ degrees

Timeline

2018 : Technology prototyping programme at CERN

2019 : Delivery of Technical Design Report to funding agencies

2020 : Detector fabrication hubs ready and logistics program defined

2024 : Construction and installation of Far Detectors

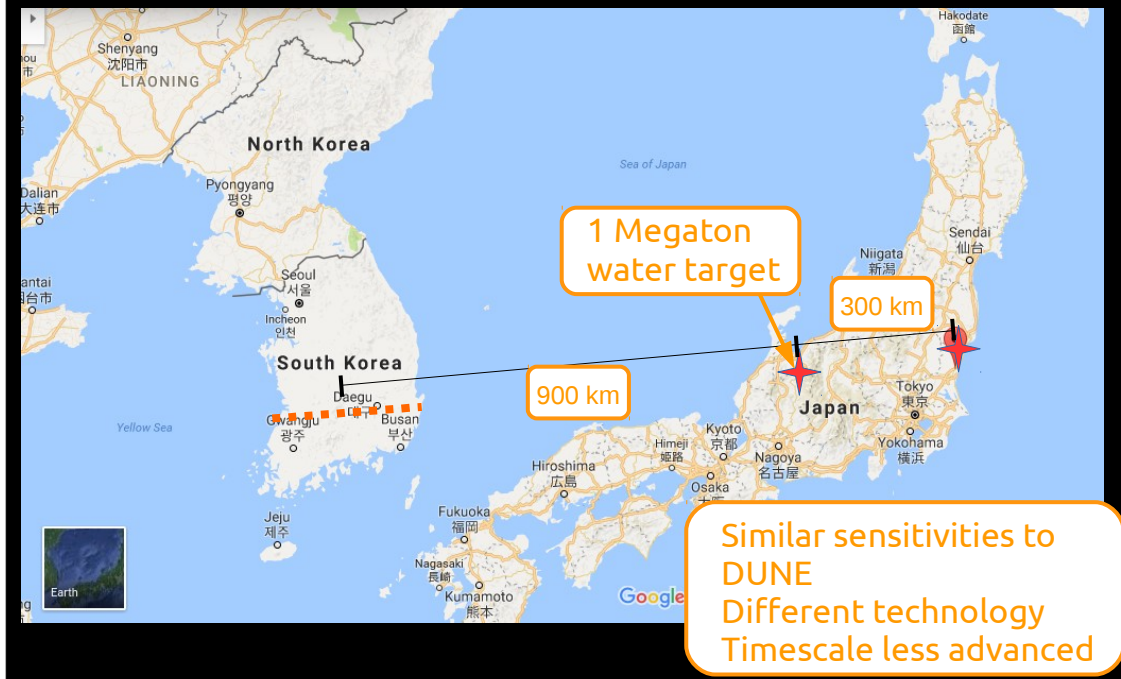
2025 : Commissioning of Far Detectors

2026 : Delivery of neutrino beam to Far site. Physics starts.

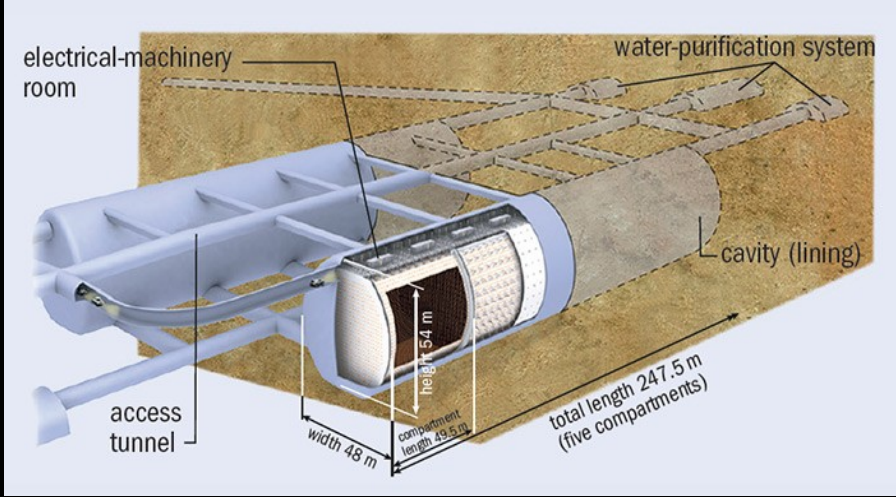
2028 : Preliminary results on mass hierarchy and δ_{CP}

2030 : Final results

Hyper-Kamiokande (The opposition)



Hyper-K Far Detector



Summary

Particle physics has studied CP violation in the quarks for years

We are only beginning to look at the leptons....and early indications are that the leptons behave significantly differently

We get all this information from long baseline experiments which are only beginning to reach the era of high-precision neutrino physics

Still 20 years of work to understand what is going on...

Assume 1 billion people eat an egg sandwich every 3 months

1.67×10^7 egg sandwiches/day

Let's say that 3 months of the year people can eat outside, and that they picnic once a month

140,000 external egg sandwiches/day

egg sandwich lifetime – 20 minutes

3000 egg sandwiches at any time

Area of egg sandwich – 15 cm x 15cm

62 m² total egg-sandwich area

Surface area of earth

500 million km²

Suppose flight paths cover area of earth uniformly

Probability of egg-sandwich/
ball bearing
intersection

1×10^{-13}

