The role of transition metals in the structure and reactivity of astrochemicals

#### Nathan J. DeYonker

#### The University of Warwick Centre for Scientific Computing 6/5/2017







#### Medium fries

#### Environmental





Prof. William Alexander @ U. Memphis

## & organometallic chemistry





Dr. Fatima Rivas @ St. Jude Children's Hospital



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 $R_2$ 

### Large fries

#### Reproducible/systematically constructed QM enzyme models





#### Are we alone in the Universe?







"I think we're going to have strong indications of life beyond Earth within a decade, and I think we're going to have definitive evidence within 20 to 30 years. We know where to look. We know how to look. In most cases, we have the technology, and we're on a path to implementing it."

-Dr. Ellen Stofan

**Previous NASA Chief Scientist** 





#### Before astrobiology, astrochemistry





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#### Molecules are out there



From "Astrochemistry and Astrobiology", Smith, Cocknell, Leach, Ed., Springer, 2013



### Sagittarius B2

- ~50% of known interstellar molecules directly observed
- Dense cloud: 300 million x mass of Sun
- Cold! -390 to 80 °F
- Weird, reactive molecules
- Origin of life?





#### 164 known interstellar molecules

| 2                     |                 | 3                  |                    | 4                  |       | 5                              | 6                    | 7                                  | 8                                | 9                                 |
|-----------------------|-----------------|--------------------|--------------------|--------------------|-------|--------------------------------|----------------------|------------------------------------|----------------------------------|-----------------------------------|
| H <sub>2</sub>        | PN              | C <sub>3</sub>     | OCS                | c-C <sub>3</sub> H | HCNO  | <b>C</b> <sub>5</sub>          | C₅H                  | C <sub>6</sub> H                   | $CH_3C_3N$                       | $CH_3C_4H$                        |
| AIF                   | SO              | CCH                | SO <sub>2</sub>    | CCCH               | HOCN  | $C_4H$                         | $H_2C_4$             | CH <sub>2</sub> CHCN               | HCOOCH <sub>3</sub>              | $CH_3CH_2CN$                      |
| AICI                  | SO <sup>+</sup> | CCO                | c-SiC <sub>2</sub> | CCCN               | HSCN  | SiC <sub>4</sub>               | $C_2H_4$             | CH₃CCH                             | CH <sub>3</sub> COOH             | $(CH_3)_2O$                       |
| <b>C</b> <sub>2</sub> | SiN             | CCS                | CO <sub>2</sub>    | CCCO               | HOOH  | $H_2C_3$                       | CH <sub>3</sub> CN   | HC <sub>5</sub> N                  | C <sub>7</sub> H                 | $CH_3CH_2OH$                      |
| CH                    | SiO             | $CH_2$             | $NH_2$             | CCCS               | HMgNC | $c-C_3H_2$                     | CH <sub>3</sub> NC   | CH₃CHO                             | $H_2C_6$                         | HC <sub>7</sub> N                 |
| CH⁺                   | SiS             | HCN                | $H_3^+$            | HCCH               | HCCO  | H <sub>2</sub> CCN             | CH <sub>3</sub> OH   | CH <sub>3</sub> NH <sub>2</sub>    | CH <sub>2</sub> OHCHO            | C <sub>8</sub> H                  |
| CN                    | CS              | HNC                | $H_2D^+$           | $NH_3$             |       | $CH_4$                         | CH₃SH                | c-CH <sub>2</sub> OCH <sub>2</sub> | CH <sub>2</sub> CCHCN            | $CH_3CONH_2$                      |
| CO                    | HF              | HCO                | SiCN               | HCCN               |       | HC <sub>3</sub> N              | HC <sub>3</sub> NH⁺  | CH <sub>2</sub> CHOH               |                                  | C <sub>8</sub> H⁻                 |
| CO <sup>+</sup>       | SH              | HCO⁺               | AINC               | HCNH <sup>+</sup>  |       | HCCNC                          | HC <sub>2</sub> CHO  | C <sub>6</sub> H⁻                  |                                  | CH <sub>3</sub> CHCH <sub>2</sub> |
| CP                    | HD              | HCS⁺               | KCN                | HNCO               |       | HCOOH                          | NH <sub>2</sub> CHO  | CH₃NCO                             |                                  |                                   |
| SiC                   | CF <sup>+</sup> | HOC <sup>+</sup>   | HCP                | HNCS               |       | CH <sub>2</sub> NH             | C <sub>5</sub> N     |                                    |                                  |                                   |
| HCI                   | PO              | $H_2O$             | CCP                | $HCO_2^+$          |       | H <sub>2</sub> CCO             | $HC_4N$              | 10                                 | 11                               | >12                               |
| KCI                   | AIO             | $H_2S$             | AIOH               | H <sub>2</sub> CO  |       | NH <sub>2</sub> CN             | C <sub>5</sub> N⁻    | $CH_3C_5N$                         | HC <sub>9</sub> N                | HC <sub>11</sub> N                |
| NH                    | HCI⁺            | HNO                | FeCN               | H <sub>2</sub> CN  |       | HNCCC                          |                      | CH <sub>3</sub> COCH <sub>3</sub>  | CH <sub>3</sub> C <sub>6</sub> H | <b>C</b> <sub>60</sub>            |
| NO                    | SH⁺             | MgNC               | $HO_2$             | $H_2CS$            |       | SiH <sub>4</sub>               |                      | (CH <sub>2</sub> OH) <sub>2</sub>  |                                  | C <sub>70</sub>                   |
| NS                    | CN⁻             | MgCN               | Si <sub>2</sub> C  | H <sub>3</sub> O⁺  |       | $H_2COH^+$                     |                      |                                    |                                  | C <sub>60</sub> +                 |
| NaCl                  | OH⁺             | NaCN               | CCN                | c-SiC <sub>3</sub> |       | C₄H⁻                           |                      |                                    |                                  |                                   |
| ОН                    | TiO             | $N_2H^+$           | $H_2O^+$           | CH <sub>3</sub>    |       | CH <sub>3</sub> O              | # of kn              | own mo                             | lecules or                       | Earth:                            |
| ArH⁺                  | NO <sup>+</sup> | $N_2O$             |                    | $PH_3$             |       | $NCCNH^{\scriptscriptstyle +}$ |                      |                                    |                                  |                                   |
| 0 <sub>2</sub>        |                 | H <sub>2</sub> Cl⁺ |                    | C <sub>3</sub> H⁺  |       | $NH_4^+$                       | $\sim 120 \text{ n}$ | nillion!                           |                                  |                                   |

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### Complexity of bioinorganic molecules on Earth









Mo-dependent nitrogenase (Fe and Mo)

#### Monomer of Dichloromuconate cycloisomerase (Mn)



# Complexity of known organometallic molecules in ISM

- Fe-C-N
- Mg-C-N
- K-C-N
- Na-C-N





When does "astrochemistry" become "astrobiology"?

- We are not close to knowing  $\boldsymbol{\boldsymbol{\otimes}}$
- Huge disconnect between molecular complexity
- **Slowly** learning more about composition of space and how chemistry is transported
- Fascinating new (experimental and theoretical) discoveries await us <sup>(2)</sup>
- What is the role of inorganic chemistry in space?



Difficulties in molecular astrochemistry

- Astrophysical observation
  - Low abundances of metals in ISM
  - Heavy atoms condense into ices and grains
- Human factors
  - Few people are searching for small inorganics



- Overlapping transitions from different molecules
- Volatile / toxic molecules





Internuclear distance









#### Observation of the FeNC molecule by laser fluorescence excitation spectroscopy

Jie Lie and Paul J. Dagdigian<sup>a)</sup> Department of Chemistry, The Johns Hopkins University, Baltimore, Maryland 21218-2685

(Received 15 September 2000; accepted 7 November 2000)

 Dagdigian in 2001 – published experimental fluorescence study of FeNC



Lucy Ziurys University of Arizona



H. Fritz Schaefer III University of Georgia



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FeNC Cite t

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www.rsc.org/advances

#### PAPER

Predicting the interactions of organometallic ruthenium ethylenediamine complexes with mononucleotides: insights from density functional theory<sup>†</sup>

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#### Observation of the FeNC molecule by laser fluorescence excitation spectroscopy

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- Dagdigian in 2001 published experimental fluorescence study of FeNC
- First project in graduate school
  - Use computational chemistry to study properties of FeCN and FeNC
- "3 months"? No way, 3 years



#### Fe-containing molecules are a hot mess

- Fe  $1s^22s^22p^63s^23p^64s^23d^64p^0$
- Electronic & vibrational energy levels are too close
- Leads to spectral perturbations





### Multireference vs single reference

- Atoms and molecules are actually linear combinations of electronic configurations
- For most atoms and organic molecules,
  - From Hund's rules, Aufbau principle, Pauli exclusion principle, LCAO-MO, dominant config > 98 % of the wave function
  - Organic exceptions:
    - Carbon dimer (C<sub>2</sub>)
    - Diradicals: molecules where two electrons occupy nearly degenerate molecular orbitals



### $C_3H_6$





Stri**glet** Svingliverteterence



### SR methods (coupled cluster)

- Single Hartree-Fock reference configuration ( $\Psi_0$ )
- Contribution of other configurations to  $\Psi \,\,^{virtual \,\, space}$ 
  - *n*-body substitutions of electrons into unoccupied MOs
  - Amplitudes of e<sup>-</sup> substitutions and their products: coupled cluster theory
  - CCSD, CCSD(T), CCSDT, CCSDTQ, etc.
- Level of e<sup>-</sup> e<sup>-</sup> interactions can be systematically improved

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- Increased substitution level = vastly increased computation time
- Short-range or dynamical correlation



core orbitals



#### MR methods

- All n-body e<sup>-</sup> e<sup>-</sup> interactions are accounted for in a limited "active space" (valence or frontier MOs)
- 1 and 2 e<sup>-</sup> substitutions from valence electrons to virtual space (MRCISD)
- MR methods appropriate when near degeneracies exist in electronic  $\Psi$ 
  - Long-range, strong, nondynamical, or static correlation



#### Degeneracies vs. near degeneracies



• Two different kinds of nondynamical electron correlation!



### Frontier MO diagram of FeCN/FeNC

(Fe d<sub>xy</sub> and d<sub>x<sup>2</sup>-y<sup>2</sup></sub> not shown - always w/ 3 e<sup>-</sup> in L = 2 MOs to form <sup>4</sup> $\Delta$  or <sup>6</sup> $\Delta$  state)



**Biogh splits:** 13658%  $4.\pm 11 + 8 + 7\% = 77\%$ 





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- A saying in computational chemistry: "The right answer for the wrong reason"
  - Parameterization
  - Picking one of hundreds of semi-quantitatively accurate methods (DFT) until finding agreement with experiment
- "Wrong answer for right reason?!"
  - MRCI versus CC
  - Unbalanced treatment of static vs dynamic correlation

JOURNAL OF CHEMICAL PHYSICS

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8 MARCH 2004

#### Low-lying electronic states of FeNC and FeCN: A theoretical journey into isomerization and quartet/sextet competition

Nathan J. DeYonker, Yukio Yamaguchi, Wesley D. Allen,<sup>a)</sup> Chaeho Pak, and Henry F. Schaefer III *Center for Computational Chemistry, University of Georgia, Athens, Georgia 30602* 

Kirk A. Peterson Department of Chemistry, Washington State University, Pullman, Washington 99164

(Received 18 August 2003; accepted 4 November 2003)

N. J. DeYonker, H. F. Schaefer, K. A. Peterson, and coauthors J. Chem. Phys. 120, 4726 (2004).



### MRCI vs CC

- Many transition metal-containing radical molecules exhibit both static and dynamical correlation
  - "Pathological cases":
    - MRCISD treats enough static but not enough dynamical
    - Conventional CC treats enough dynamical but not enough static
- Over last 20 years
  - MRCI developments = slow
  - MRCC developments = VERY slow
  - CC developments = steady





### Work since 2004

- Some experimental labs tried (and failed) to reproduce the Dagdigian work and synthesize FeCN and FeNC
  - L. Ziurys group (Arizona)
  - M. Heaven group (Emory)
- Some theoretical groups were skeptical of (and ignored) our coupled cluster results

– MRCISD must be correct!

- T. Hirano (Ochanomizu) and P. Jensen (Bergische)



#### CENTER FOR ADVANCED SCIENTIFIC COMPUTING AND MODELING UNIVERSITY OF NORTH TEXAS





#### CASCASCAM CENTER FOR ADVANCED SCIENTIFIC COMPUTING AND MODELING UNIVERSITY OF NORTH TEXAS Assessing "pathology" of inorganic molecules







W. Jiang, N. J. DeYonker, J. J. Determan, A. K. Wilson, J. Phys. Chem A. 116, 870 (2012).

W. Jiang, N. J. DeYonker, A. K. Wilson, J. Chem. Theory Comput. 8, 460 (2012).

www.aci.org

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#### Fourier-transform microwave spectroscopy of FeCN ( $X^4\Delta_i$ ): Confirmation of the quartet electronic ground state

L.N. Zack, J. Min, B.J. Harris<sup>1</sup>, M.A. Flory<sup>2</sup>, L.M. Ziurys\*

Department of Chemistry, Department of Astronomy and Steward Observatory, 933 North Cherry Avenue, University of Arizona, Tucson, AZ 85721, USA

#### Millimeter-wave rotational spectroscopy of FeCN (X ${}^{4}\Delta_{i}$ ) and FeNC (X ${}^{6}\Delta_{i}$ ): Determining the lowest energy isomer

M. A. Flory and L. M. Ziurys<sup>a)</sup> Department of Chemistry and Department of Astronomy, Steward Observatory, University of Arizona, 933 N. Cherry Ave., Tucson, Arizona 85721, USA

#### DETECTION OF FeCN ( $X^4\Delta_i$ ) IN IRC+10216: A NEW INTERSTELLAR MOLECULE

L. N. ZACK<sup>1</sup>, D. T. HALFEN<sup>1,2</sup>, AND L. M. ZIURYS<sup>1,2</sup>

<sup>1</sup> Department of Chemistry, University of Arizona, P.O. Box 210041, 1306 East University Boulevard, Tucson, AZ 85721, USA
<sup>2</sup> Department of Astronomy and Steward Observatory, Arizona Radio Observatory, University of Arizona, 933 North Cherry Avenue, Tucson, AZ 85721, USA Received 2011 March 24; accepted 2011 April 20; published 2011 May 10

- FeCN (low spin) and FeNC (high spin) have a different ground state electronic configuration
- Our coupled cluster results were correct!

Zack, L. N.; Min, J.; Harris, B. J.; Flory, M. A.; Ziurys, L.M.; *Chem. Phys. Lett.* **2011,** *514* (4-6), 202-206. Flory, M. A.; Ziurys, L. M.; *J. Chem. Phys.* **2011,** *135* (18), 184303. Zack, L. N.; Halfen, D. T.; Ziurys, L. M., *Astrophys. J. Lett.* **2011,** *733* (2), L36.

### 10 years later, revisit FeCN/FeNC

- Calibration of inorganic chemistry and spectroscopy
- Confident that CC can often outperform MRCI
- New technological developments
  - Improved basis sets treatment of 1-electron space
  - Better treatment of core/valence (CV) interactions
  - Scalar relativistic effects
  - Arbitrary order coupled cluster (CCSDTQ and beyond)
  - Parallel computing
  - Spin-orbit coupling

FeH: N. J. DeYonker, W. D. Allen, *J. Chem. Phys.*, **137**, 234303 (2012). VCI<sup>+</sup>: N. J. DeYonker, D. T. Halfen, W. D. Allen, L. M. Ziurys, *J. Chem. Phys.*, **141**, 204302 (2014).







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DeYonker, N. J.; J. Phys. Chem. A 2015, 119 (1), 215.

### Current work: TM monoacetylides

- M-C≡CH isoelectronic with MCN/MNC
  - Many polyynes (CC<sub>n</sub>R, R = H,C,N,P) known in ISM
  - Organometallic catalysis
- Expt. data for M = Cr, Ni, Cu, Zn
- More discrepancies between CCSD(T) and MRCISD+Q
- 10 MCCH 18 months vs.
   FeCN&FeNC 3 years

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| Metal | MCN<br>ground<br>state | Computed<br>MCCH<br>ground<br>state |  |
|-------|------------------------|-------------------------------------|--|
| Sc 🤇  | зД                     | 1Σ+                                 |  |
| Ti    | <sup>4</sup> Ф         | $^{4}\Phi$                          |  |
| V     | <sup>5</sup> Δ         | <sup>5</sup> Δ                      |  |
| Cr    | 6 <b>Σ</b> +           | <sup>6</sup> Σ+                     |  |
| Mn    | <sup>7</sup> Σ+        | <sup>7</sup> Σ+                     |  |
| Fe    | 4Δ                     | <sup>6</sup> ∆?                     |  |
| Со    | 5Φ                     | <sup>3</sup> Φ, <sup>5</sup> Φ      |  |
| Ni    | <sup>2</sup> Δ         | <sup>2</sup> Δ                      |  |
| Cu    | <sup>1</sup> Σ+        | <sup>1</sup> Σ <sup>+</sup>         |  |
| Zn    | <sup>2</sup> Σ+        | <sup>2</sup> Σ+                     |  |



www.q-chem.org

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#### DOI: 10.1002/qua.25206



w/ Shelby Dickerson (now at U. South Carolina)

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#### FeCCH déjà vu



S. D. Dickerson, N. J. DeYonker, Int. J. Quantum. Chem., 117, 104 (2017).



#### TM rovibrational spectroscopy

- CuOH, CuCN, CuCCH w/ Ryan Fortenberry (Georgia Southern)
- "Black box" vibrational quartic force fields
- Different prescription for inorganic species
  - Relativistic effects
  - Improved CV treatment



|            | Comp. (cm <sup>-1</sup> ) | Expt.  |
|------------|---------------------------|--------|
| $v_{OH}$   | 3662.1                    |        |
| $V_{CuO}$  | 625.3                     | 625(1) |
| $v_{bend}$ | 735.1                     | 743(1) |

#### **POSS reactions**

- Exploration of reactions on ice/grain surfaces with ionic metal defects (Ni<sup>+</sup>, Co<sup>+</sup>, Fe<sup>+</sup>)
  - PolyOligoSilSesquioxanes (siliceous clusters)
  - Interstellar rxns on grain/ice/mineral surfaces
  - Heterogeneous catalysis
  - Double-hybrid DFT & MP2-F12b calculations



M. Fioroni, N. J. DeYonker, ChemPhysChem, 17, 3390 (2016).



### From astrochemistry to astrobiology!

- Building molecular complexity
- Regioselectivity
  - Mechanism of butyraldehyde from propene



- Enantioselectivity
  - Non-racemic mixtures of origin-of-life chemicals
  - Mechanism of methyl butanal from butene





### Outlook

- Studying gas phase TM-containing radicals is quite challenging
- Reactivity of inorganics to produce more complex biorelevant molecules will chronically stretch *ab initio* theory to its limits
- We can assess single reference & multireference methods

- Calibration of high-accuracy rovibrational and vibronic spectroscopy techniques is underway
- Multiscale modeling of reactions on grain & ice surfaces

# Computational astrochemistry is awesome!

- Building trust
  - Terrestrial experimentalists need to believe in computations
    - Finally working together to understand and characterize undiscovered molecules
    - Increased throughput in astronomical observation
- How can we always efficiently get "right answer for right reason"?





DeYonker Lab:
 Thomas Summers
 Hannah Lozano
 Kate Charbonnet
 Dr. Qianyi Cheng



Lucy Ziurys

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Ryan Fortenberry



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